

Active Galaxies



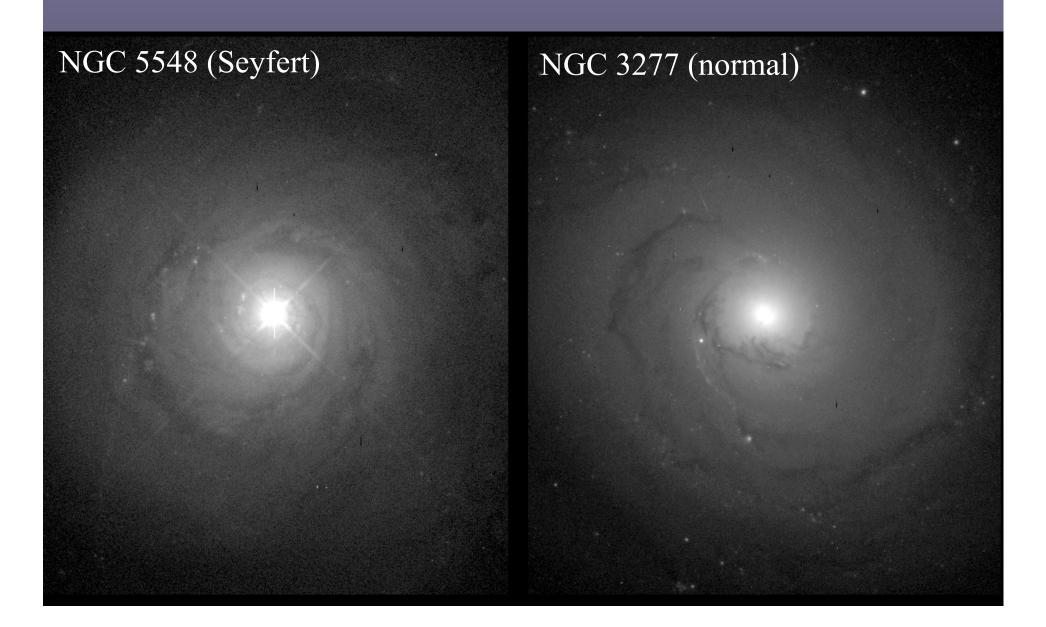


Outline

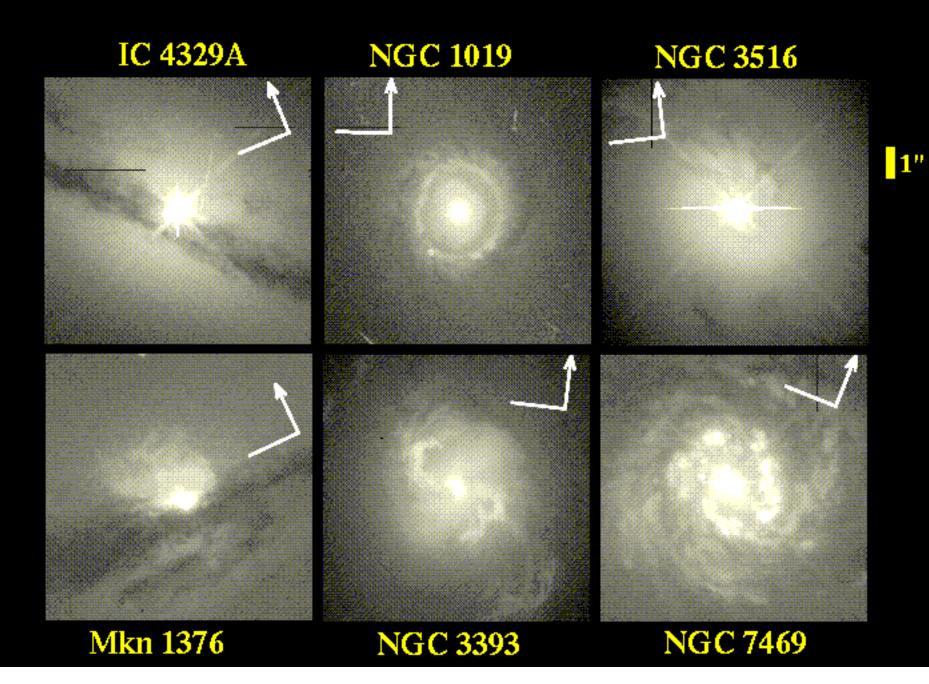
- Normal (boring) galaxies
- Active (fun) galaxies
- The extragalactic zoo: Quasars, Blazars, Radio Galaxies, BL Lacs, Seyferts, Optically Violent Variables, GHz Peaked Spectrum, Compact Symmetric Objects, ...
- Components of an active galaxy
 - Supermassive Black Holes
 - Jets
 - Hot spots and the evolution of radio galaxies



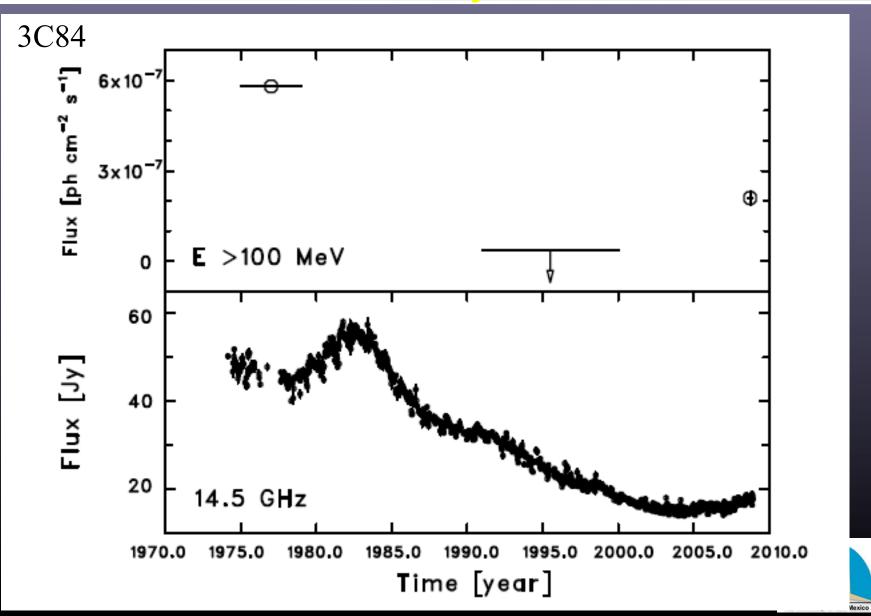
Seyfert vs Normal Galaxy



Seyfert Nuclei – HST Planetary Camera

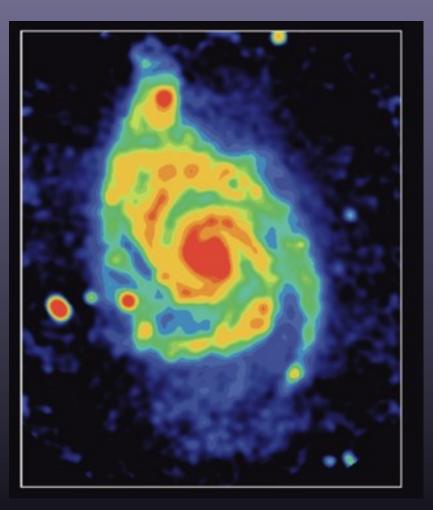


Variability



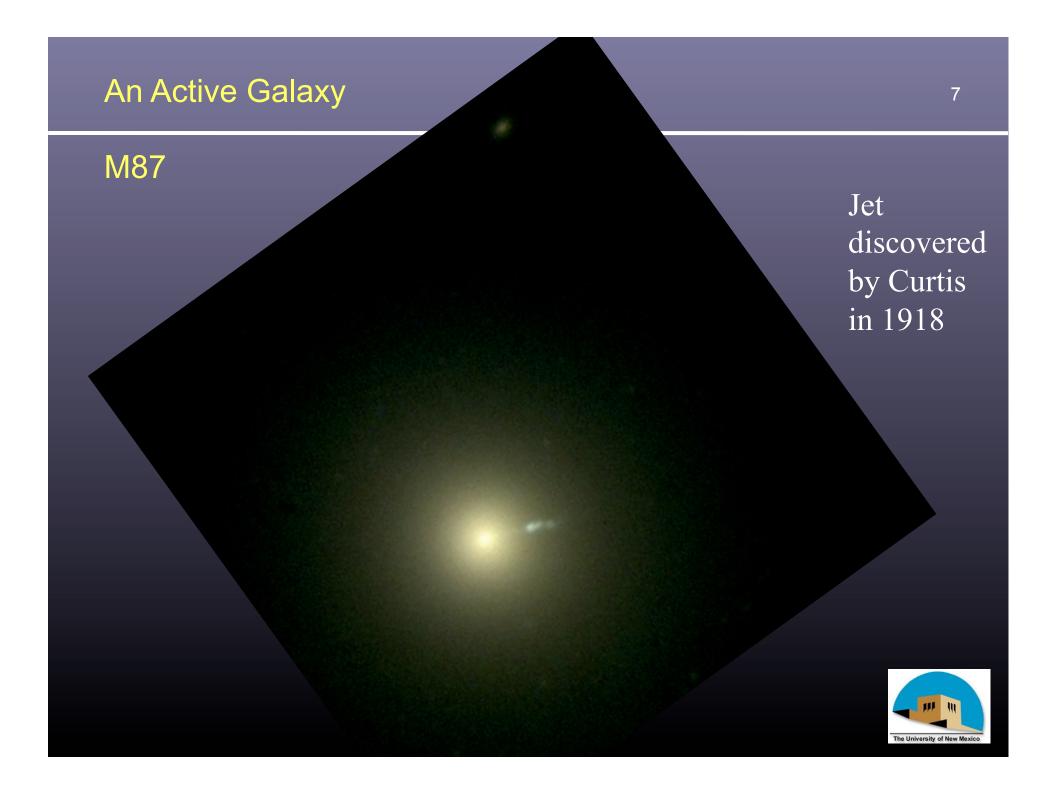
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Normal Galaxies in Radio Continuum



M51 - Beck et al.





An Active Galaxy

M87 HST Image

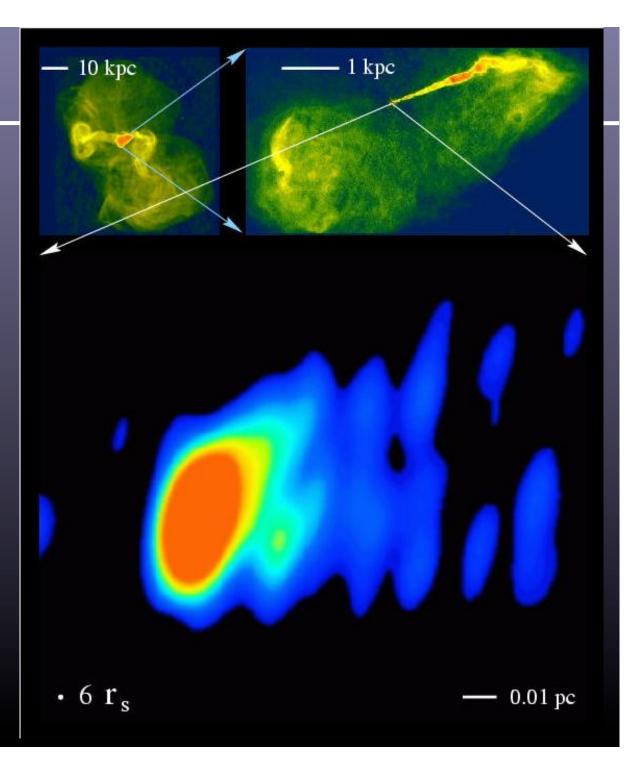
M87 Distance = 15 Mpc Jet length = 2 kpc



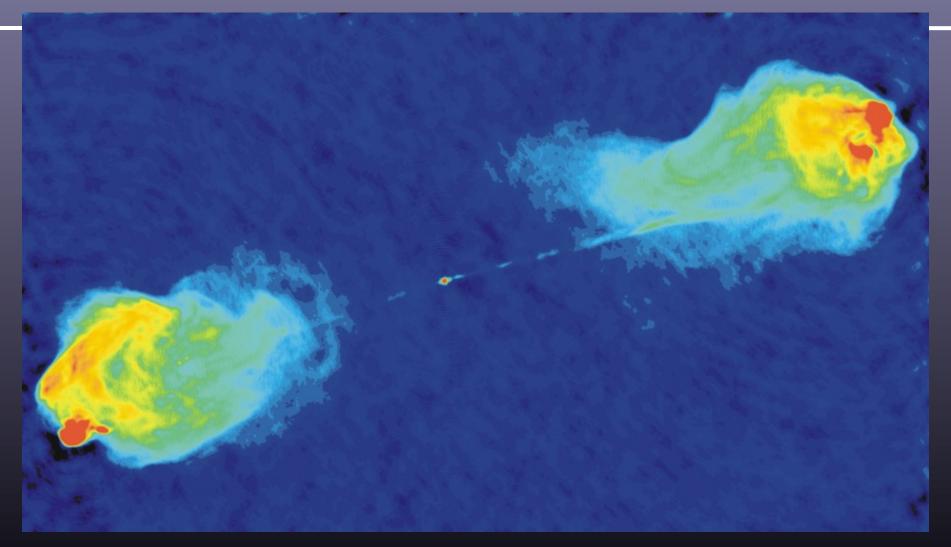
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An Active Galaxy

M87 VLA and VLBA Radio Images



Cygnus A – prototypical radio galaxy -- FR II

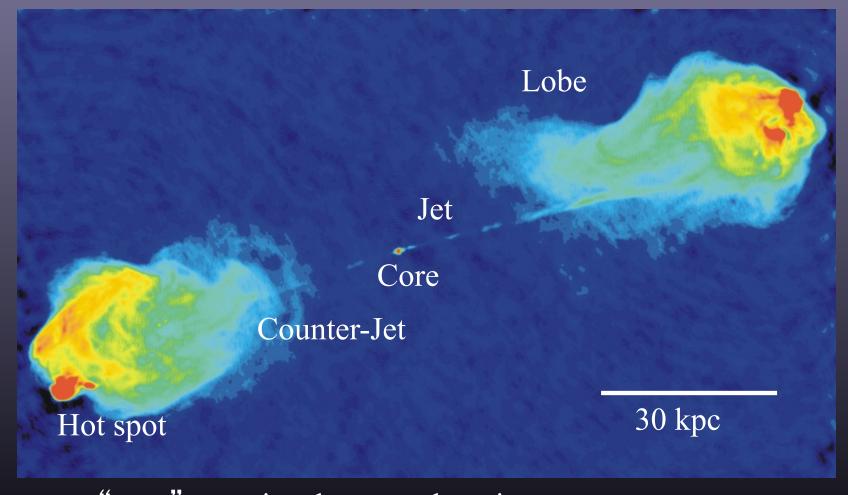


Carilli et al.



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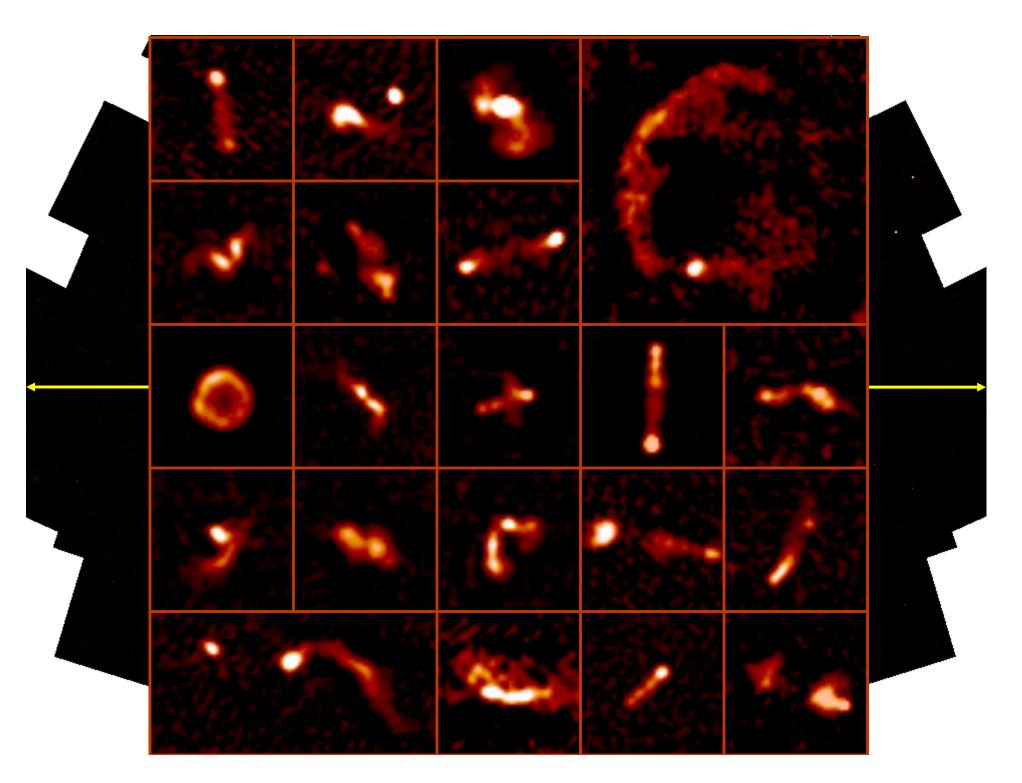
Components of an AGN





"core" contains the central engine

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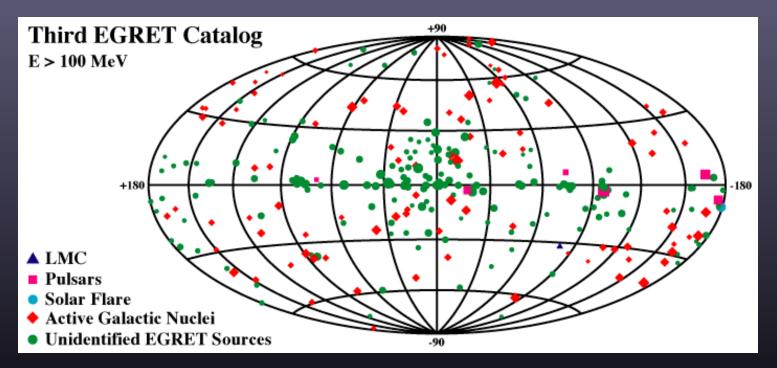


The EGRET Era

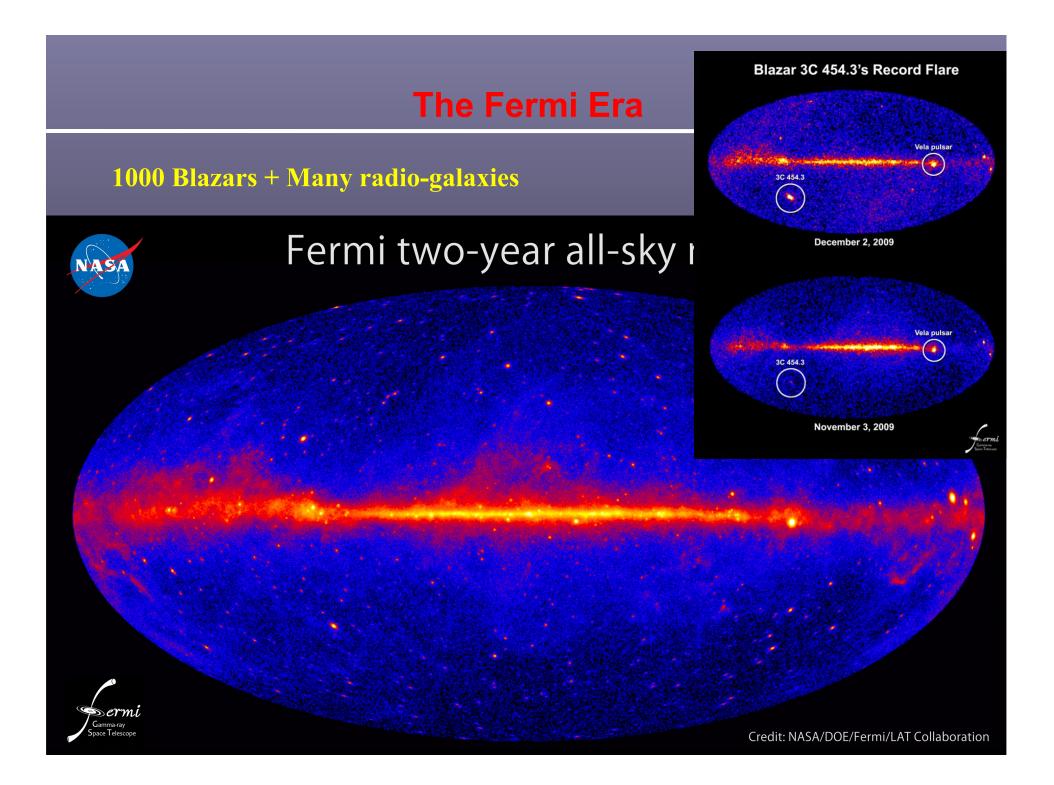
AGN dominate the extragalactic high-energy sky

• 67 Blazars

• marginal detection of a couple of radio-galaxies







Clicker Question:

What is the ultimate energy source in an Active Galaxy?

A: Nuclear fusion in stars.

B: Nuclear fission in stars.

C: Accretion of material onto a supermassive black hole.

D: Direct conversion of matter into energy in the hot spots.



Clicker Question:

The end of the Jet in a radio galaxy is usually marked with a:

A: hot spot.

B: lobe.

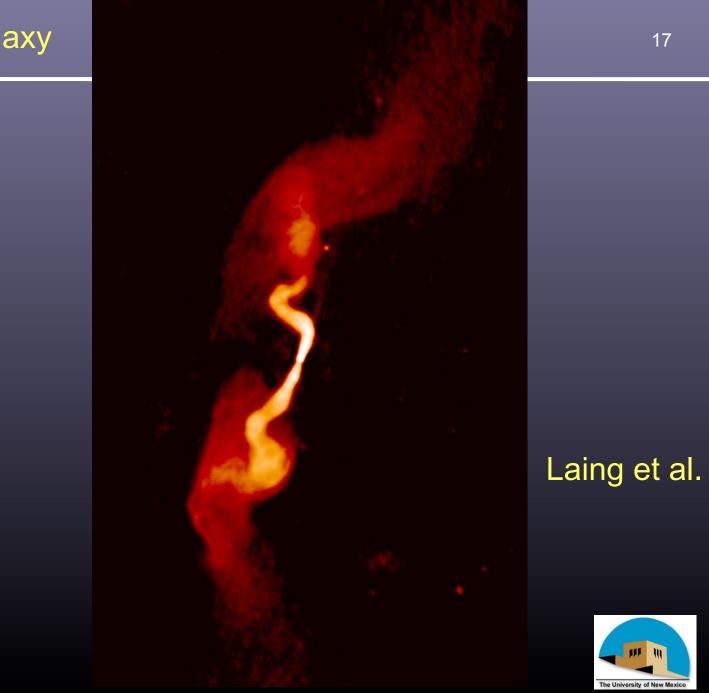
C: core.

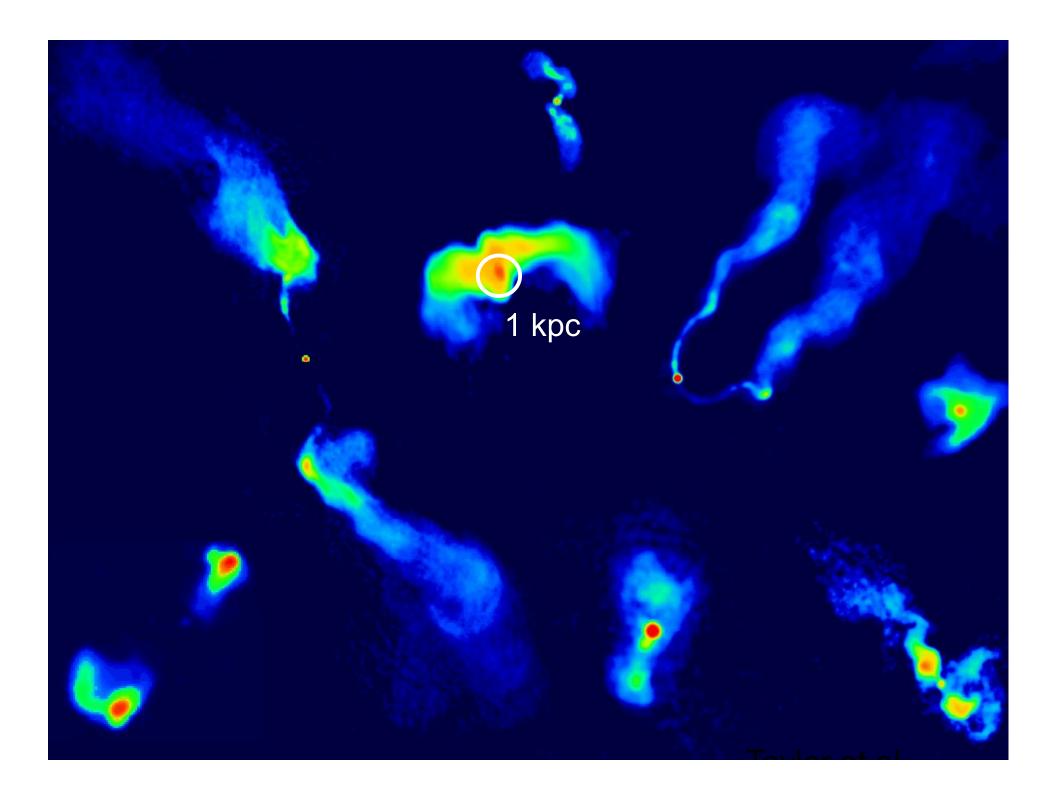
D: 'X'



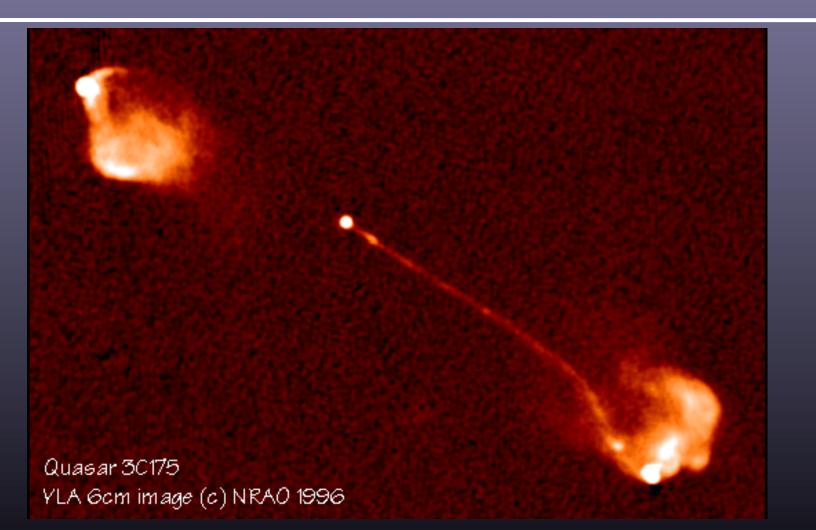
An Active Galaxy

3C31 FR I



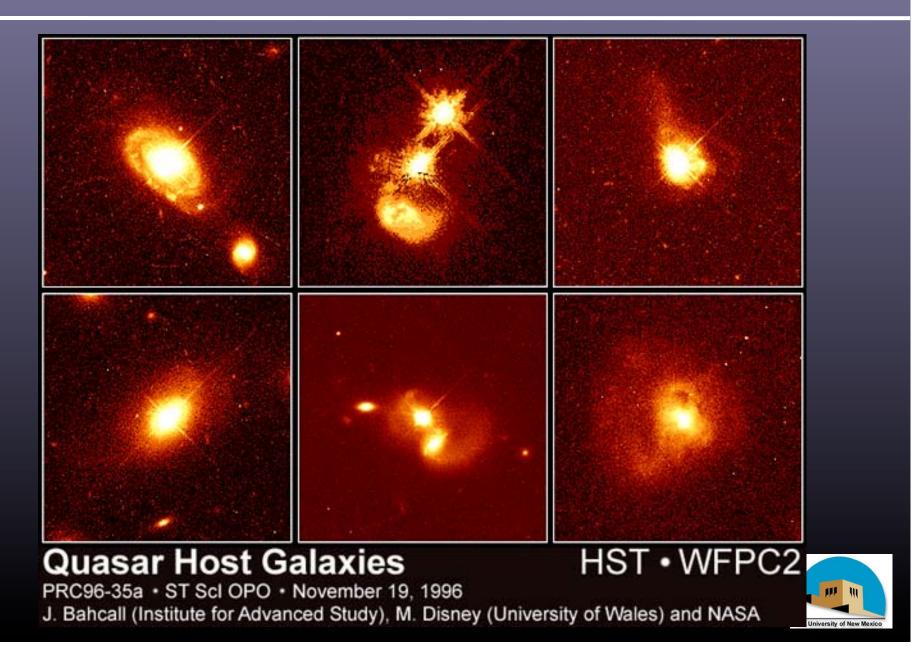


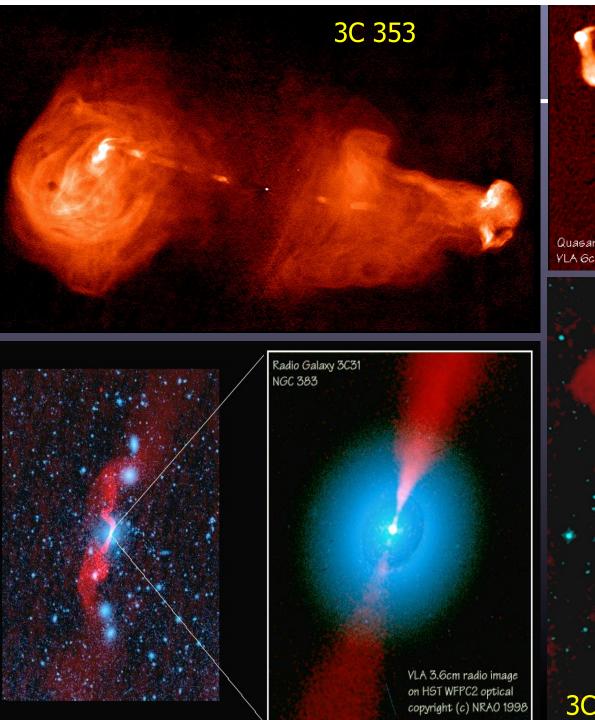
Quasar

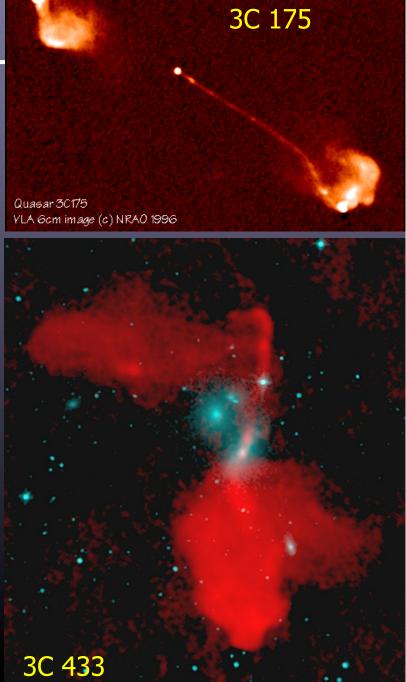




Quasar Host Galaxies

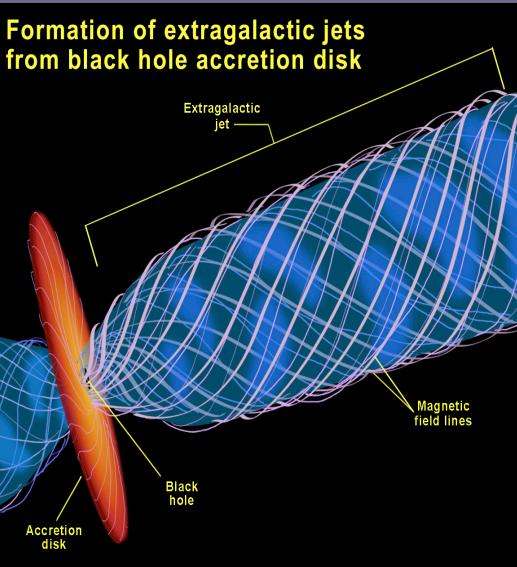




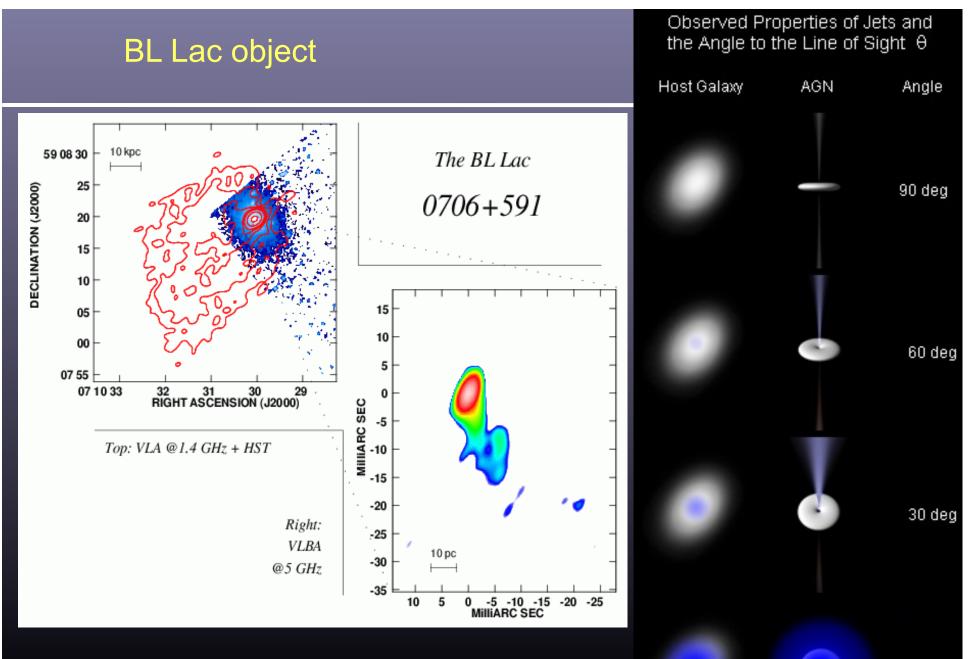


Radio Jets - Theory

- Accretion of gas onto a massive central black hole releases tremendous amounts of energy
- Magnetic field collimates outflow and accelerates particles to close to the speed of light



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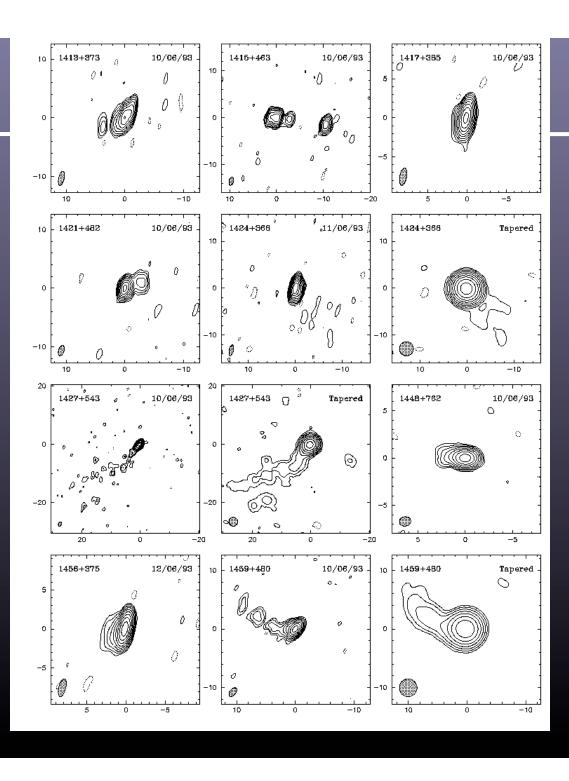


Giroletti et al 2004

0 deg

VLBA Images At 5 GHz

Taylor et al. 1994

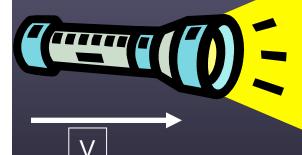


30 pc

24



Headlight effect



- Beam becomes focused.
- Same amount of light concentrated in a smaller area
- Flashlight appears brighter!



Superluminal motion

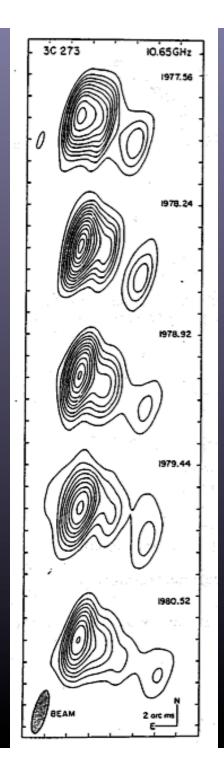
Pearson et al. 1981

Constant expansion observed for 3C273 at Rate = 0.76 +/- 0.04 milliarcseconds/year

Redshift z=0.158 so D = 704 Mpc Assuming $H_0 = 75$ km/s/Mpc

1 milliarcsecond = 2.5 pc at 704 Mpc

Velocity = 2.5 pc/mas * 0.76 mas/y = 2 pc/year * 3.26 light-year/pc = 6.5 light-year/year = 6.5 c !!!



Clicker Question:

Most compact jets are one-sided because:

A: The jets are intrinsically one-sided (like a jet from a hose).

B: The jets are strongly beamed and one side is pointing closer to us.

C: On one side they are compressed as they run into material in the interstellar medium.

D: Gravitational lensing by the supermassive black hole.



Clicker Question:

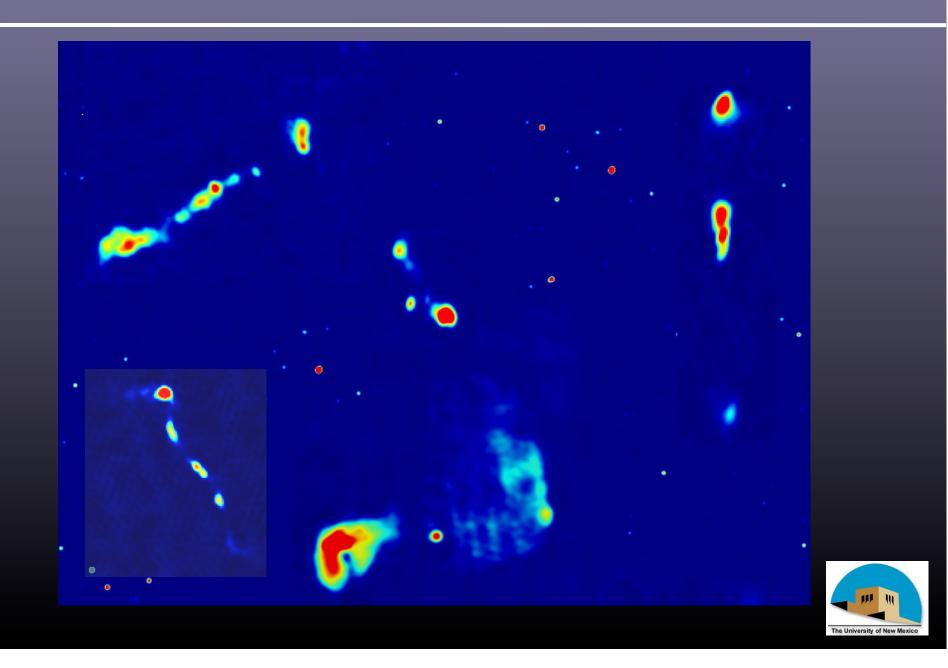
Some extragalactic jets have components that appear to move faster than light (superluminal motion). This happens because:

A: Its an illusion because the jets are pointed at us and moving close to the speed of light.

- B: The jets come out of a black hole.
- C: The jets are very light.
- D: The jets are made of tachyons.



Compact Symmetric Objects (CSOs)



Evolution

4C31.04

100 pc

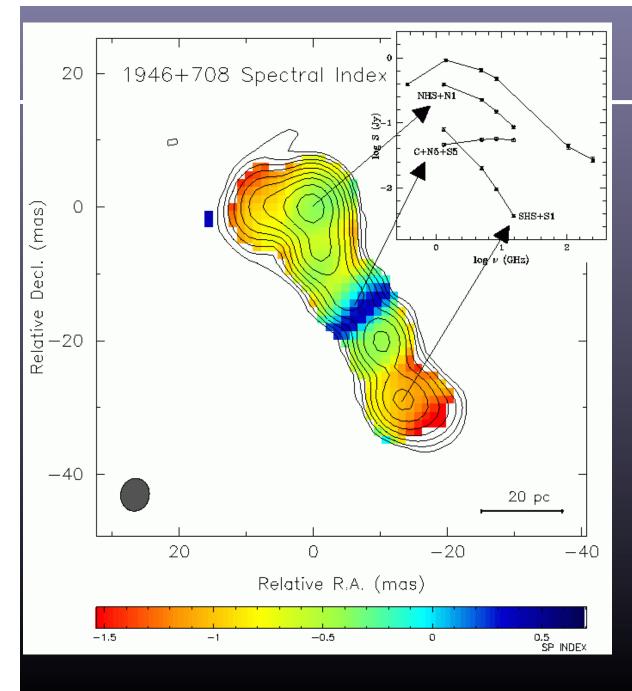




100 kpc



30



Free-free absorption in 31 1946+708

Peck & Taylor (2001) Spectral index map from 1.3/5 GHz VLBI observations



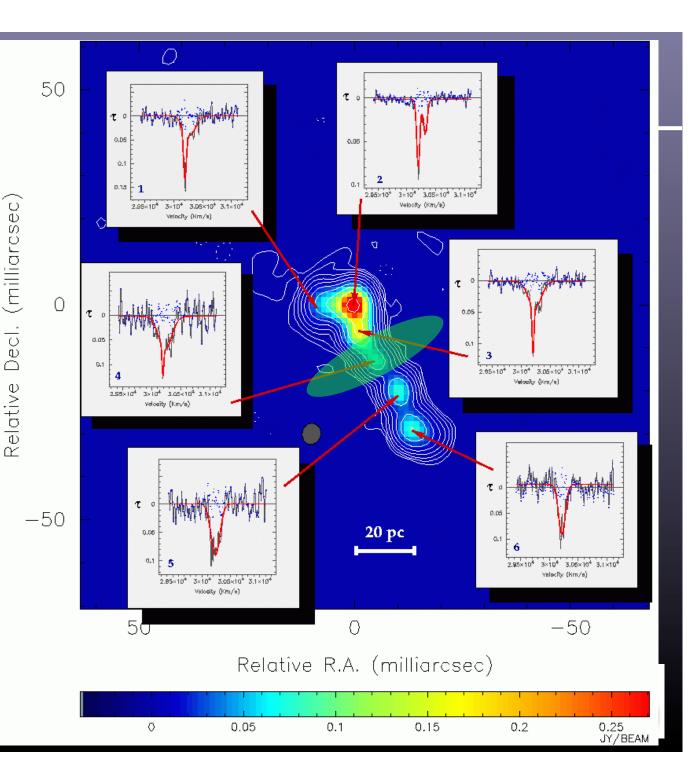
HI absorption in 1946+708

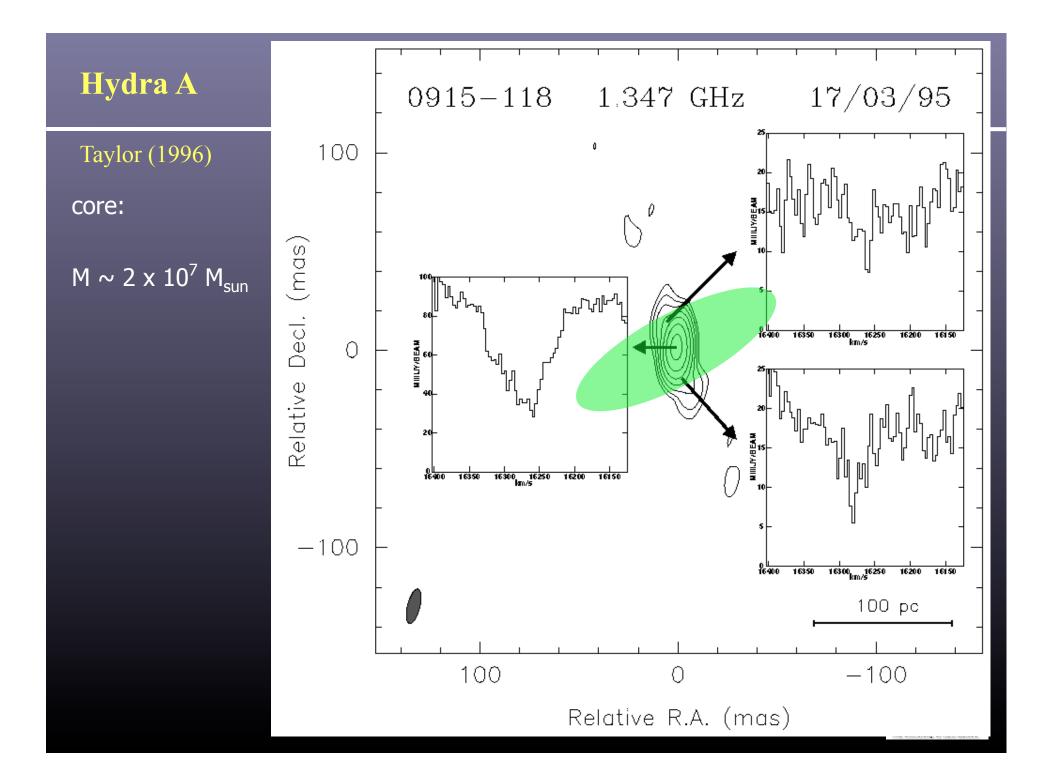
Peck & Taylor (2001)

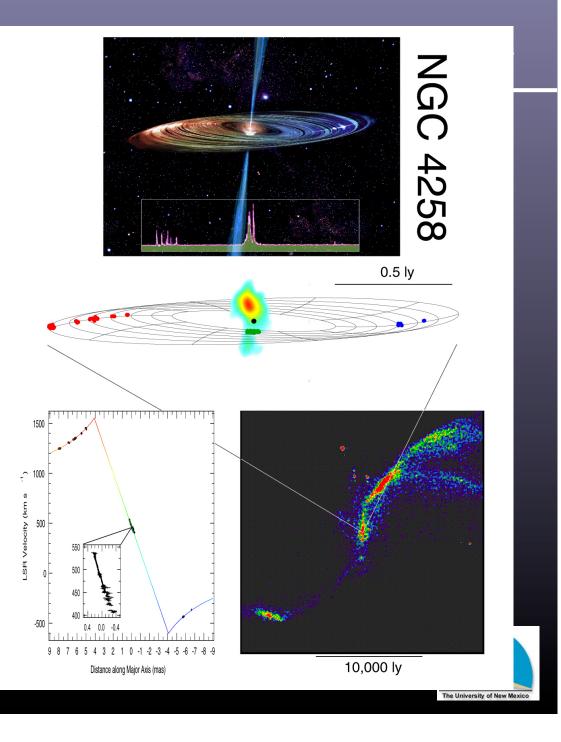
"Global" VLBI observations

core:

 $M \sim 10^8 M_{sun}$





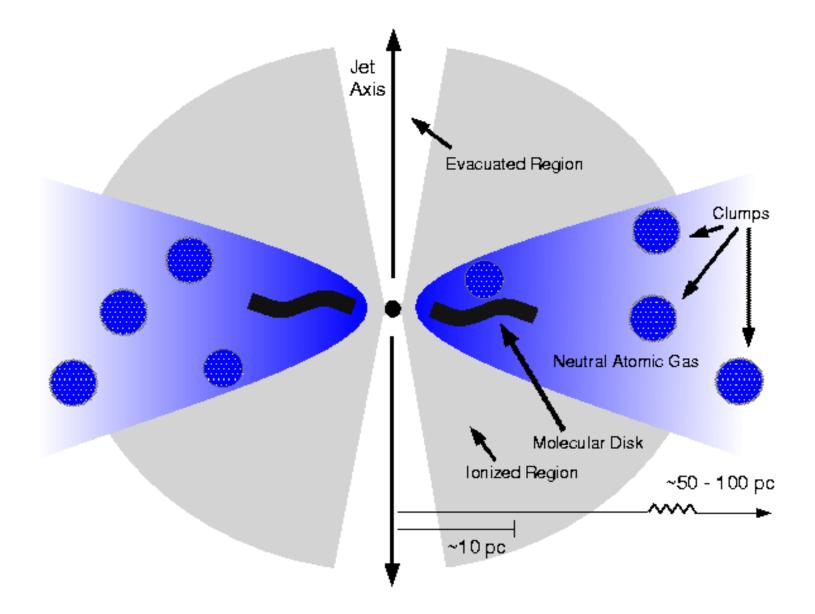


NGC 4258

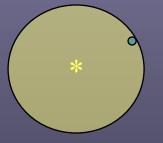
 Considered best evidence of a supermassive black hole
Can estimate central mass

Can estimate distance to host galaxy

image courtesy Lincoln Greenhill (see Miyoshi et al 1995 Herrnstein et al 1999)



Accretion



v = $(2GM/R)^{1/2}$ = c $(R_g/R)^{1/2}$

 $L = 0.5 M c^2 (R_g/R)$

$L = \varepsilon \dot{M} c^2$

c = speed of light v = velocity R = object radius R_g = gravitational radius ϵ = efficiency

 ϵ = 3 x 10⁻⁴ for a white dwarf, 0.15 for a neutron star ϵ up to 0.5 for a rotating black hole Compared to 7 x 10⁻³ for hydrogen fusion

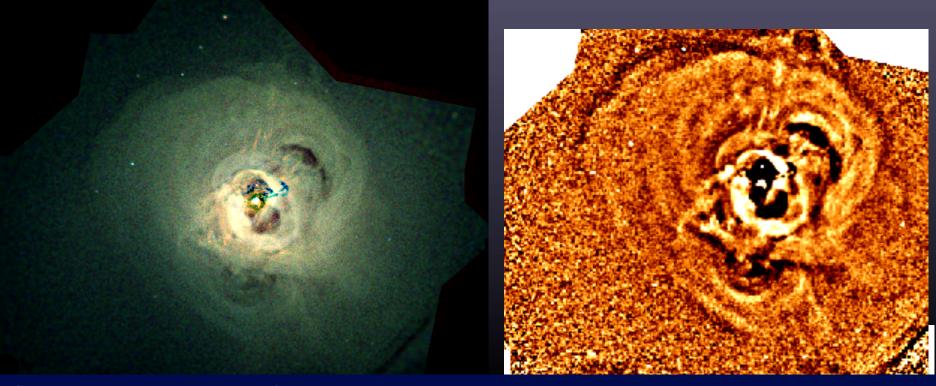
Accretion is a very efficient energy source



AGN Feedback

AGNs feedback energy to the cluster gas, prevent further cooling

Fabian, Forman: studies of nearby clusters ⇒ In Perseus and in M87, directed energy from jets is being redistributed via bubbles blown, compression waves ...



Chandra 900 ks image of Perseus cluster, unsharp-masked at right (Fabian et al. 2006)

NASA animation of sound waves produced by 3C84 38



Clicker Question:

What is the most efficient energy source in the Universe?A: Fusion of hydrogenB: Accretion onto a black holeC: Coal-fired power plantsD: Wind turbines.



Clicker Question:

What sources dominate the sky at radio wavelengths?

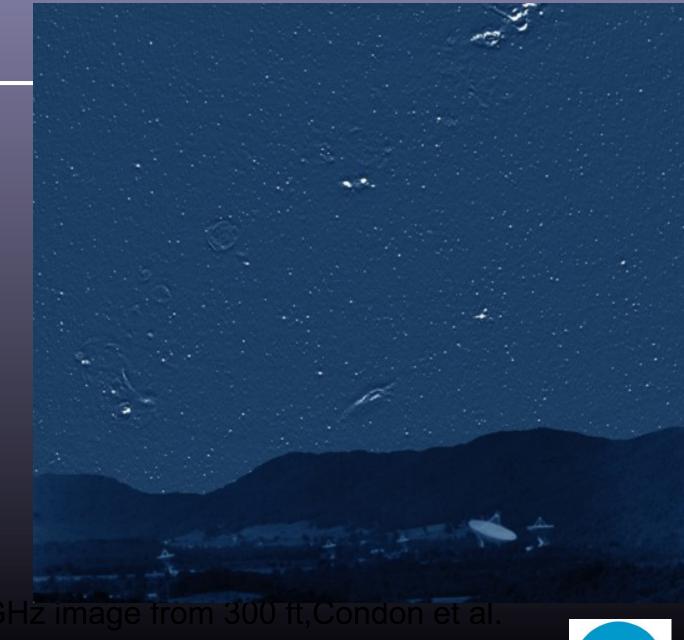
A: Stars

- B: Active Galaxies
- C: Normal Galaxies
- D: Globular clusters



The Radio Sky

5 GHz 300 ft telescope





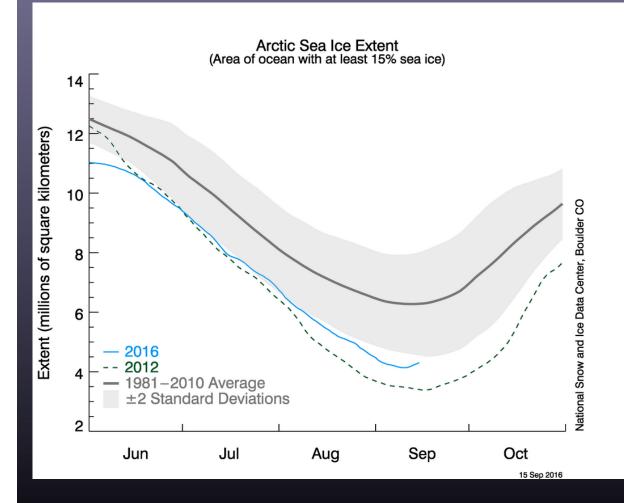






FOOD FOR Thought

Melting Sea Ice



- An area twice the size of Texas has melted away since 1979 (over 50% decrease).
- Ice 40% thinner. (Rothrock,D.A, et al. 1999)
- Ice free summer in 20 years?
- Bering Sea Ice Sheet also retreating (Science 3/10/06).



