

# LWA Outreach South of the Border

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## The Large Millimeter Telescope



Currently at 30 meter  
(could go to 50-m)

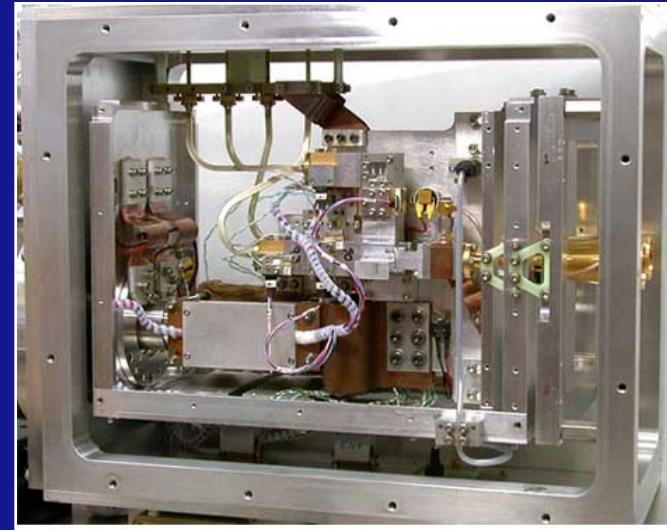
4600 meter elevation  
1 - 3 mm observing

Excellent potential as a  
mm-VLBI station

2<sup>nd</sup> Inauguration on  
May 17<sup>th</sup>  
Observing by Fall 2011?

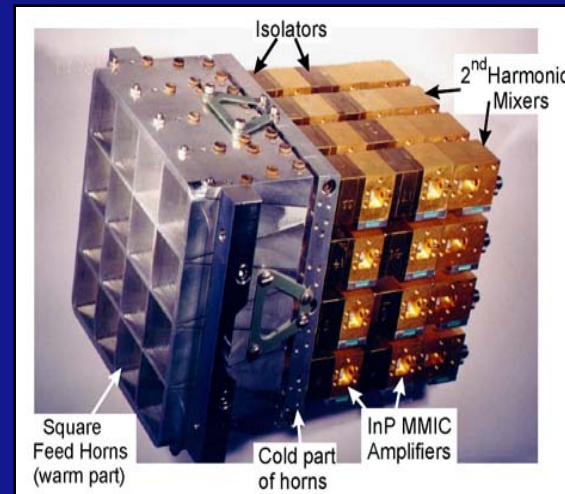
# LMT Instrumentation

- ◆ Ultra-wideband receiver covering 75 – 111 GHz in a single tuning
- ◆ 4 pixels, dual polarization
- ◆ Analog auto-correlator with 31 MHz resolution or 100 km/s over full 36 GHz
- ◆ Searches for CO in high redshift ( $z > 3$ ) galaxies with exceptional baseline stability



Redshift  
Search  
Receiver

- ◆ 32 pixels in a dual polarization 4x4 array
- ◆ 85 – 116 GHz
- ◆ Noise temperature of 50 – 80 K
- ◆ Ideal for mapping molecular clouds and star formation regions



SEQUIOA  
Focal Plane  
Array

Also available is the AzTEC 144 pixel bolometer array, operating at 2.1 and 1.1 mm  
Successfully commissioned on the JCMT in 2005

# MEXART: 140 MHz Dipole Array



- IPS solar wind studies by UNAM Geophysicists
- ◆ Design and development in the USA and India
- ◆ Analog Butler Matrix

Operation Frequency  
Basic Antenna Element  
Number of Elements  
Element Arrangement  
Bandwidth  
Angular Resolution

139.65 MHz  
Full wavelength dipole  
4096  
64 E-W lines of 64 dipoles each  
1 MHz  
 $1^\circ \times 1^\circ$

E-W row  
64 dipoles

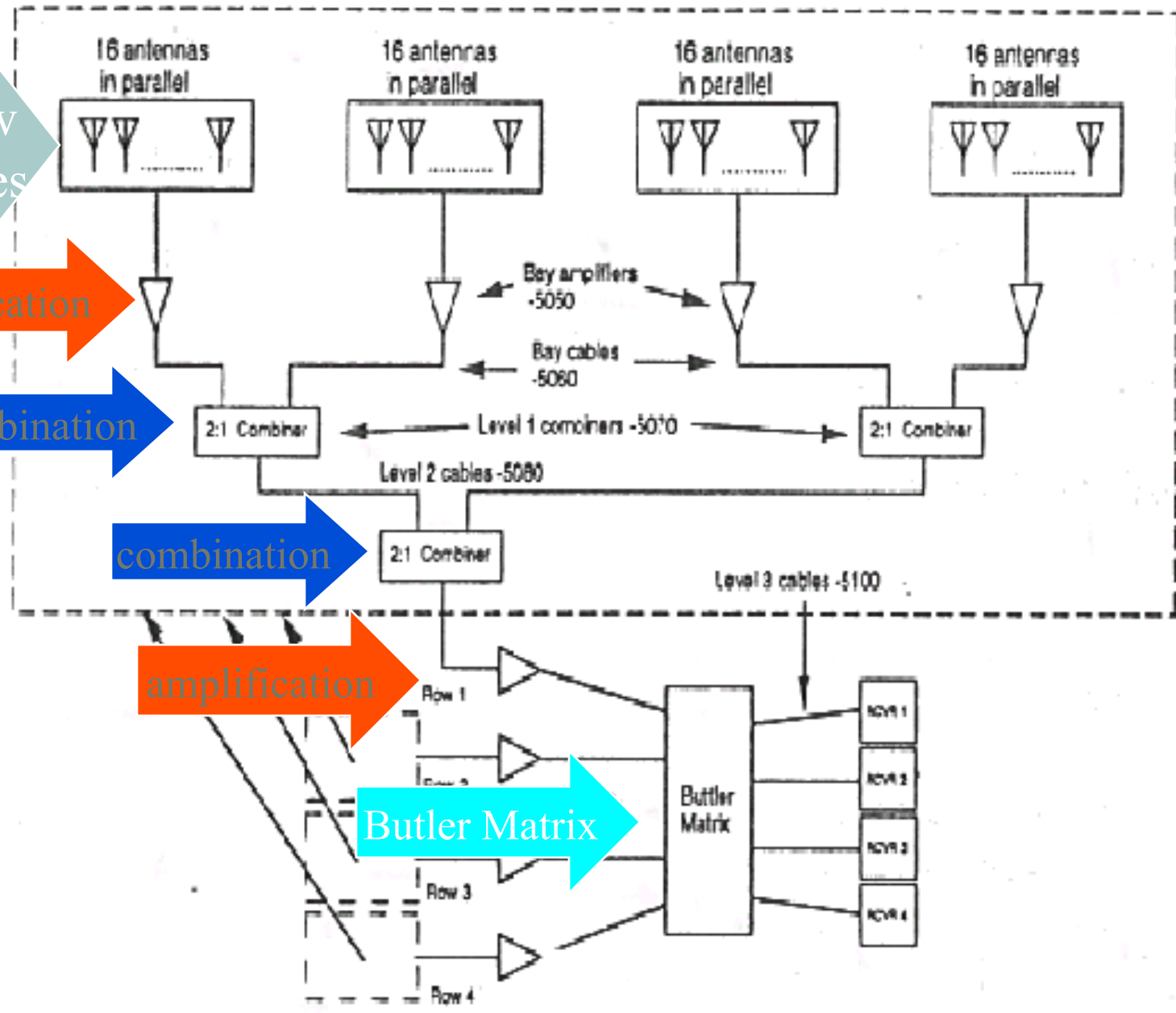
amplification

combination

combination

amplification

Butler Matrix



# EVLA

+

# VLBA

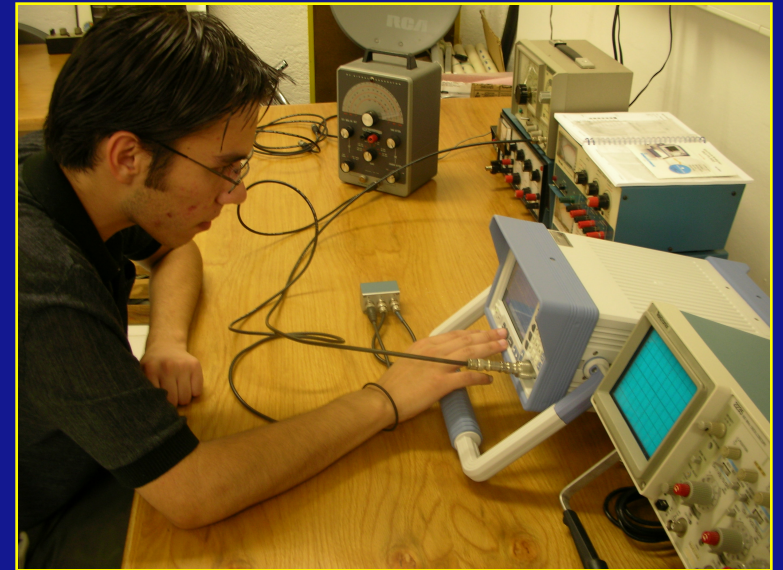


Mexico contributes 1.76 M\$  
NRAO outfits 2 antennas  
Ensures Mexican access to ALMA



Mexico contributes 0.4 M\$  
to buy disk drives and  
playback units

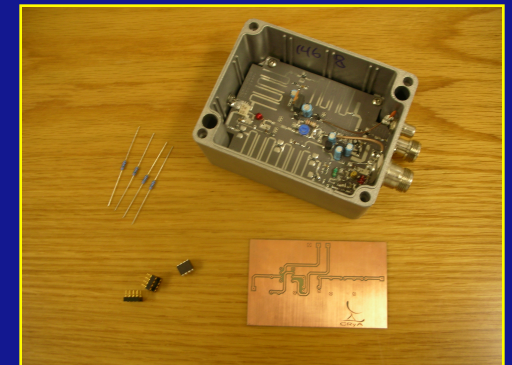
# CRyA Radio Observatory



Bare-bones RF laboratory

Two Haystack SRT systems  
C, Ku-band projects underway  
Interferometer projects underway  
at 20 MHz and 1400 MHz

Mostly undergraduate level,  
But with potential for  
masters/PhD projects



Lots of outreach to schools

# Radio Jove: A project for high schools & colleges



**Project Radio JOVE**

Learning Science by Listening to the All-Natural Jupiter and Solar Radio Stations



Project initiated by NASA

Observations of solar and jovian storms

High school level

Introduction to:  
Electronics  
Antennas & Radio  
Data analysis  
RFI  
& many more...





# What's a Radio Jove System?



15m  $\lambda/2$  dipole(s)



20.1 MHz DC Rx



PC



Optional filter  
& Noise diode



Final Data via Sky Pipe

# Mexico City High Schools using Radio Jove

- ◆ First system installed in 2005
- ◆ 10 high schools currently participating
- ◆ Plans to extend to technology institutes
- ◆ Multiple projects, not just Radio Jove



# How does the LWA fit in?

1. Mexico has radio instrumentation needs that it currently can't meet (mostly for lack of knowledge and experience)
2. Corollary: we need to train ourselves to do radio instrumentation
3. CRyA radio lab needs challenging projects for students
4. High schools need to keep pushing the envelope
5. Worthwhile goal is to involve technology institutes
6. Mexican participation in LWA (with \$) is feasible

# How LWA fits in

# Step 1: Broadband Solar Telescopes

12 antennas



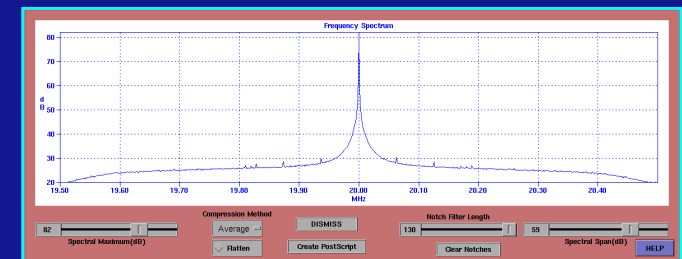
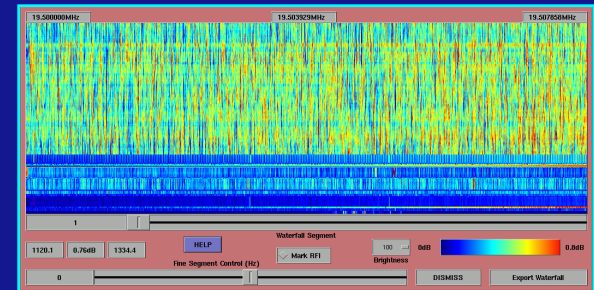
USRP 1



4 channel input  
(2 I/Q pairs)  
64 Ms/s sampling  
12 bit resolution  
DC to 2400 MHz  
Altera FPGA

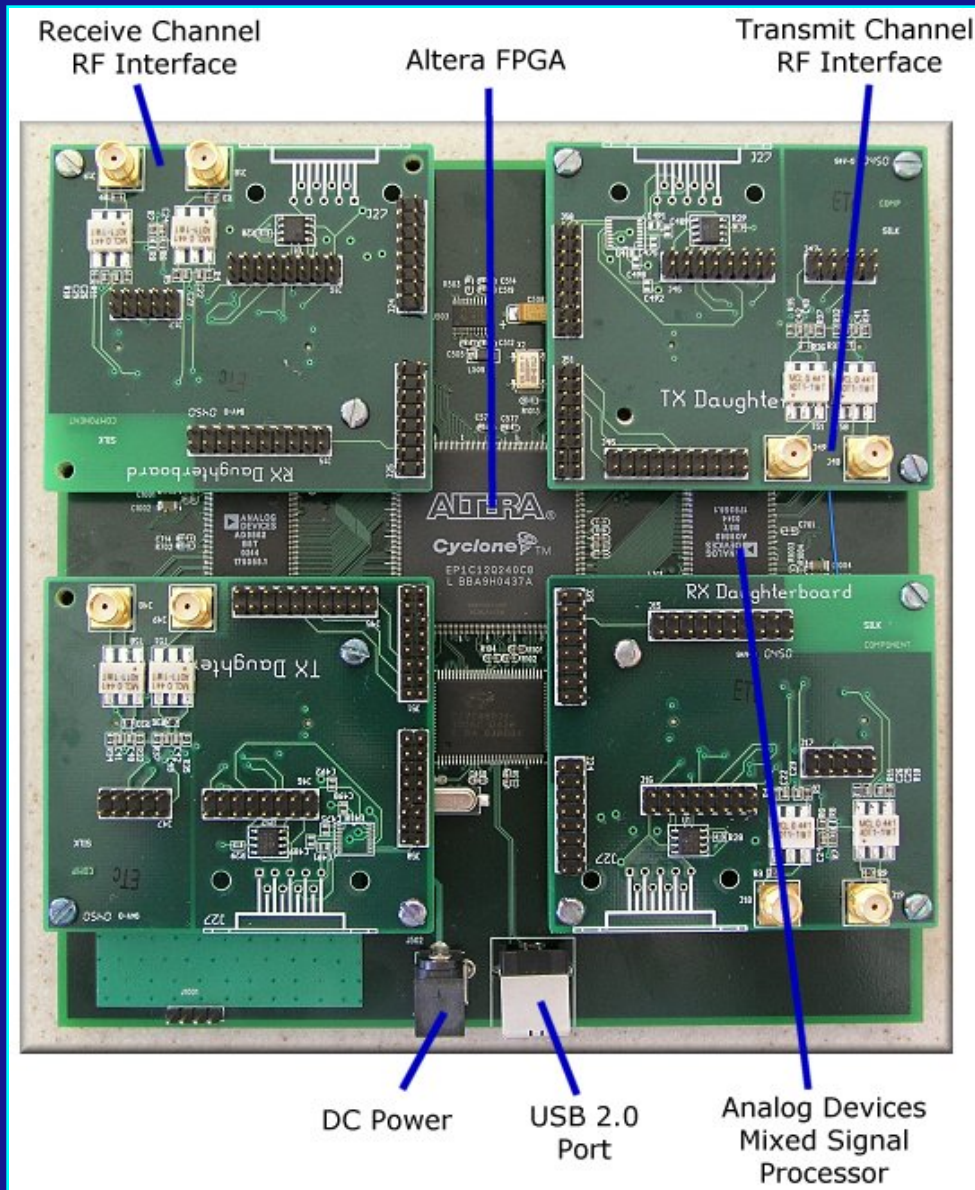
The screenshot displays the GnuRadio GUI with several control panels:

- Detector Value:** 3865066.0, Peak Value: 4255046.5
- Transients:** 03:40:31 264 3.20 3.17
- RF Controls:** Tuner Frequency (Hz) 20000000.000000, Sky Frequency (Hz) 20000000.000000, GCA 1.000000, GC B 1.000000, RF Gain 65, DC Gain 5, Mult. x10.
- Continuum Controls:** Offset 0.0, Integration 15, Refmut 1.0, SHOW button.
- SETI Analysis:** Dicke Mode: OFF, Sigma\_K 2.75, SETI Integration 45.00, SHOW button.
- Other controls:** Pulsar Rate -1.0000000000, Folding Const 0.1, DM, Delay (ms), PhaseErr, Power Threshold 2.5, Transient Duration 3, Current Declination -15.000000, Current RA, STAT indicator.



## GnuRadio Software Defined Radio

# Ettus Research & National Instruments



## Available Rx modules

- DC - 30 MHz ✓
- 50 - 860 MHz ✓
- 800 - 2400 MHz

USRP1 capable of 16 MHz bandwidth

Our bandwidth set by PC (6 core 3.3 GHz CPU)  
We expect 4 MHz with swept operation

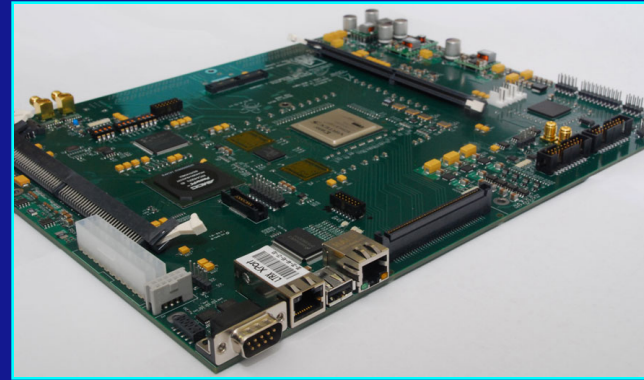
Small cluster would be ideal

## Step 2: More ambitious projects



more LWA antennas

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ROACH or RHINO board

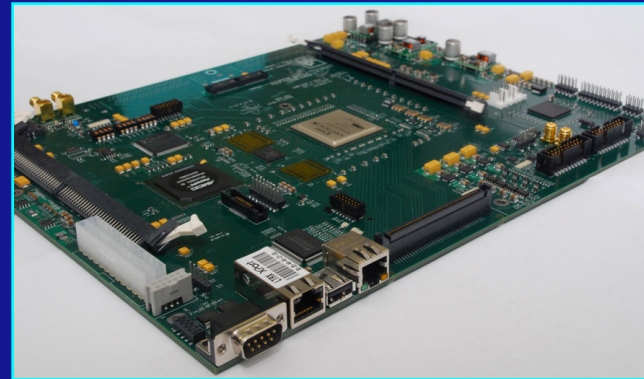
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Correlator for interferometry  
Digital Beamformer  
Gain valuable experience and  
involve Technology Institutes

## Step 3: Even *more* ambitious projects



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LOTS more LWA antennas

Multiple ROACH and/or  
RHINO boards

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LWA South station in northern Mexico?  
Fabrication of antennas & front ends in Mexico?  
Digital beamformer for MEXART?  
Digital spectrometer for LMT?

# Interference Study at 139.65 MHz +/- 1 MHz

