



Transients at Low Radio Frequencies

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TIME DOMAIN

VS

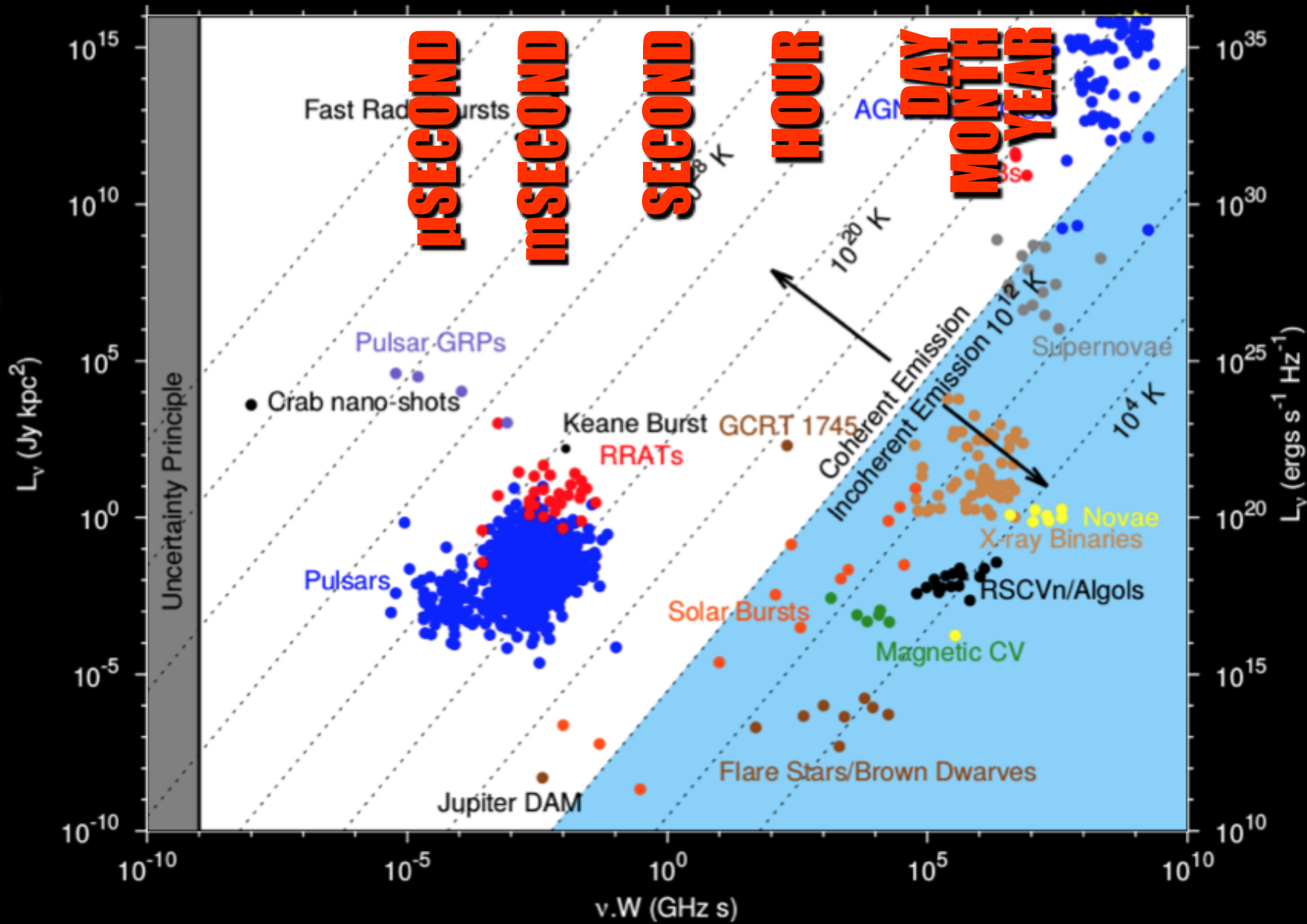
TRANSIENT

BOOM!

whhhhooooossss

sshhh

$$T_B \sim \frac{L}{fW}$$



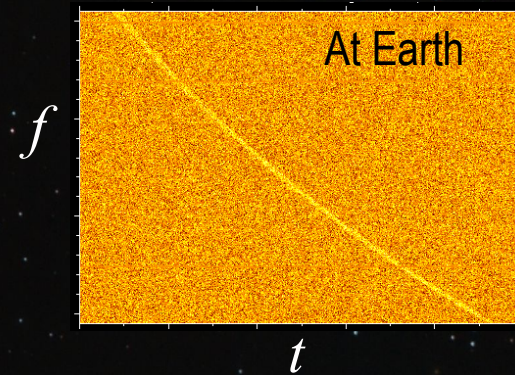
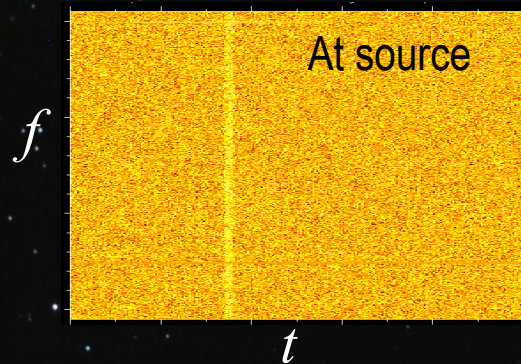
Fender et al (2015)

Unique regimes: Slow and Fast Transients

	Physics:	Instruments/strategy:
FAST	<ul style="list-style-type: none">★ Intervening media★ Huge brightness temperatures	<ul style="list-style-type: none">★ Phased/beamformed★ Computation heavy★ Enormous data sets★ Real-time identification
BOTH	<ul style="list-style-type: none">★ CATAclysms!	<ul style="list-style-type: none">★ Commensal★ Automated classification/ machine learning
SLOW	<ul style="list-style-type: none">★ Ejecta afterglows★ Relativistic jets★ Kinetic feedback measurements	<ul style="list-style-type: none">★ Image plane

Dispersion

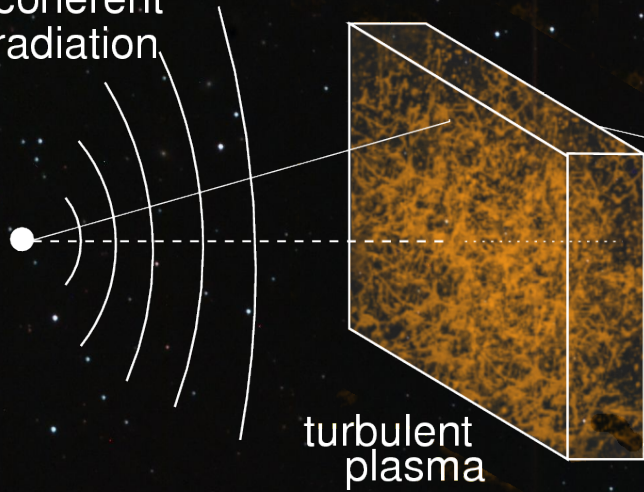
$$\sim f^{-2}$$



Scattering

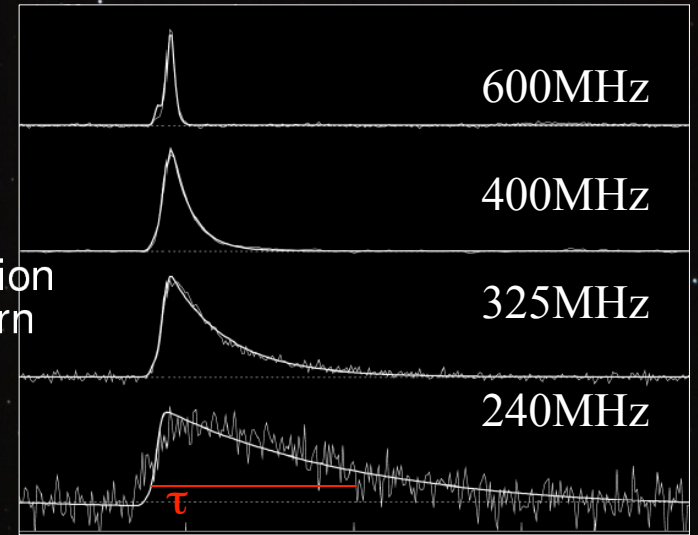
$$\sim f^{-4}$$

spatially
coherent
radiation



randomly
distorted
wavefronts

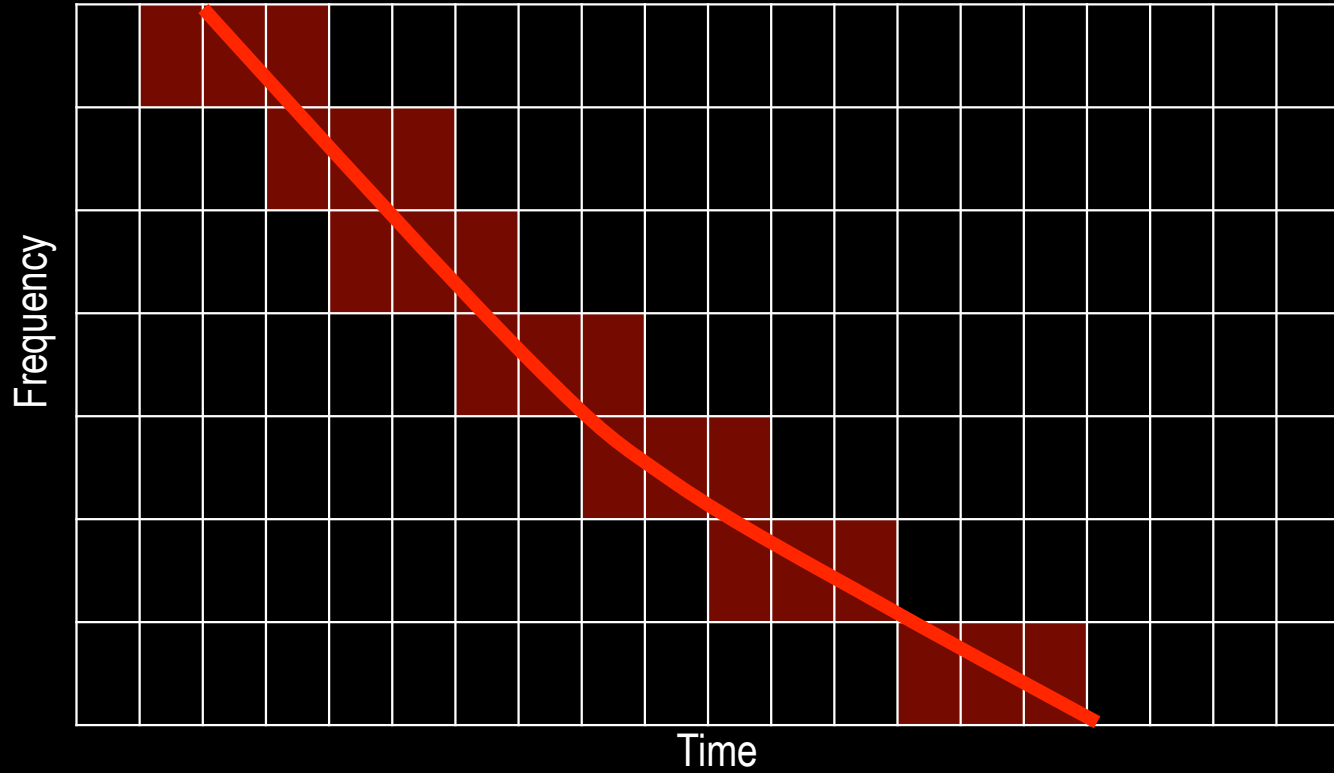
diffraction
pattern



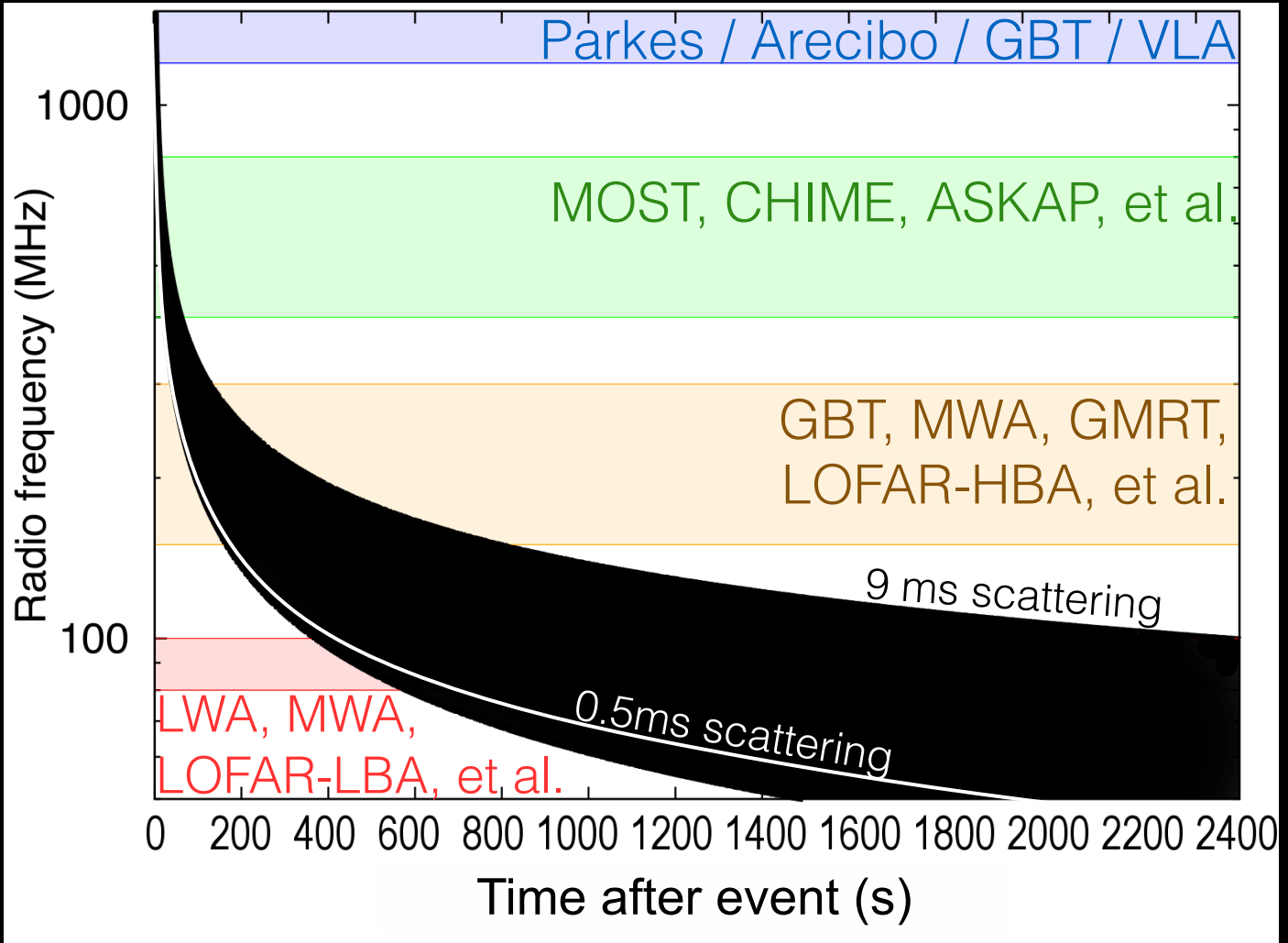
Lorimer+Kramer05

Instrumental Requirements

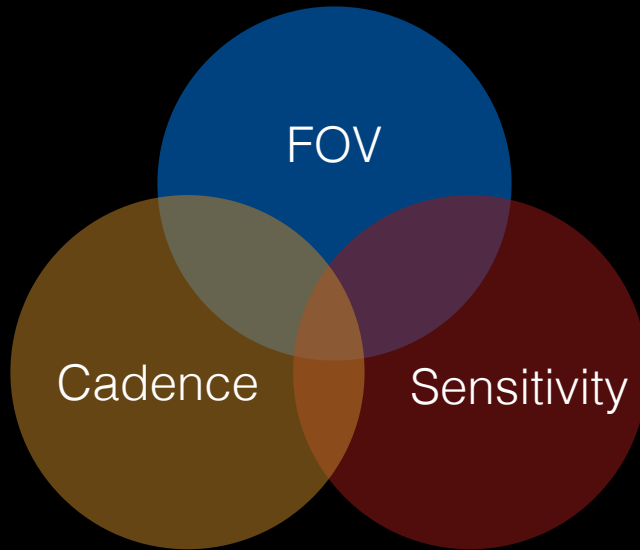
$$[\text{Detected width}]^2 = [\text{Intrinsic width}]^2 + [\text{Scattering}]^2 + \underbrace{[\text{Sampling Time}]^2 + [\text{Dispersion error}]^2}_{\text{Instrumental broadening}}$$



Dispersion:
1000 pc/cc

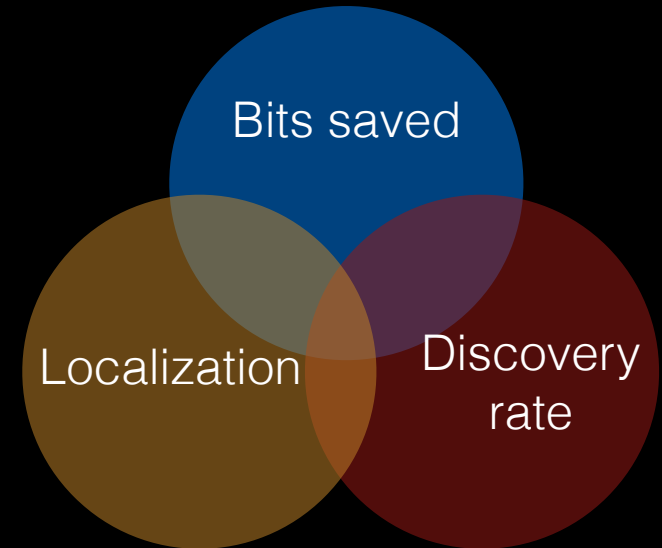


General transients:



- ★ FOV
- ★ Sensitivity
 - ★ Bandwidth
 - ★ Collecting area
 - ★ Freq. selection
- ★ Cadence

Fast transients:



- ★ FOV
- ★ Sensitivity
 - ★ Bandwidth
 - ★ Collecting area
 - ★ Freq. selection
 - ★ Sample time
 - ★ Channelization
- ★ Polarization
- ★ Localization
- ★ Huge supercomputer

So what do we know so far?

X-Ray Binaries



Matter falling from a star companion onto a black hole



Internal collapse of a star

Brown Dwarfs, Flare Stars



Outbursts

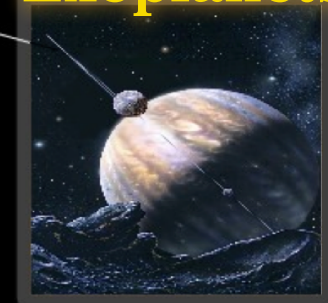
Tidal Disruption



Disruption of stars by black holes -

TRANSIENTS

Exoplanets



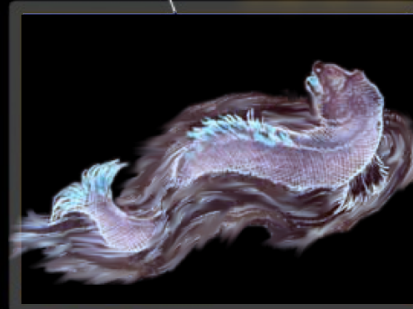
Flares

Active Nuclei

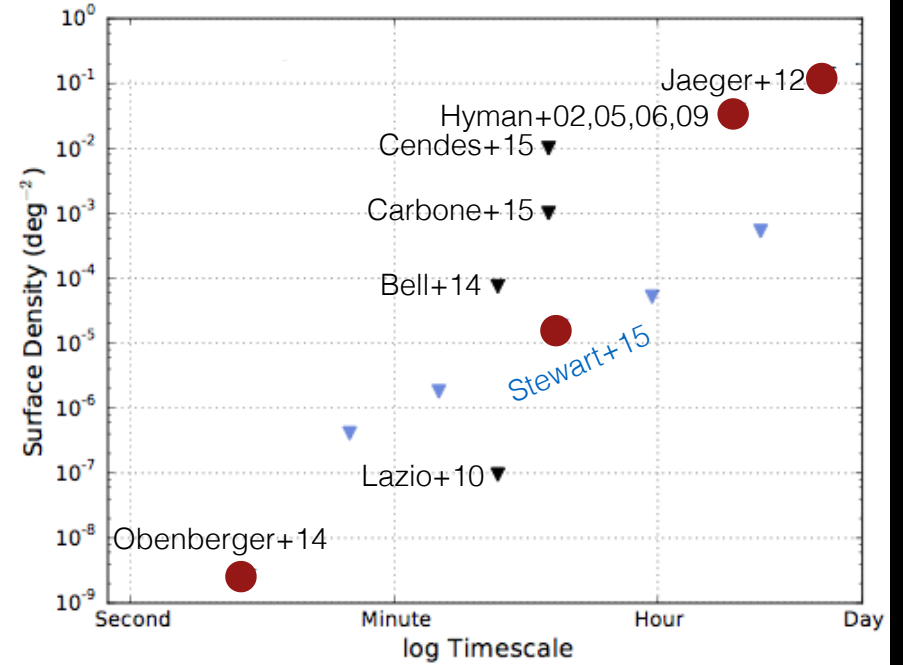
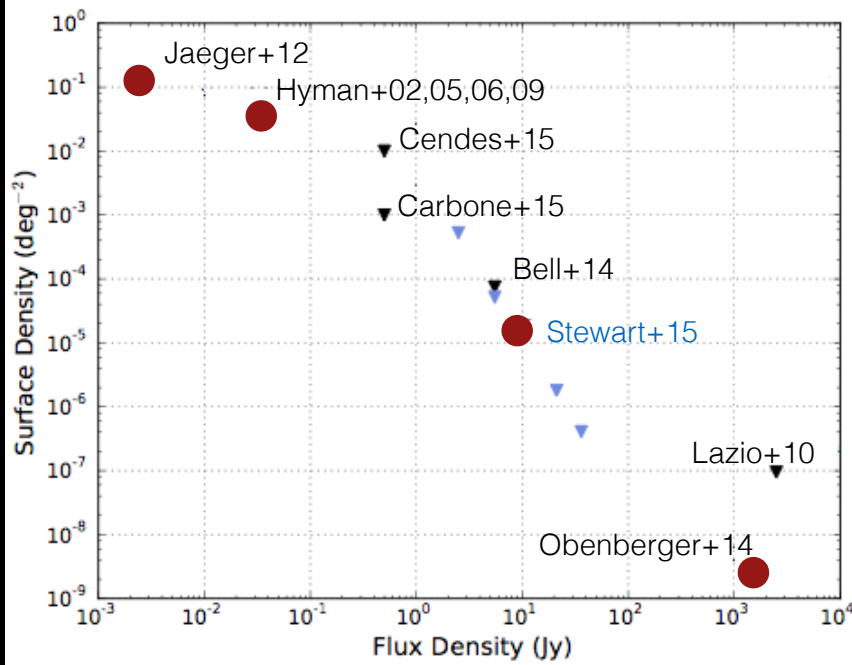


Matter falling onto a super-massive black hole in the centre of a galaxy

?!?!?!!!



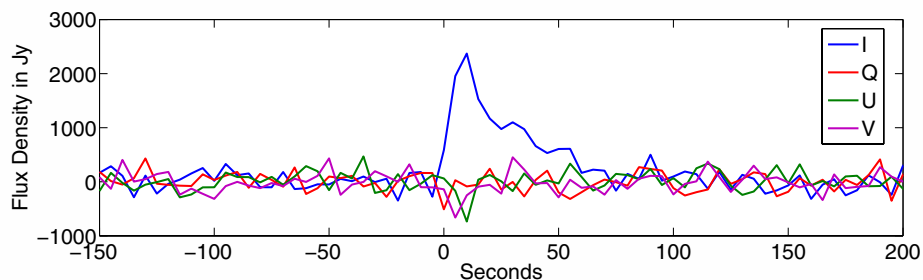
Blind Discoveries!



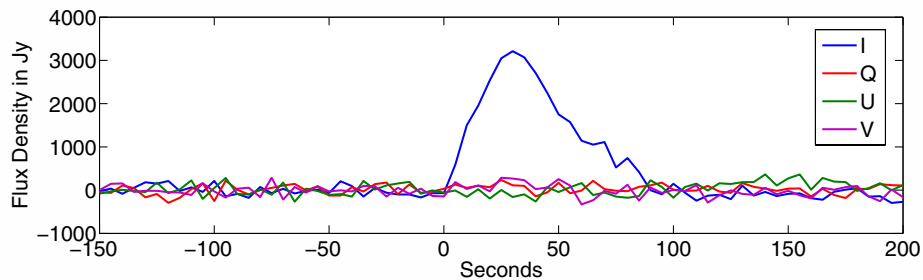
Stewart et al. (2015)

Blind

37.9 MHz
DM < 450 pc/cc

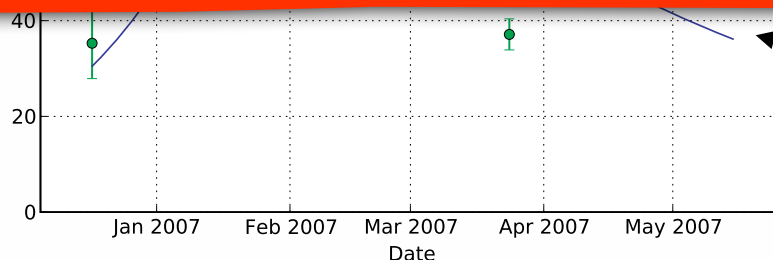


29.9 MHz
DM < 250 pc/cc



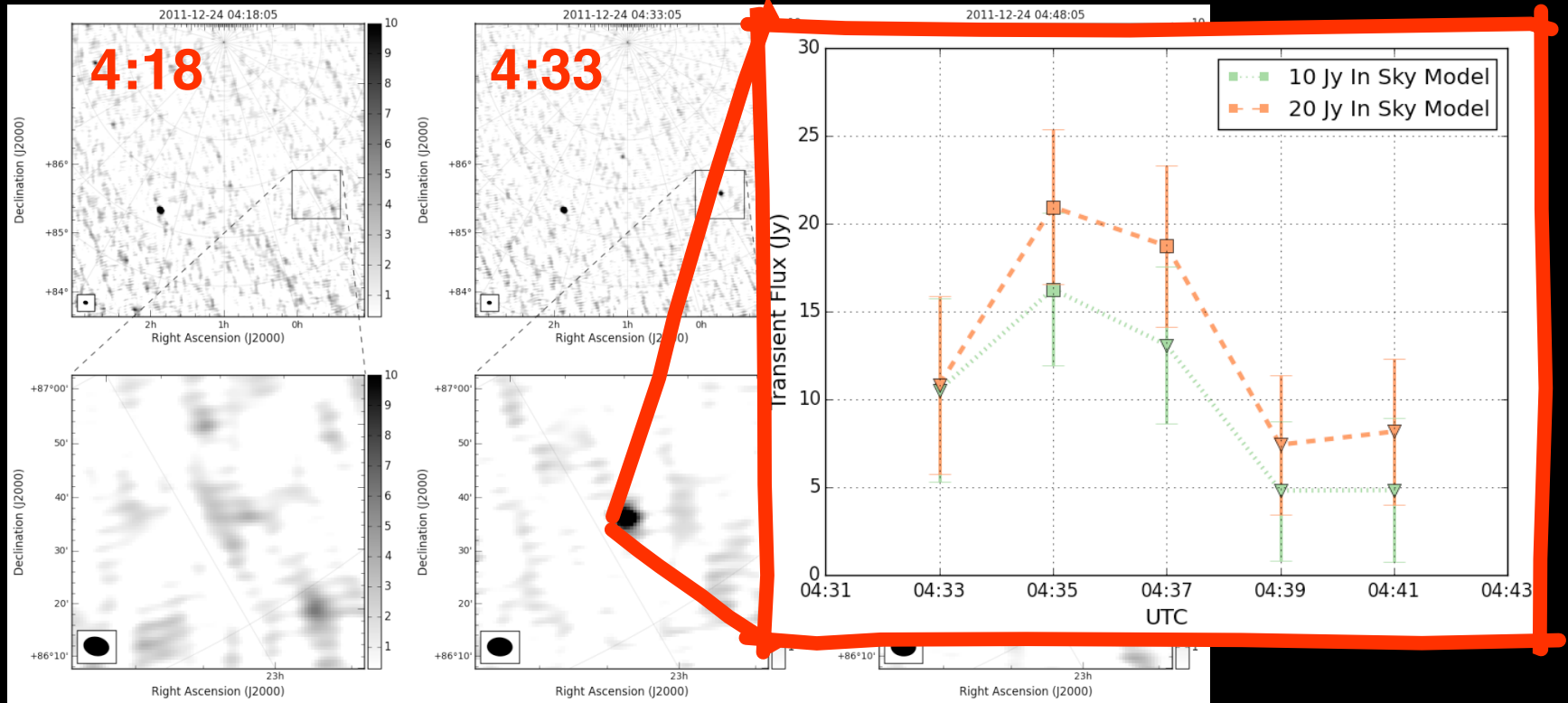
Obenberger
et al.(2014)

Hyman+
GC surveys



GCRT 1742-3001:
Six-month variable; steep
spectrum

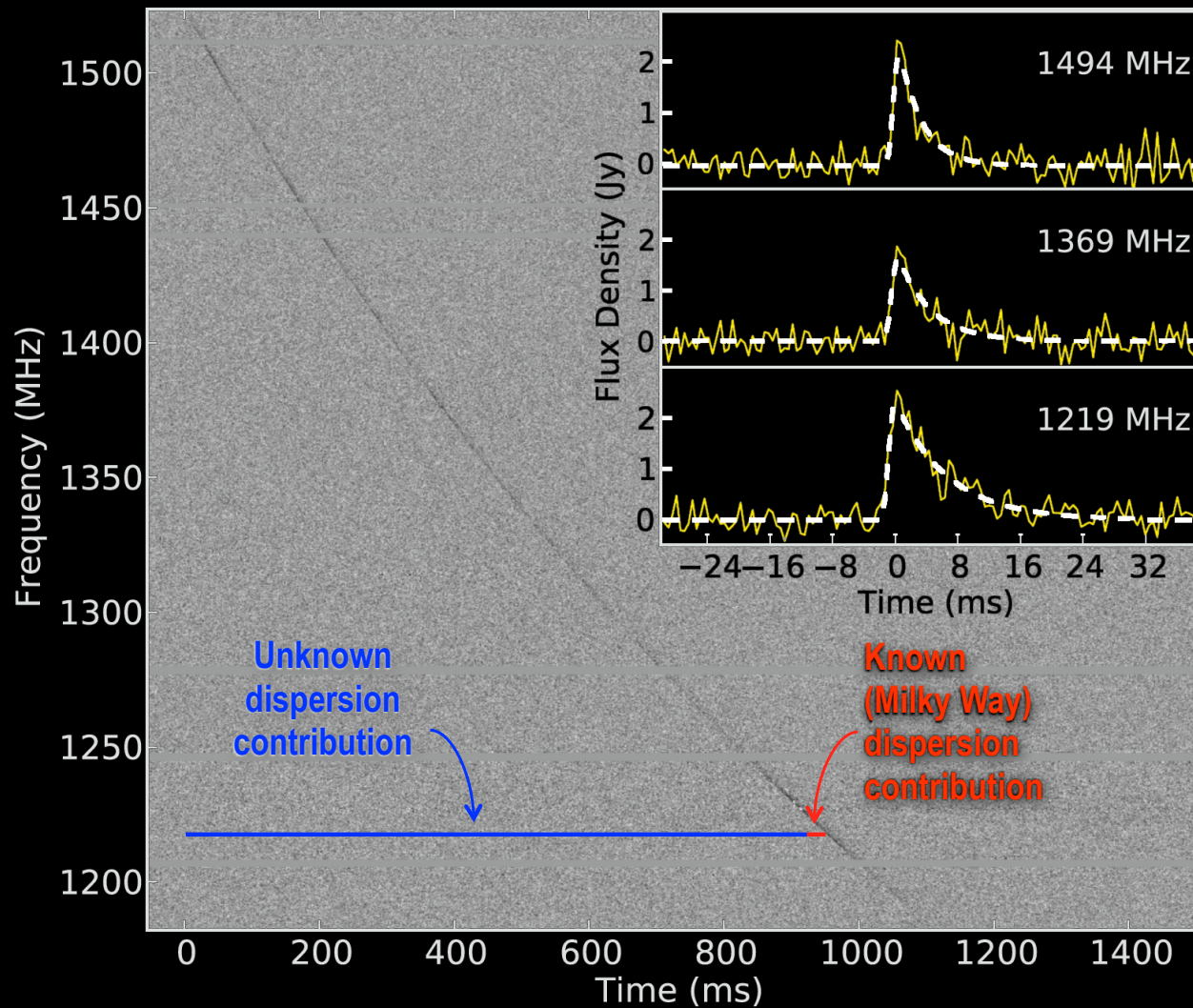
Christmas eve 2011, North (celestial) Pole



Stewart et al. (2015)

The future...

Fast Radio Bursts (FRBs)

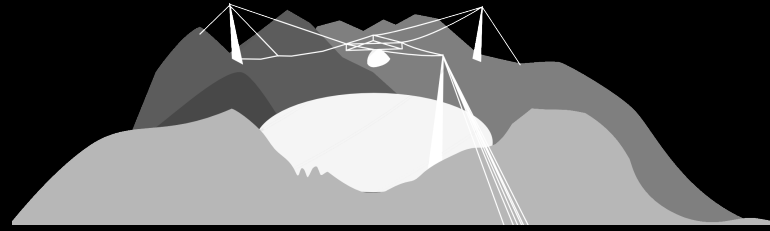


FRB fast facts

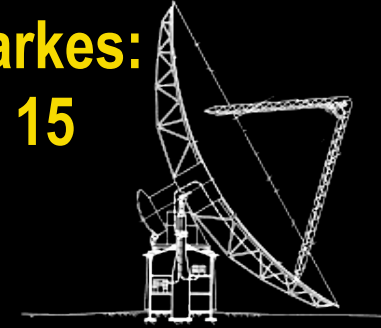
GBT: 1



Arecibo: 1

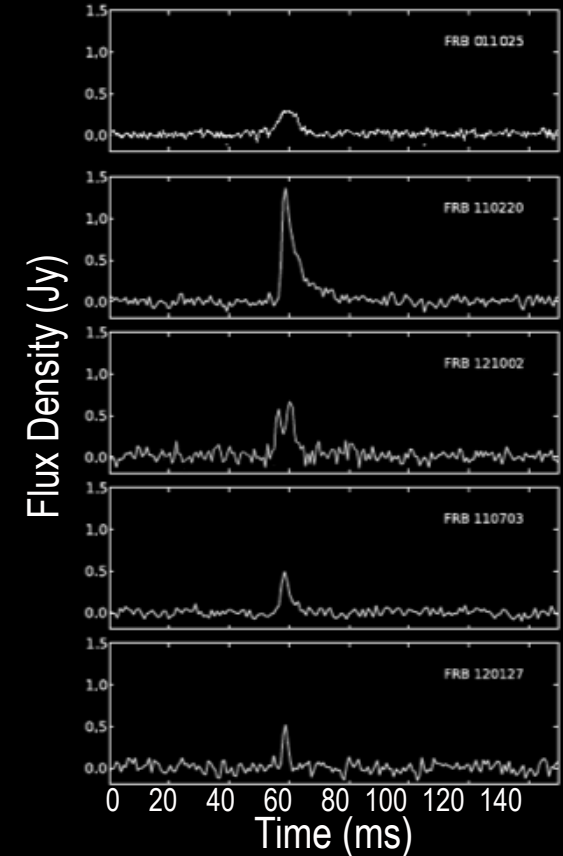


Parkes:
15

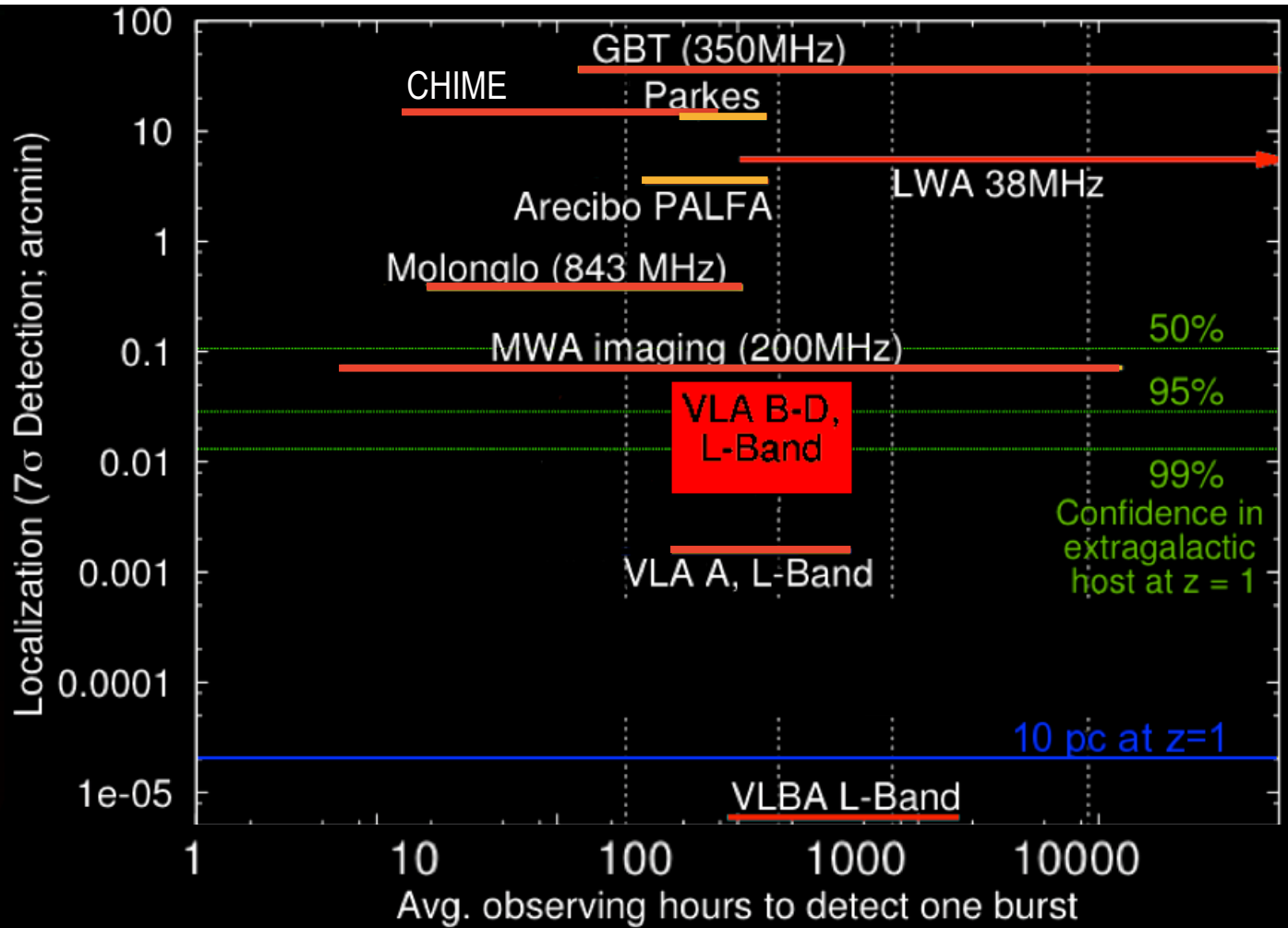


AT 1 GHz:

- ★ 17 detected (10 published)
- ★ $f = 0.7 - 1.5$ GHz
- ★ 2500-10000/sky/day
- ★ Extragalactic
- ★ ~50% scattered, $\langle \tau \rangle \sim 7$ ms

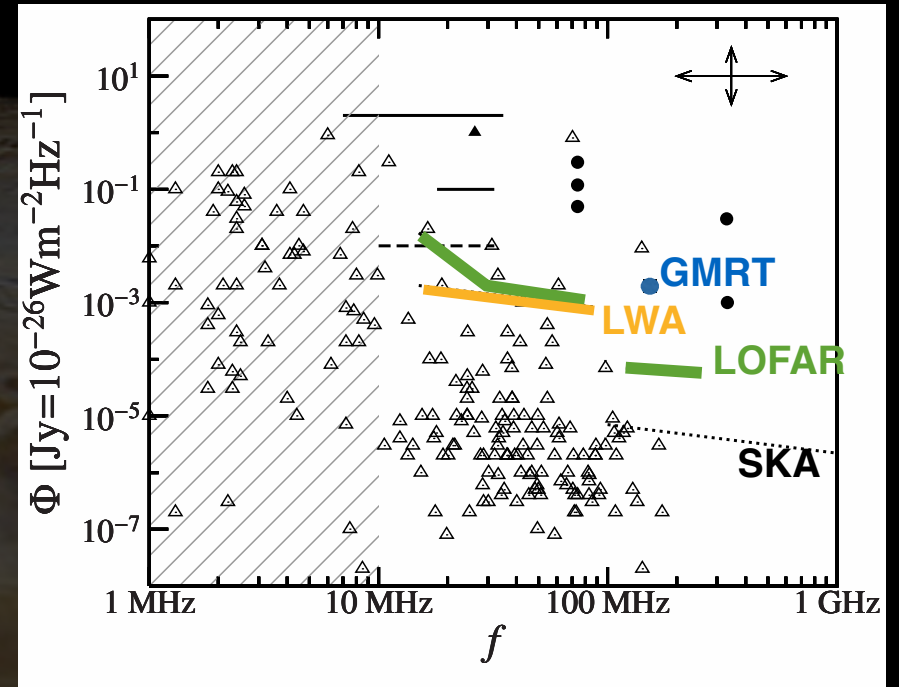


FRB optimization



Future: Hot Jupiters

- ★ Jupiter: magnetospheric cyclotron masers
- ★ Extrasolar planets: same processes?



Greissmeier et al. (2007)

Future: Multi-messenger

- LIGO/Virgo: detects a signal
- Prompt radio signal delayed by up to minutes, hours
- NS-NS merger, Cosmic string cusps

Yancey et al. (2015)

Conclusions

- Transient science
 - Huge range of targets
 - Afterglows, jet phenomena, exoplanets, intergalactic medium, and more
- Discoveries are commencing!
- Unexplored parameter space:
coherent one-off events at any distance