

The Extragalactic Radio Sky at Faint Flux Densities

Dr Carole Jackson Research School of Astronomy & Astrophysics October 2002







Probing deep fields...



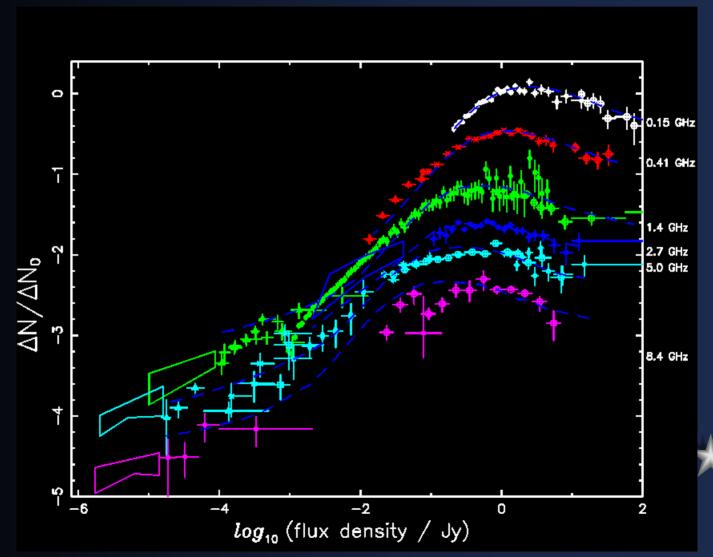
~ 3000 galaxies

13 radio sources

Radio waveband samples different population of galaxies

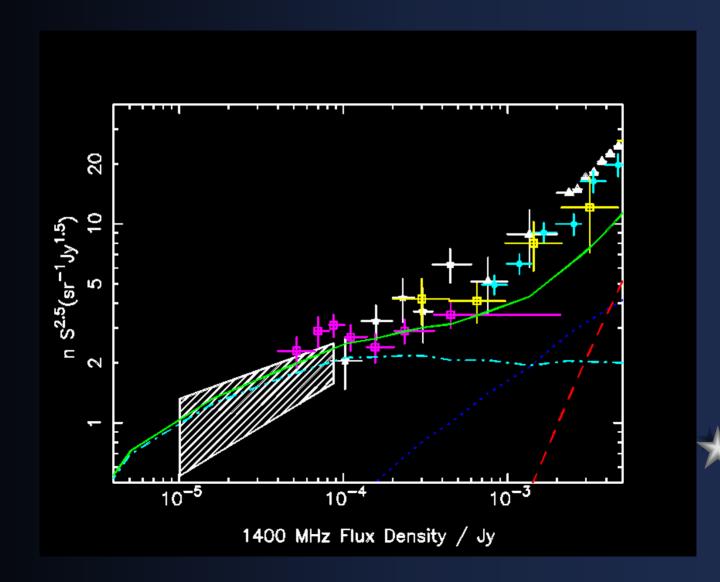


Current Deep Radio Surveys

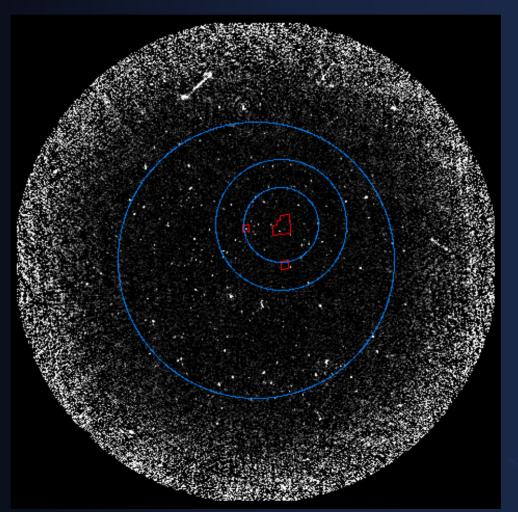




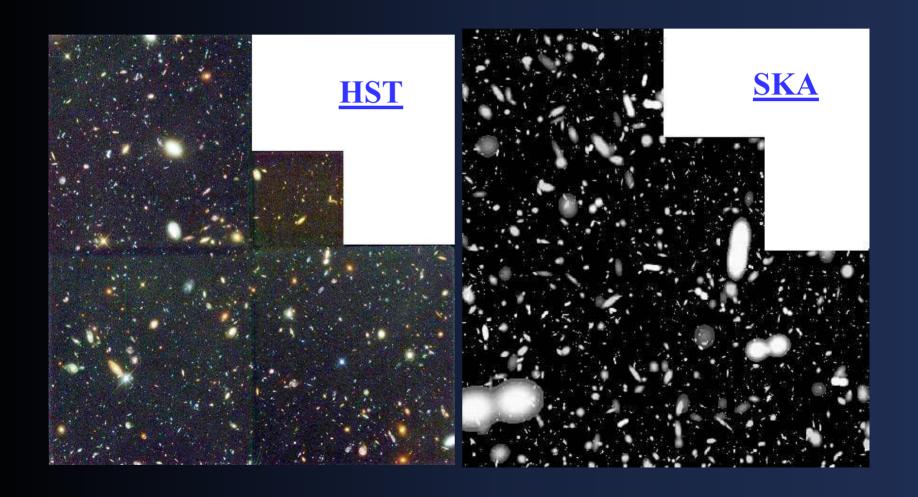
Current Deep Radio Surveys



HDF-s ATCA Image (rms=7 microJy)

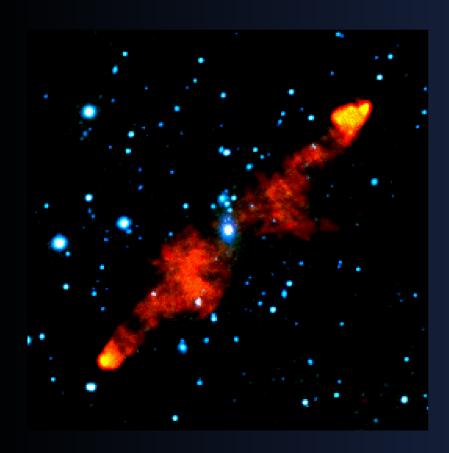


The future 1 nJy at 1.4 GHz?



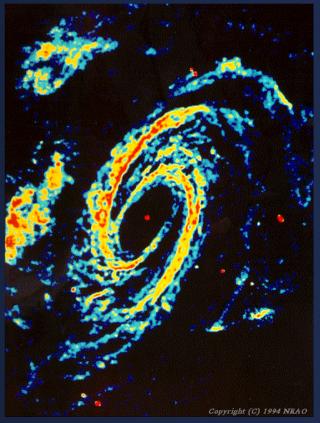
Radio-loud AGN (Quasars & radio galaxies)

CSIRO ATCA PKS 2356-61 FRII RG



Starburst galaxies

NRAO VLA M81 spiral galaxy









Physical characteristics Of the source populations

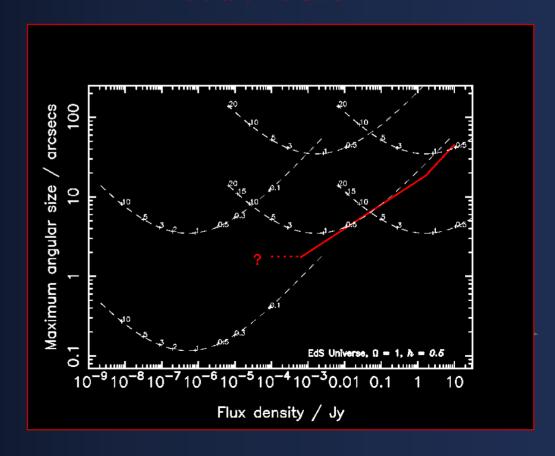
Spectral Shape

Simple $S \propto v^{\alpha}$ with -0.7 or fitted spectral model

-Ignores (peaked) lowfrequency population (if there is one)

-Ignores GHz-peaked sources

Source Sizes





Recipe for predicting the radio sky from the LRLF + Evolution

For the 3 radio galaxy populations (FRI, FRII & SB):

- -Determine the LRLF & Evolution Adopt reasonable evolution type (LDDE) Use source counts & complete samples to constrain model
- -Transpose Frequency if required
- Calculate source density (sky area, z distr)
- Adopt reasonable source sizes & shapes Randomly place & orient sources on sky

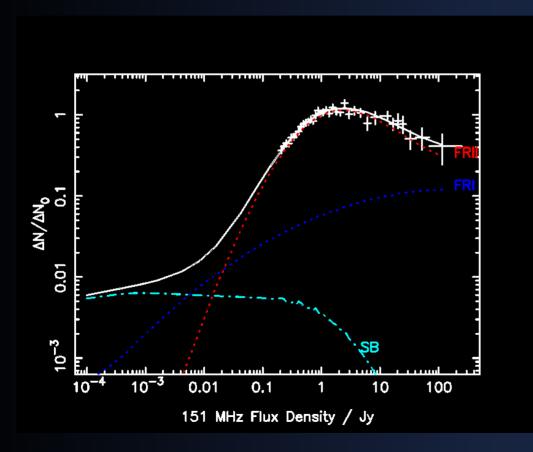


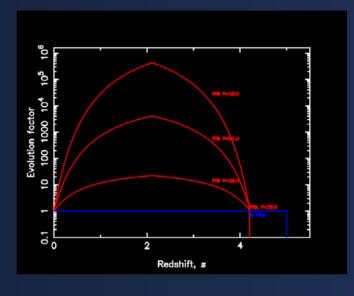




FRI & FRII Evolution & the LRLF

Best-fit to 151 MHz source count





LDDE - FRIIs strongly evolving, FRIs not









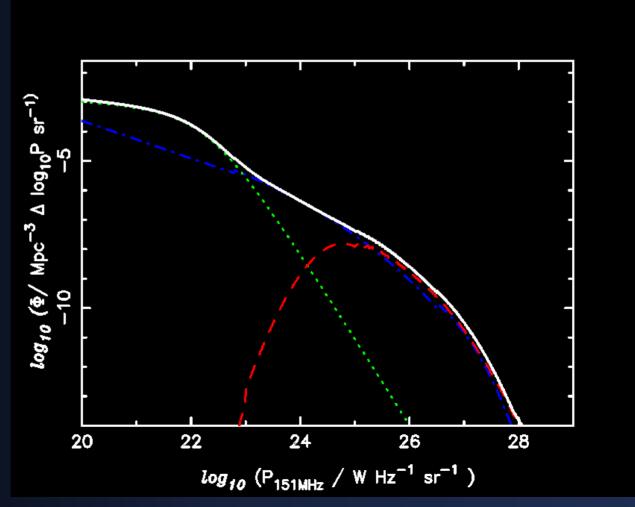
Evolution & the LRLF

LRLF from best-Fit model.

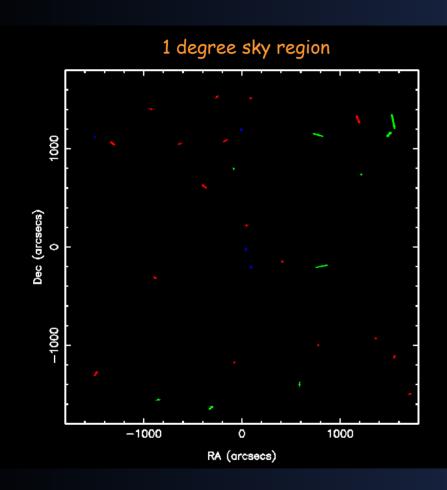
Starburst galaxy

LRLF from 2dFGRS-NVSS (Sadler et al 2002)

Evolution from HDF (Haarsma et al 2000)

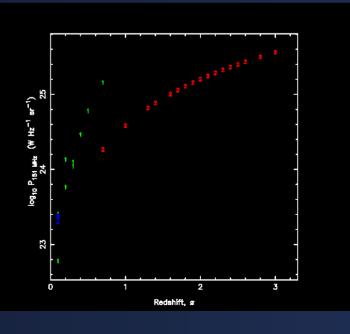




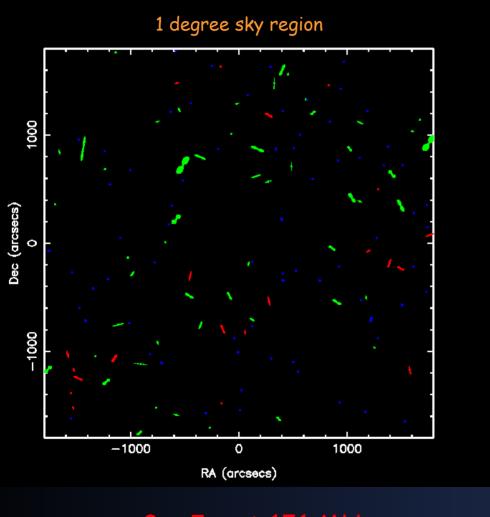


10 mJy at 151 MHz

P-z distribution

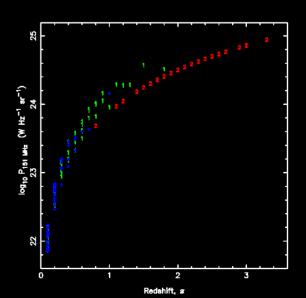






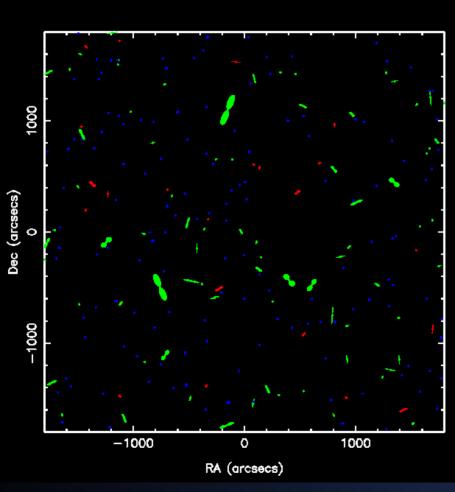
2 mJy at 151 MHz

P-z distribution



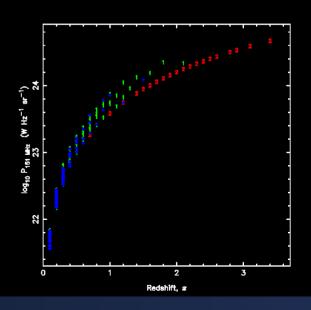


1 degree sky region



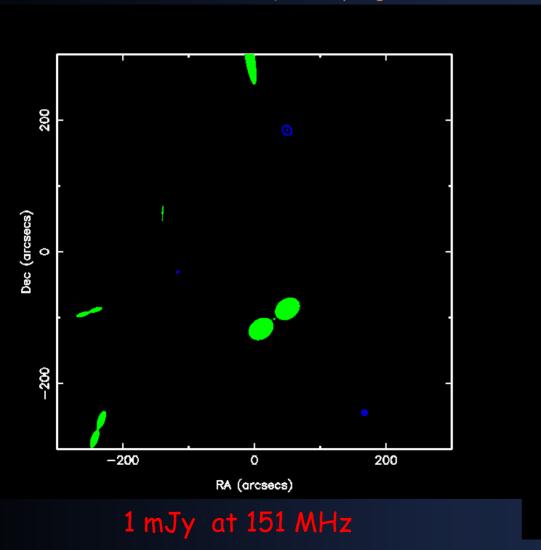
1 mJy at 151 MHz

P-z distribution

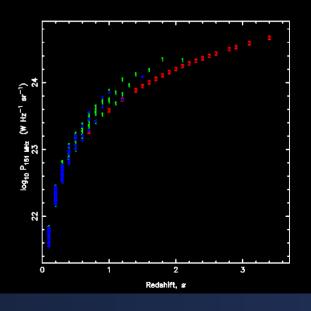




10 arcmin square sky region

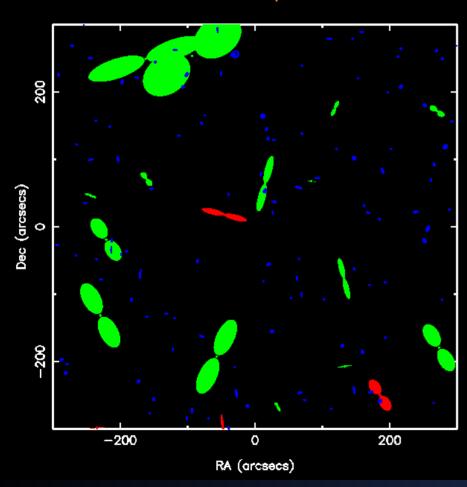


P-7 distribution



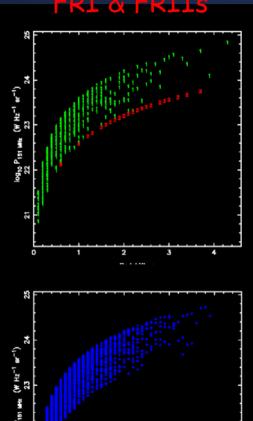


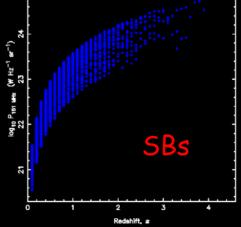




0.1 mJy at 151 MHz

FRI & FRIIs







From simulated skies - predict resolution required - fraction of sources 'overlapped' (line of sight)

Assumes no frequency-size dependence (probably ok up to 1 GHz?)

Assumes no size-RG age dependence

FRI+FRII LRLF + evolution (?) - degenerate, really FR-split ???

Starburst LRLF + evolution (?) - difficult to determine (HDF small sample)
- Late-type galaxies in LRLF?

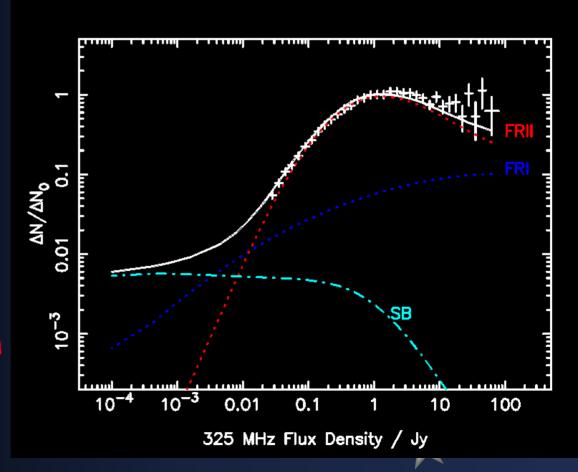
Other populations?



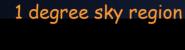
325 MHz - WENSS source count

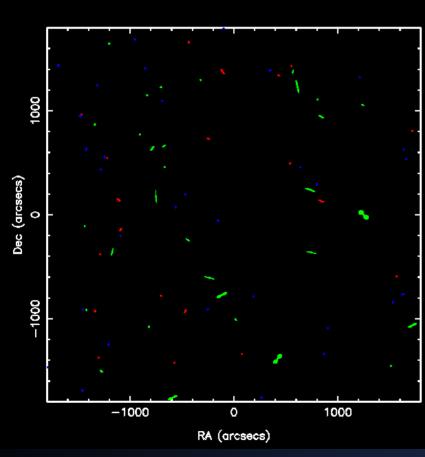
Transpose LRLFs to 325 MHz & generate model count to 0.1mJy

Molonglo Demonstrator project - to 300 MHz science - HI absorption against bright RGs

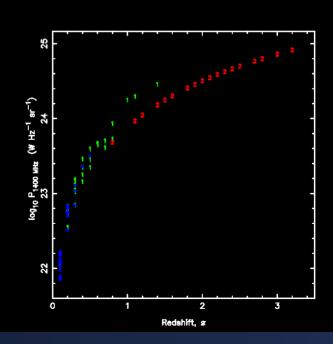






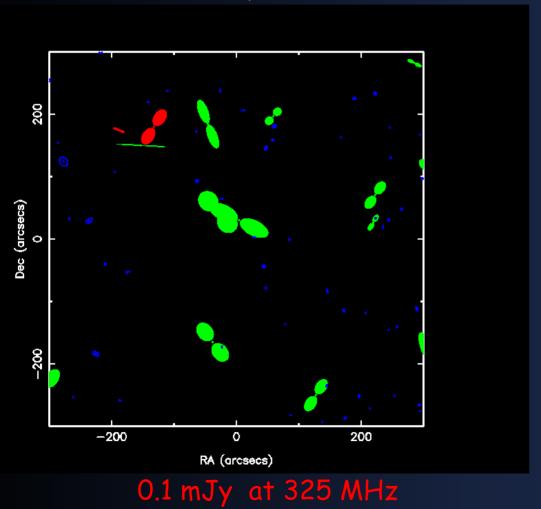


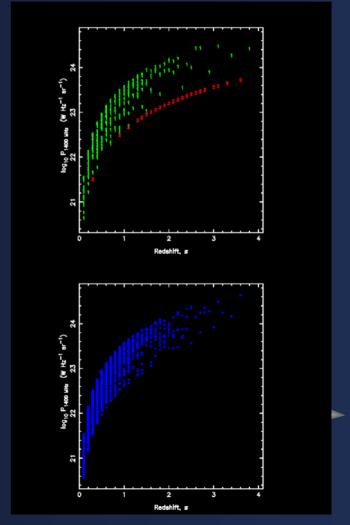
2 mJy at 325 MHz





10 arcmin square







cjackson@mso.anu.edu.au

