

Precision Primary Beam Mapping with ECHO

Daniel Jacobs
NSF Fellow — ASU

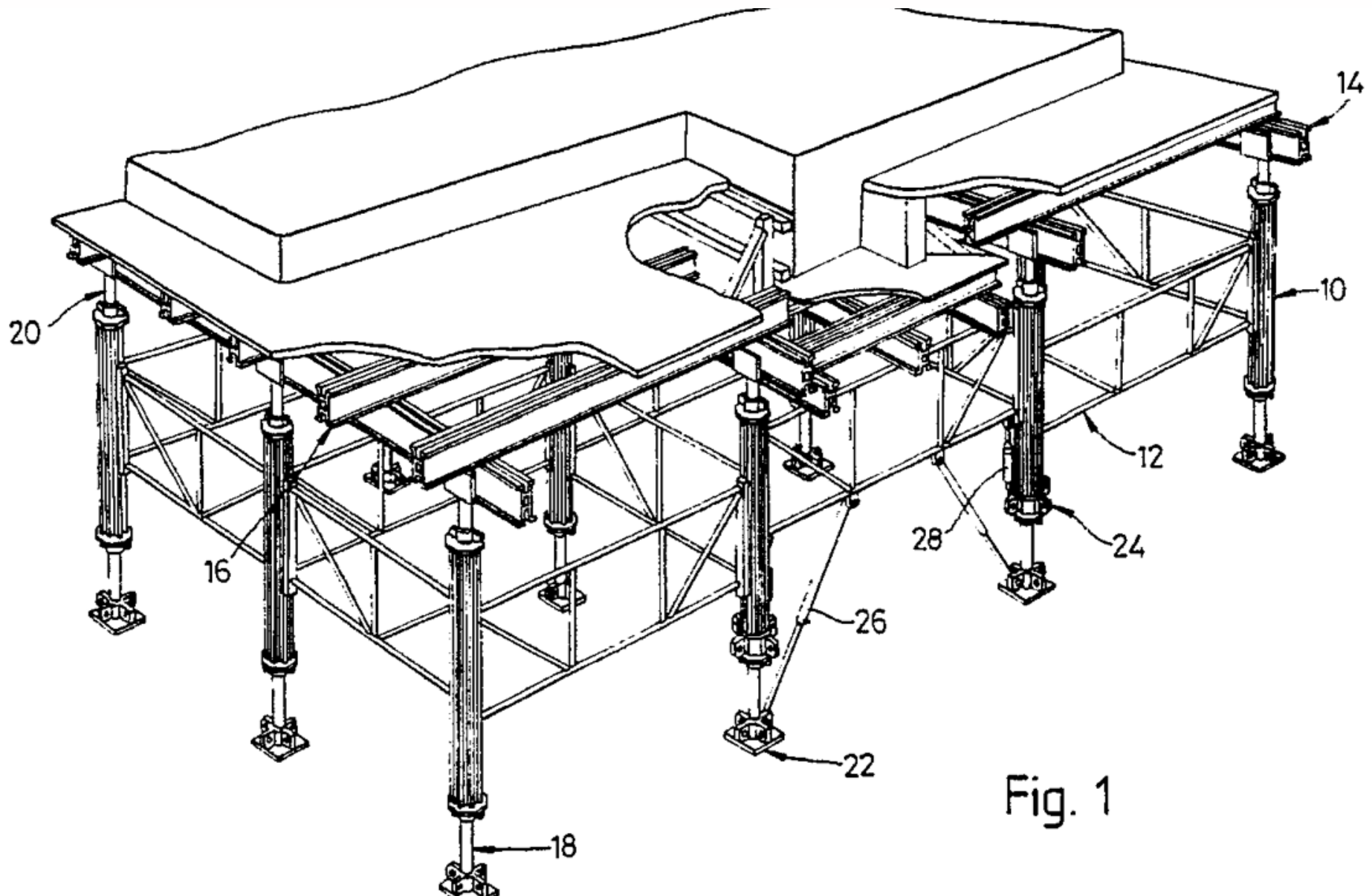


Fig. 1

THE EXTERNAL CALIBRATOR FOR HYDROGEN OBSERVATORIES

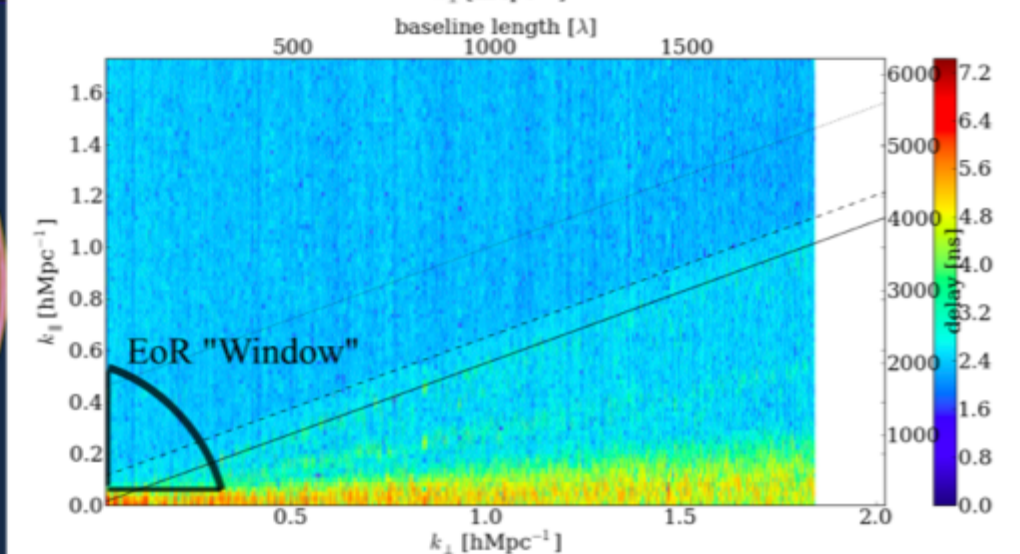
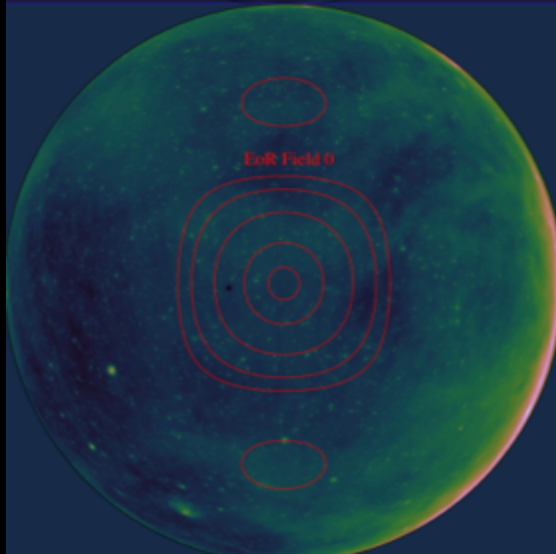
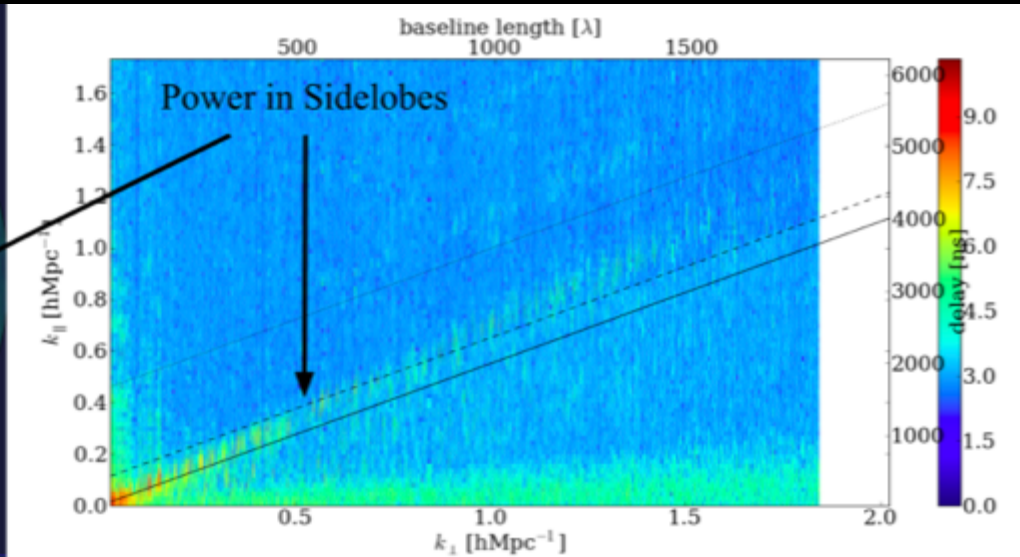
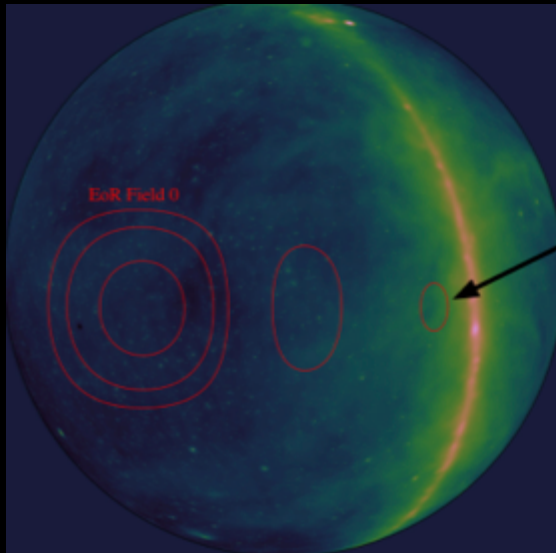
ASU: D. Jacobs, J. Bowman

J. Burba, L. Turner, B. Stinnett, M. Busch, M.
Leatham, V. Serrano, M. Denney, D. Nelson

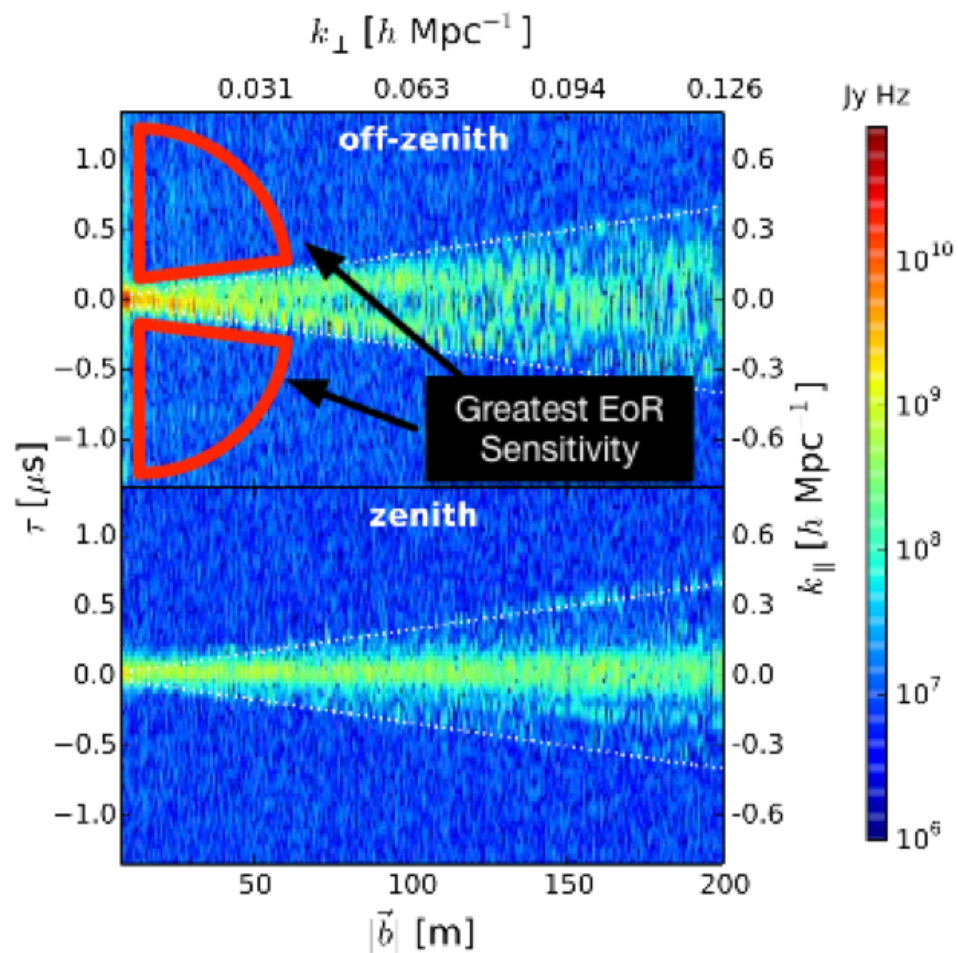
MIT: A. Neben



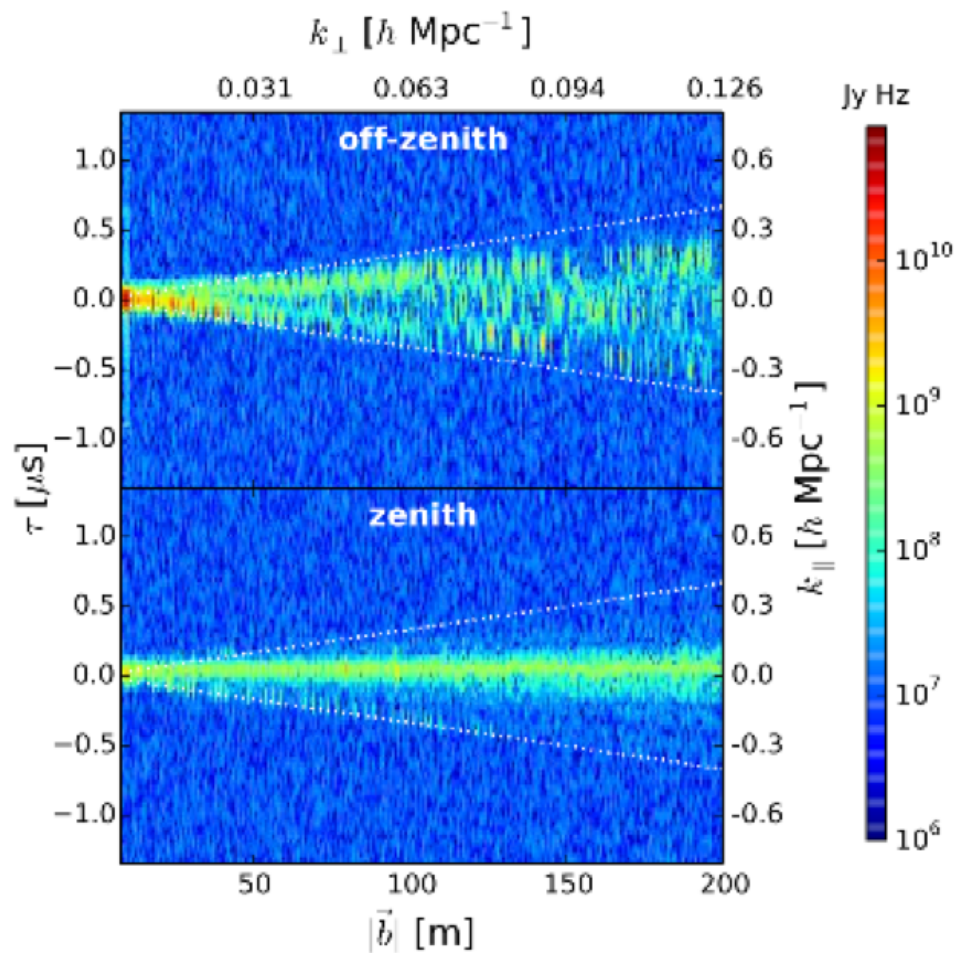
THE BEAM IN 21CM POWER SPECTRA

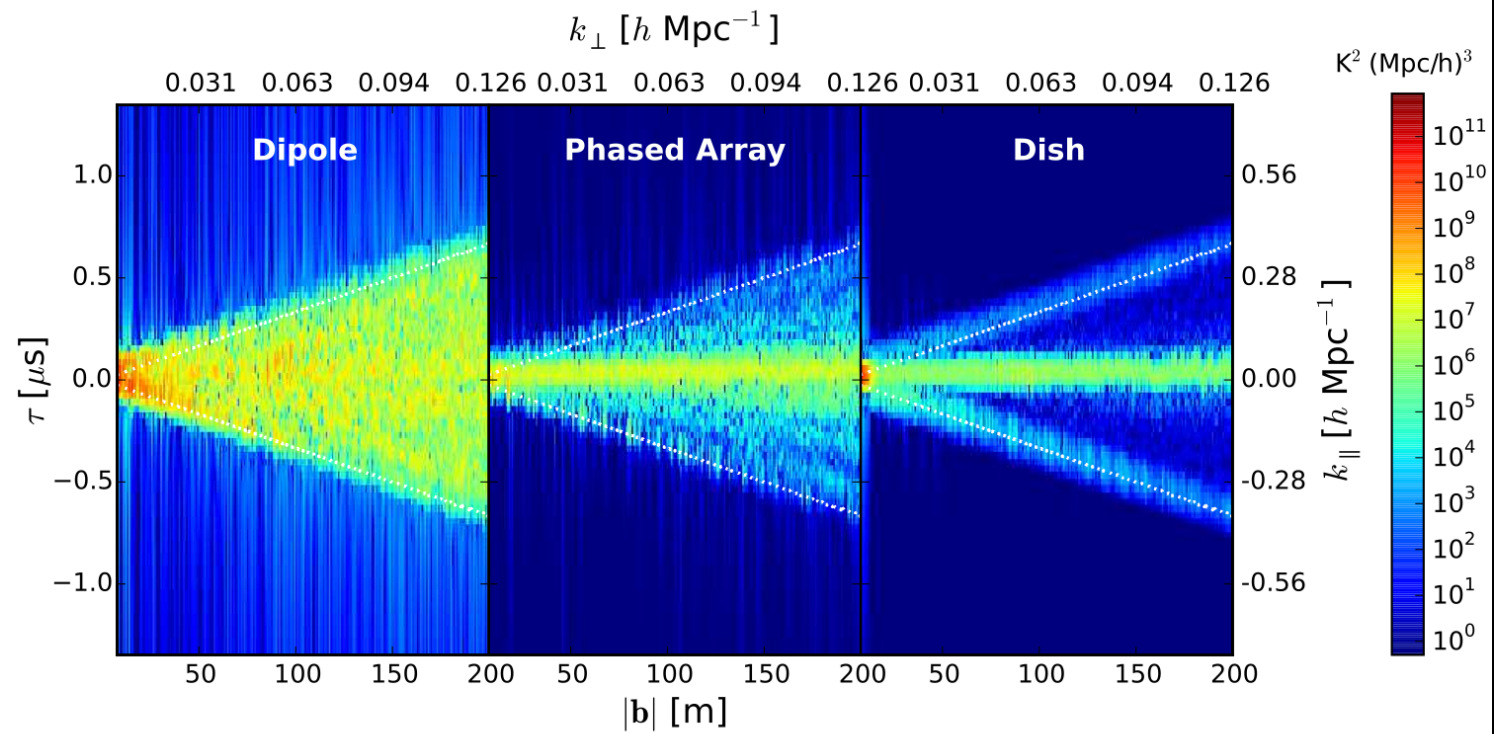


Data



Simulation

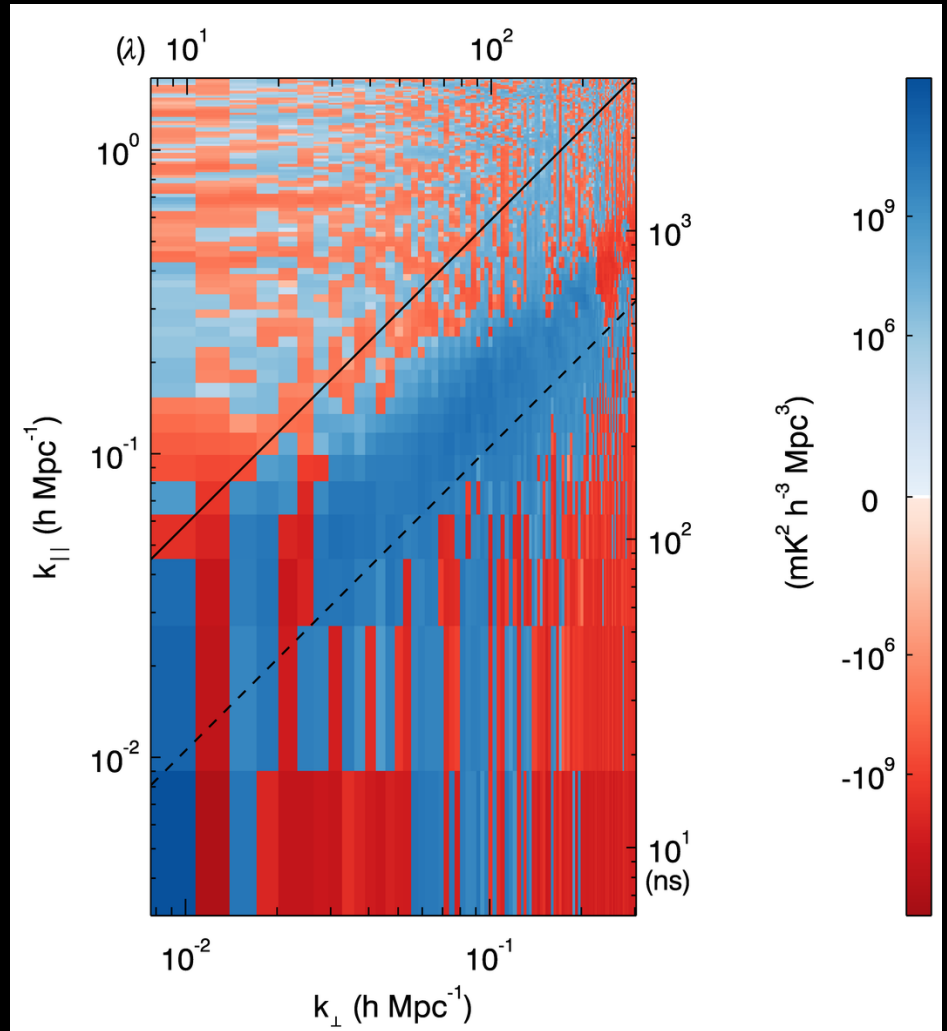
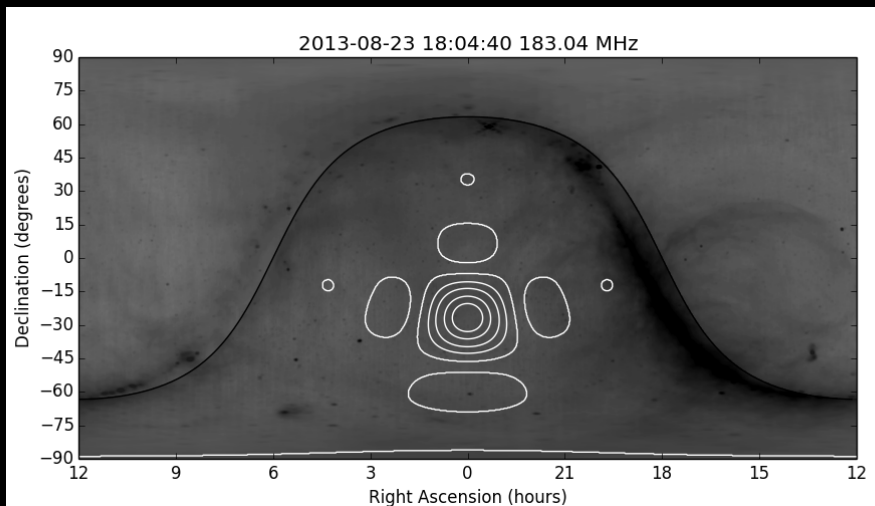




Thyagarajan et al (2015)

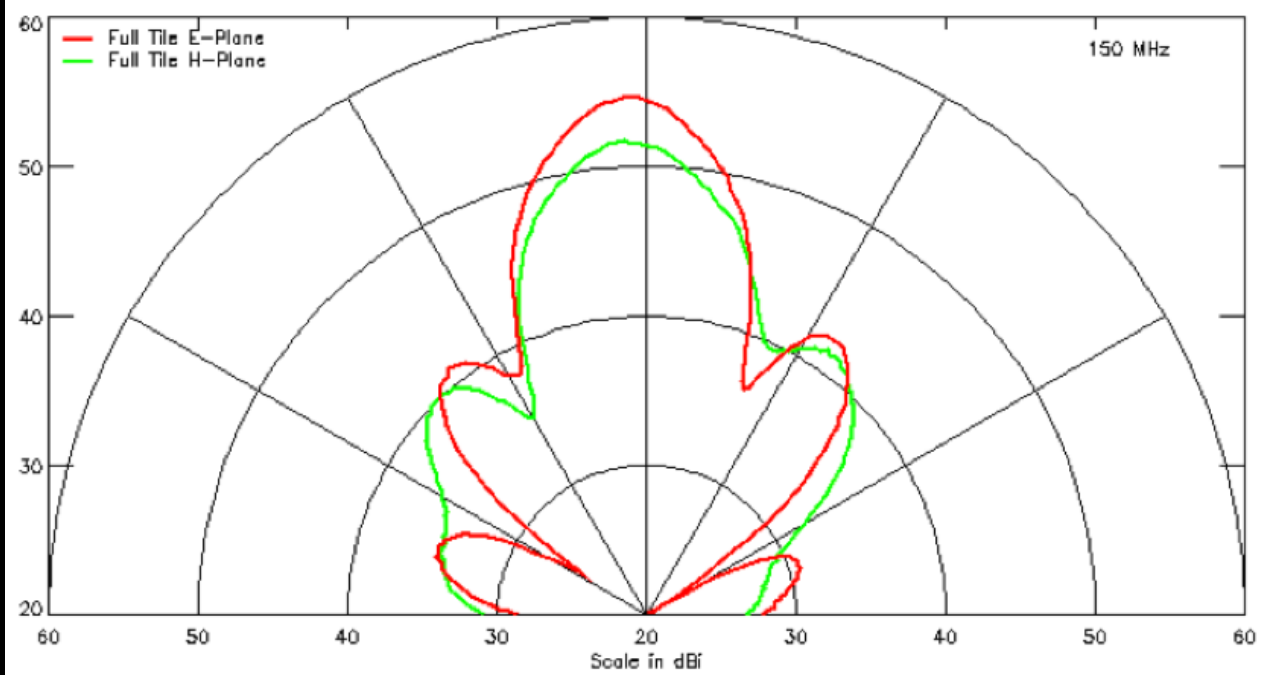
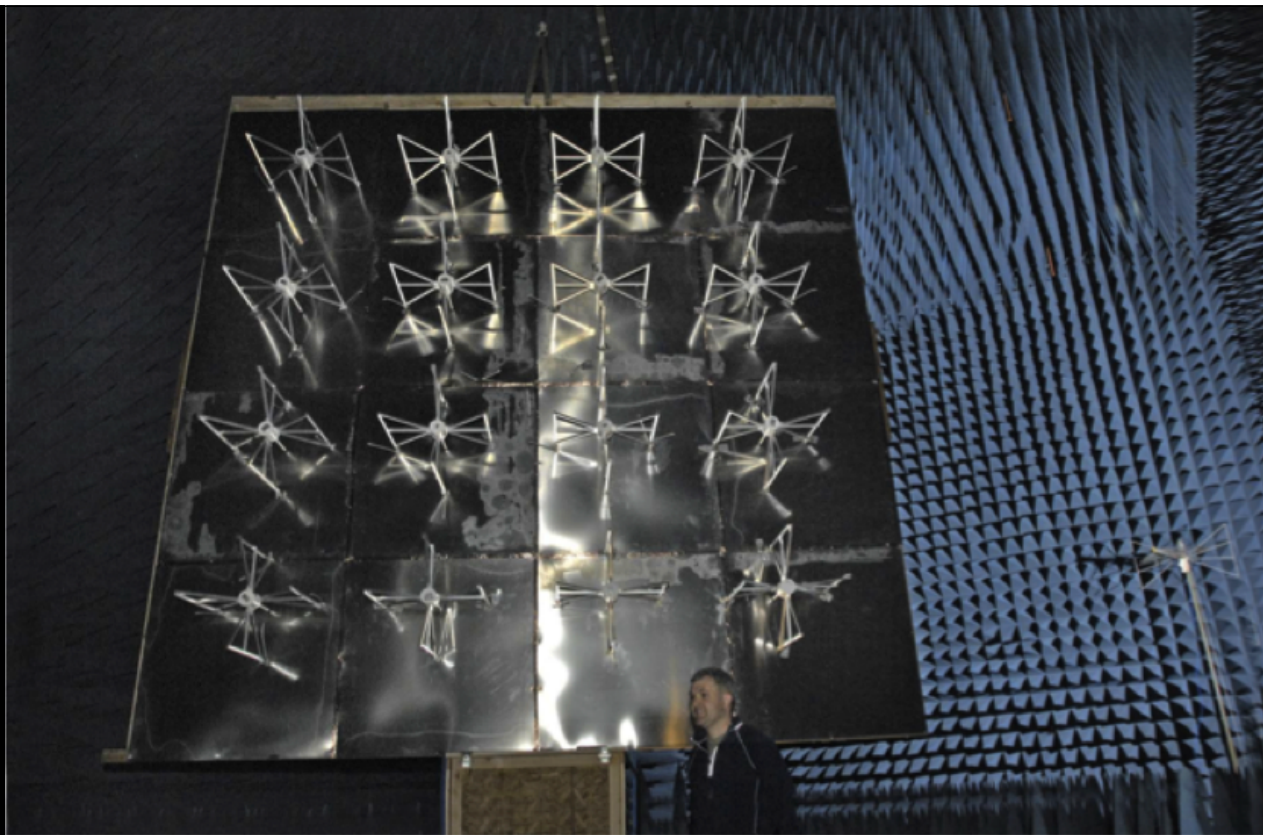
BEAM PRECISION LIMITS SUBTRACTION

- Sidelobe sources bleed higher in k_{\parallel}

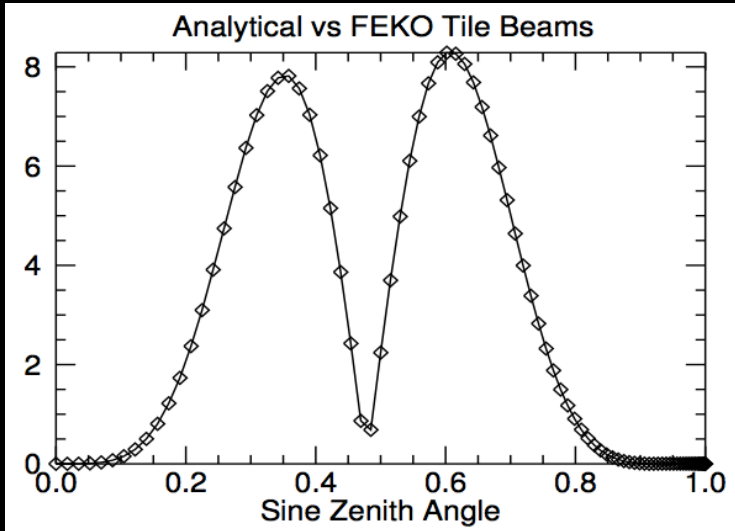


FOR BEAM REQUIREMENTS

- accuracy for model subtraction
preliminary spec 1% at -40dB (pending more modeling)
—> this is likely below the known manufacturing tolerance, implying a requirement to measure the "as-built system"
- similar requirement on precision (ie as built, beam to beam variation) to understand limits of redundancy
- horizon to horizon response (regardless of FoV)



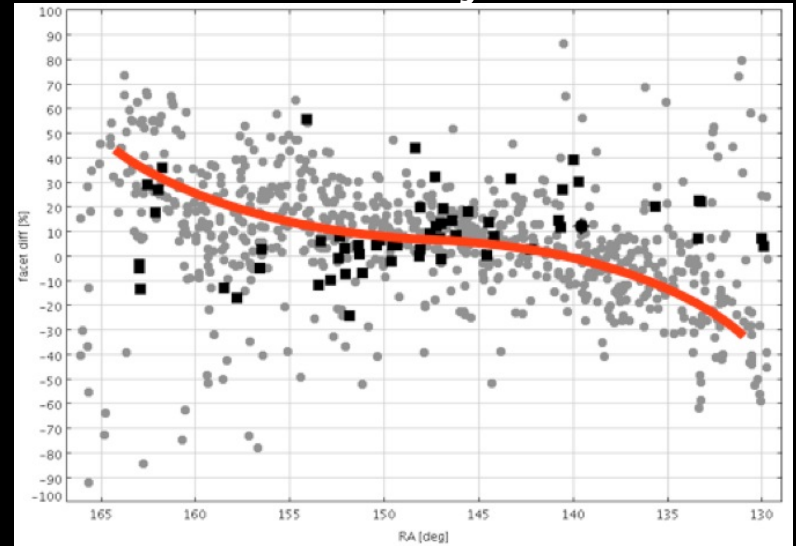
Model Variance



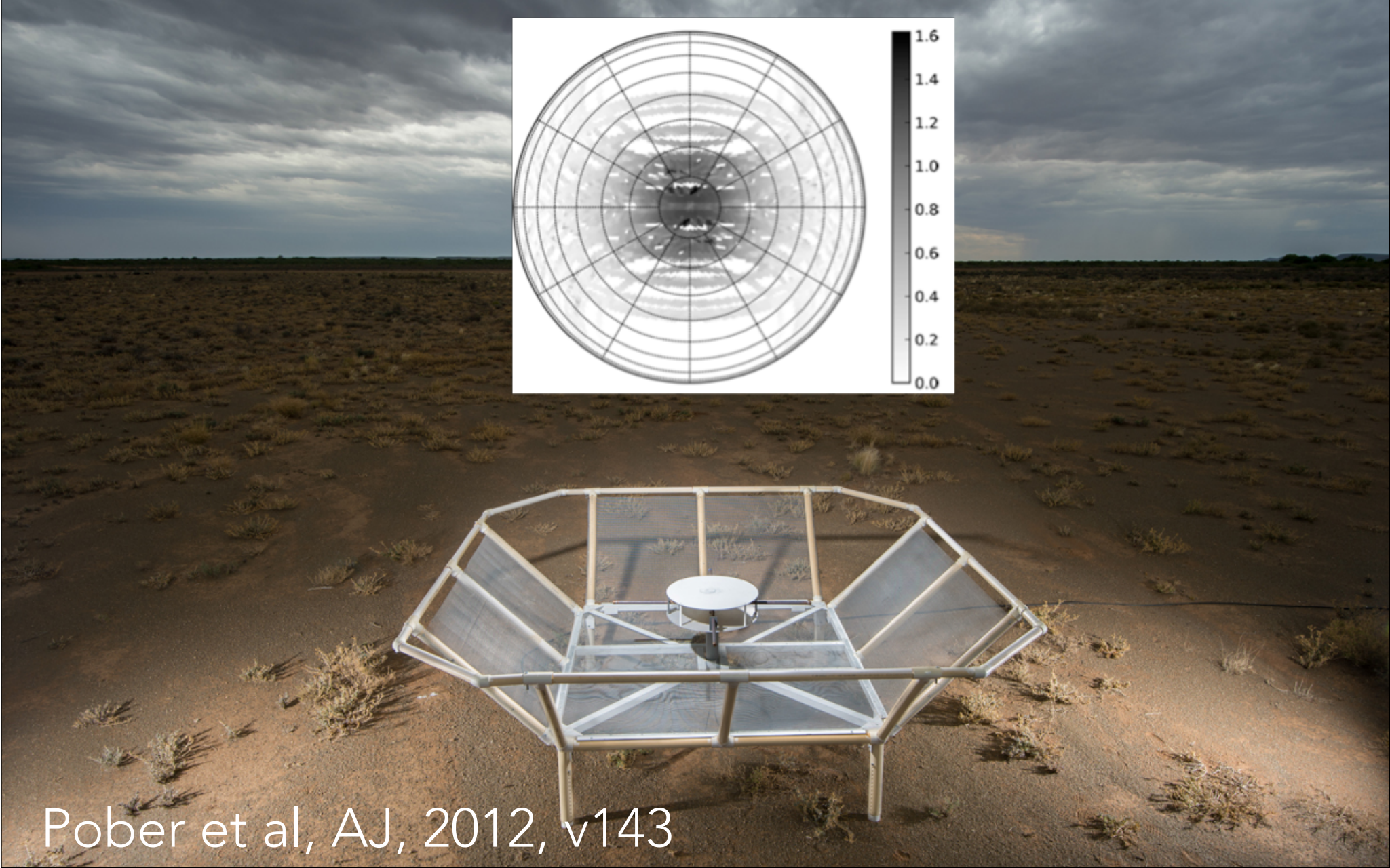
Analysis by Ben McKinley et al

see also Sutinjo et al, Rad Sci, 2015

Flux uncertainty



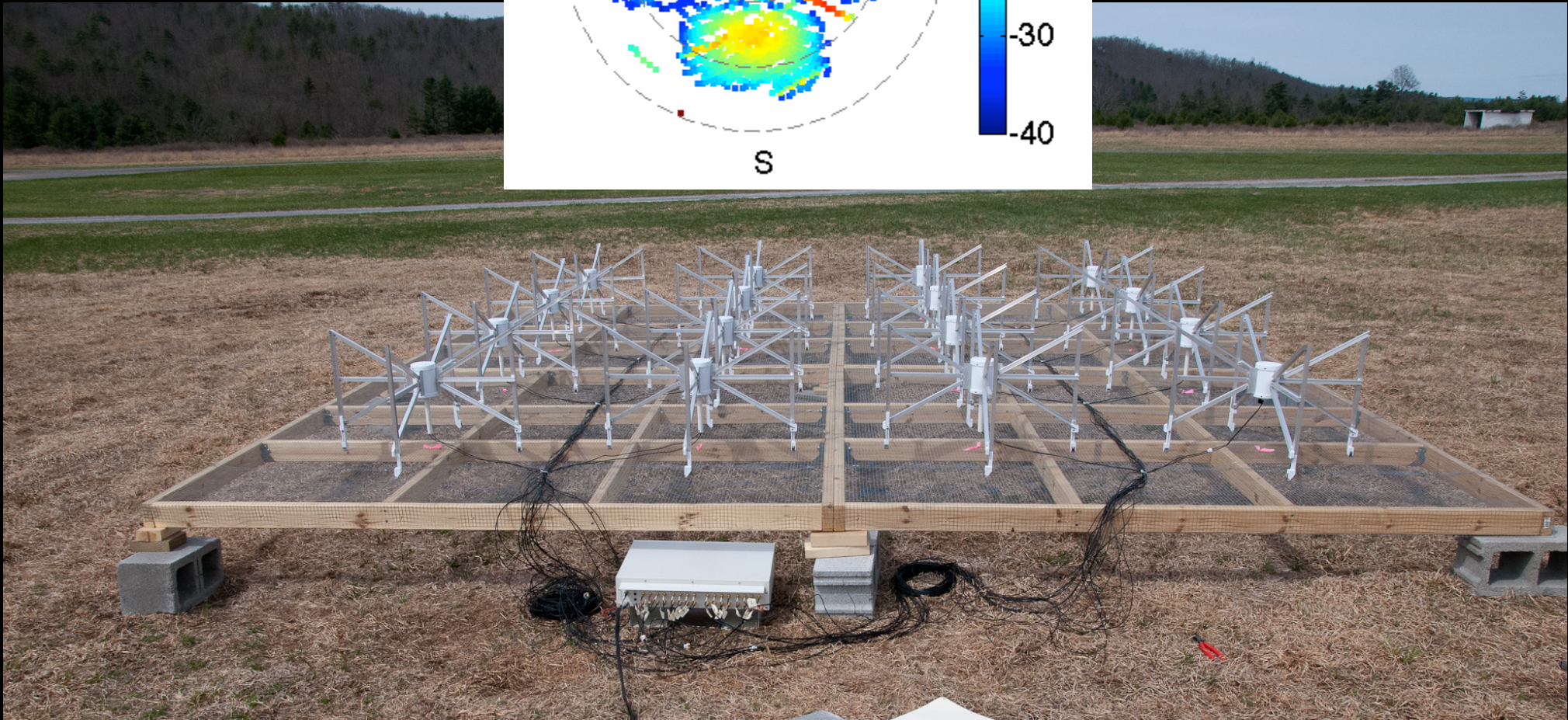
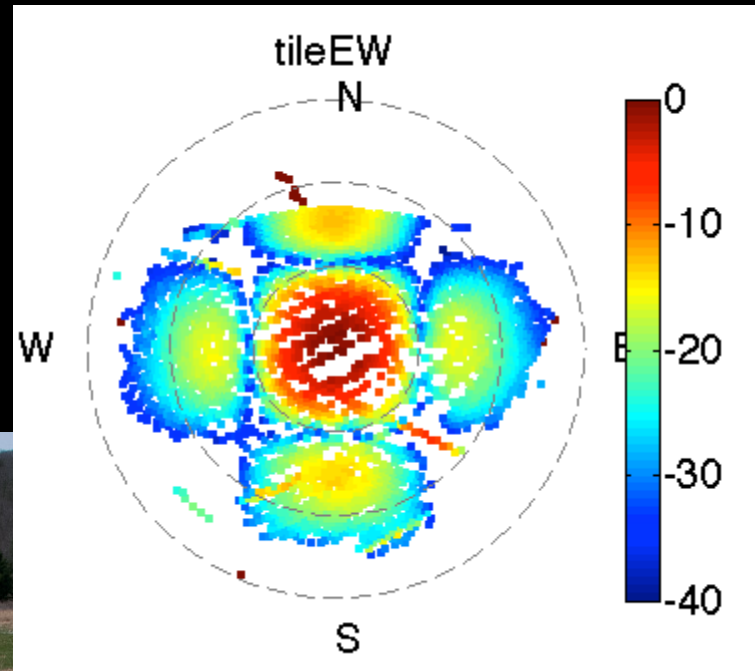
Jacobs et al 2013



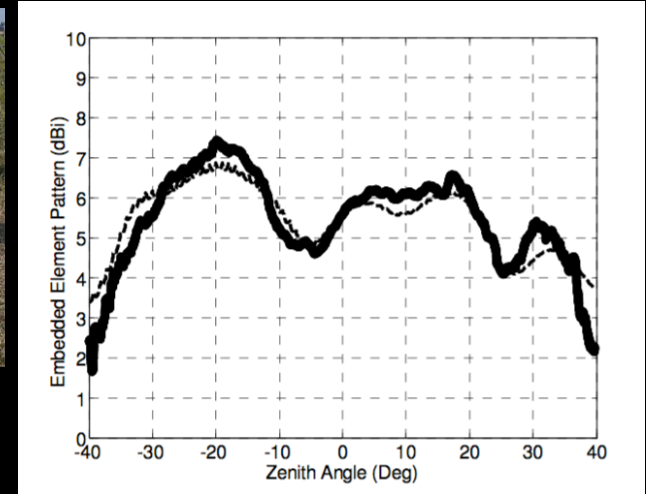
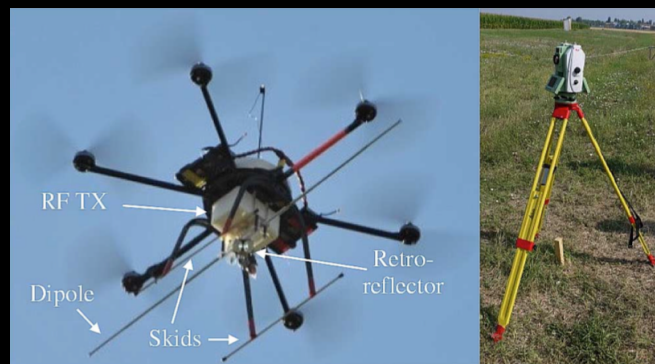
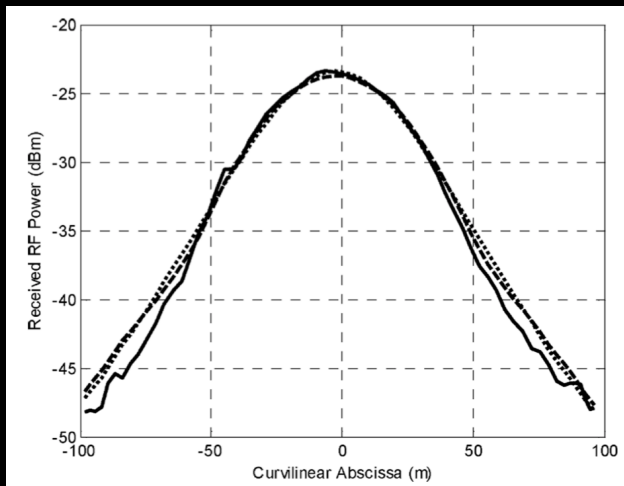
Pober et al, AJ, 2012, v143

Using ORBCOMM

Neben et al Radio Science, 2015, vol 50



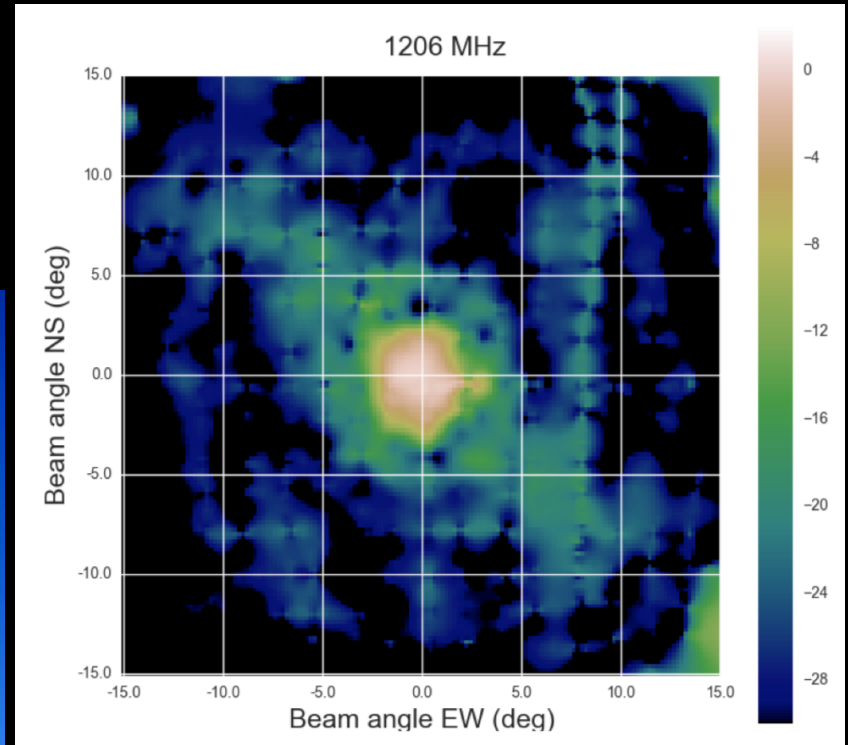
SKA Aperture Array Verification Program



virone et al IEEE AWPL, 2014

virone et al APS IEEE, 2014

Chang et al arxiv:1505.05885
~1GHz



EXTERNAL CALIBRATOR FOR HYDROGEN OBSERVATORIES

Source: VCO synthesizer
(137-2GHz)

Drone: 3DR X8



Antenna: Bicollog bowtie
100-2Ghz







HERA dish

Launch point

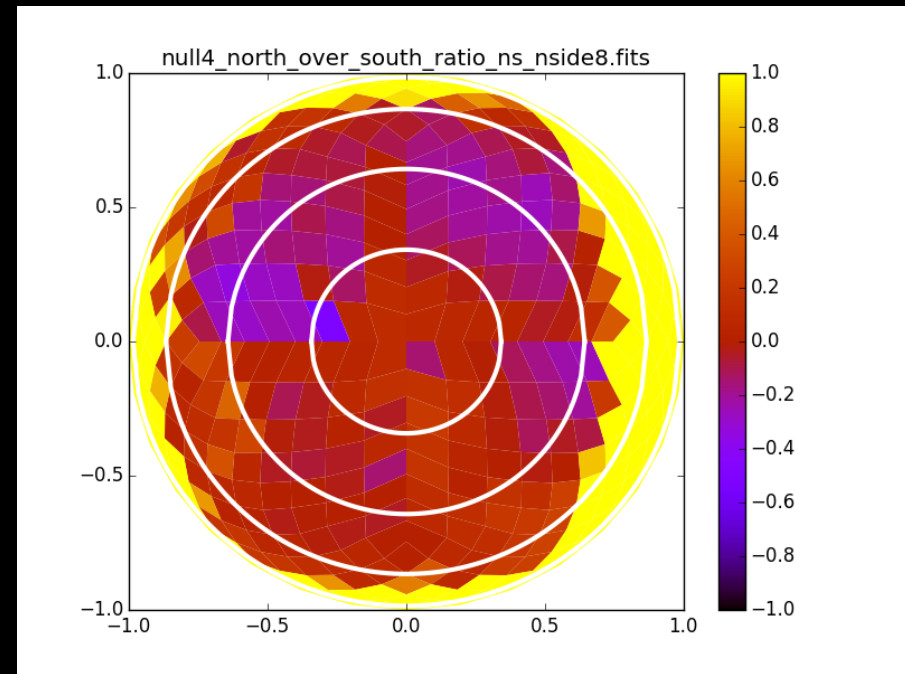
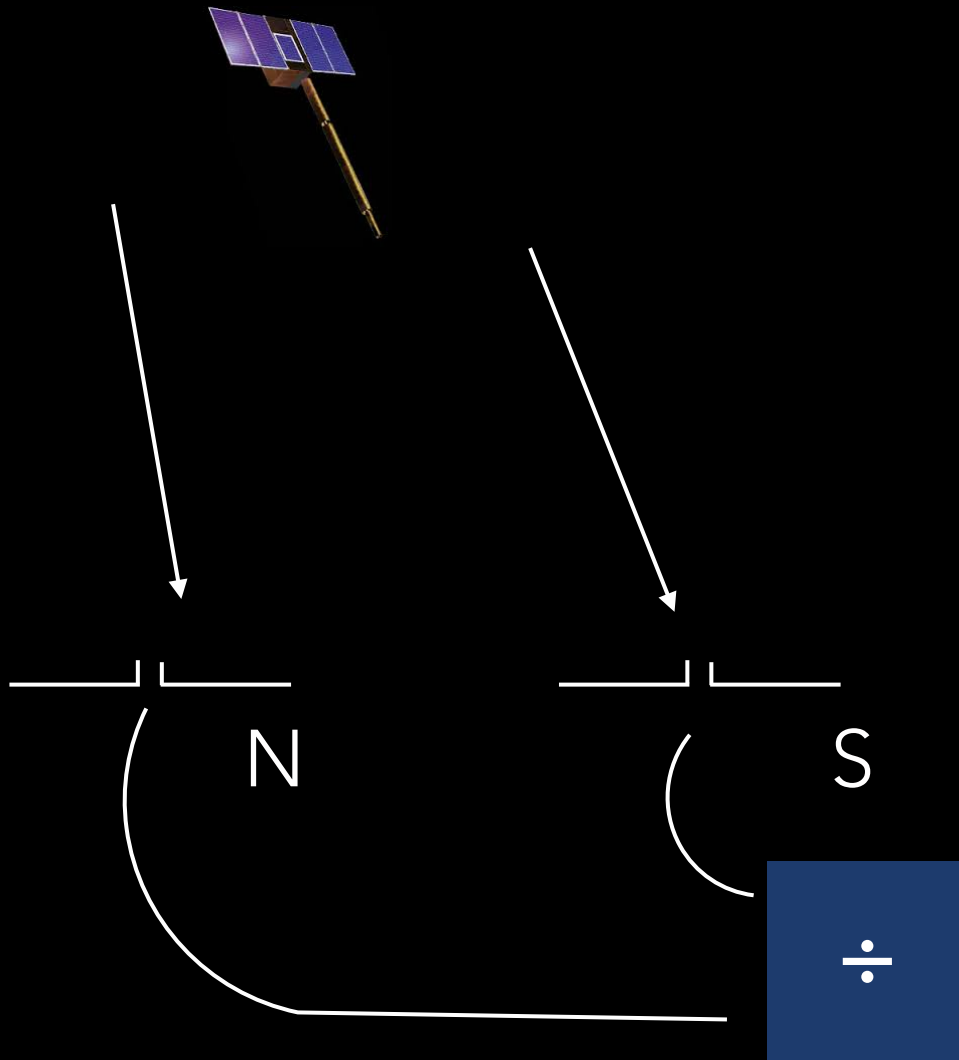
North ORBCOMM
calibration dipoles

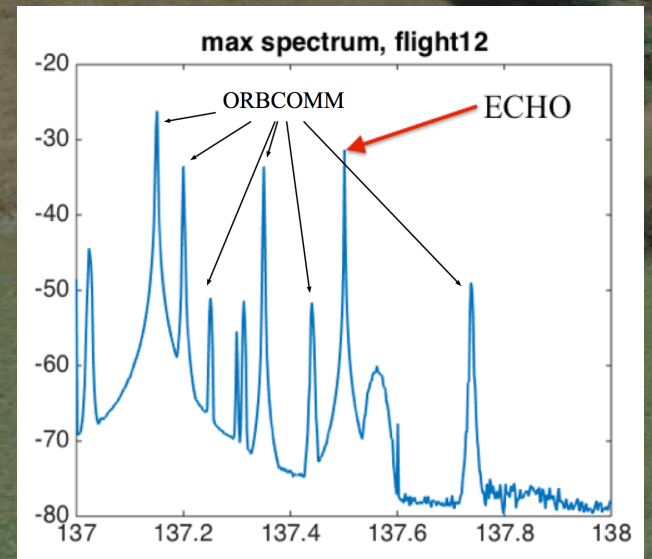
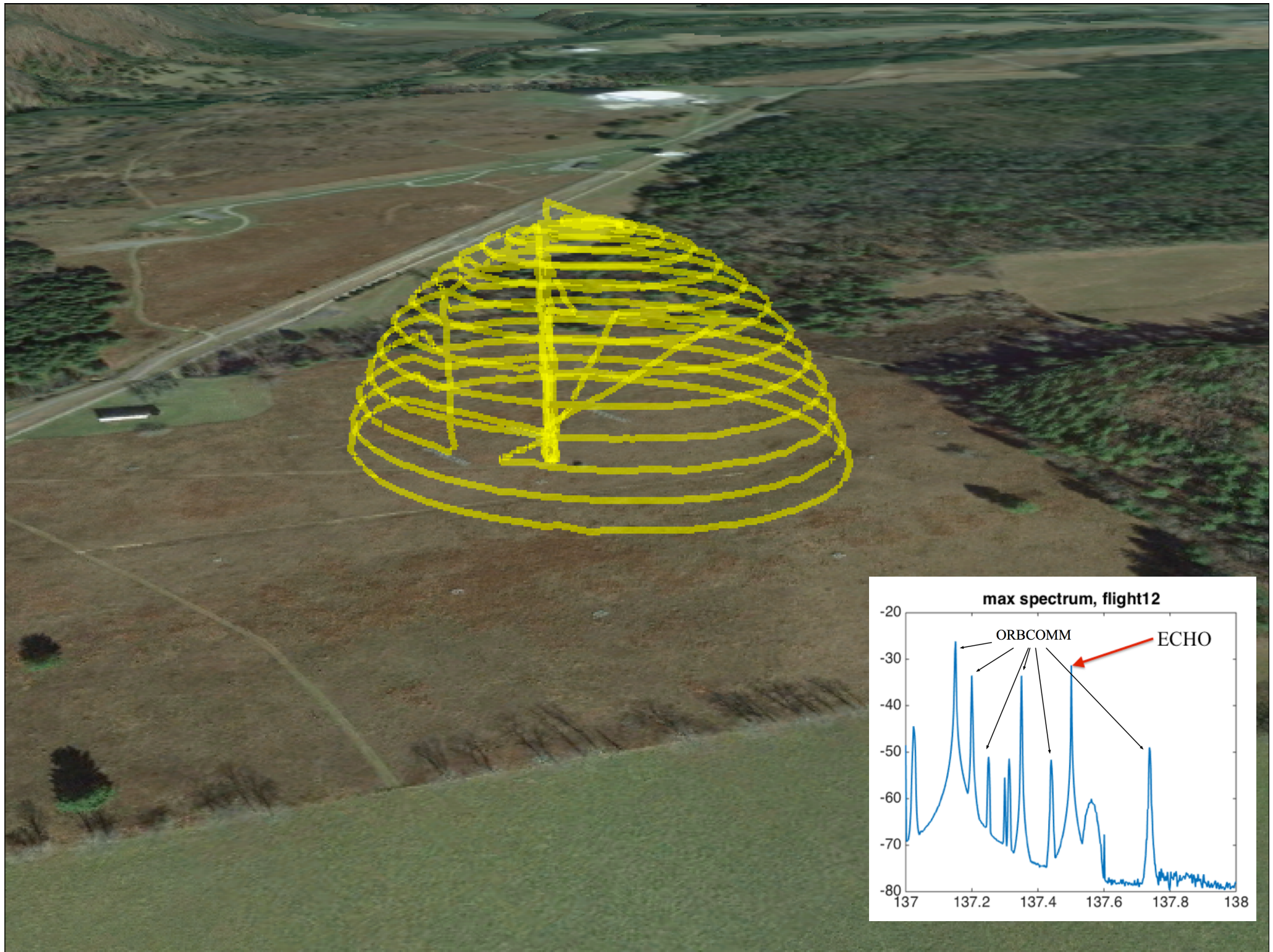
South ORBCOMM
calibration dipoles

50m

N

ORBCOMM NULL CALIBRATION





Looking from above

Healpix ring flight paths



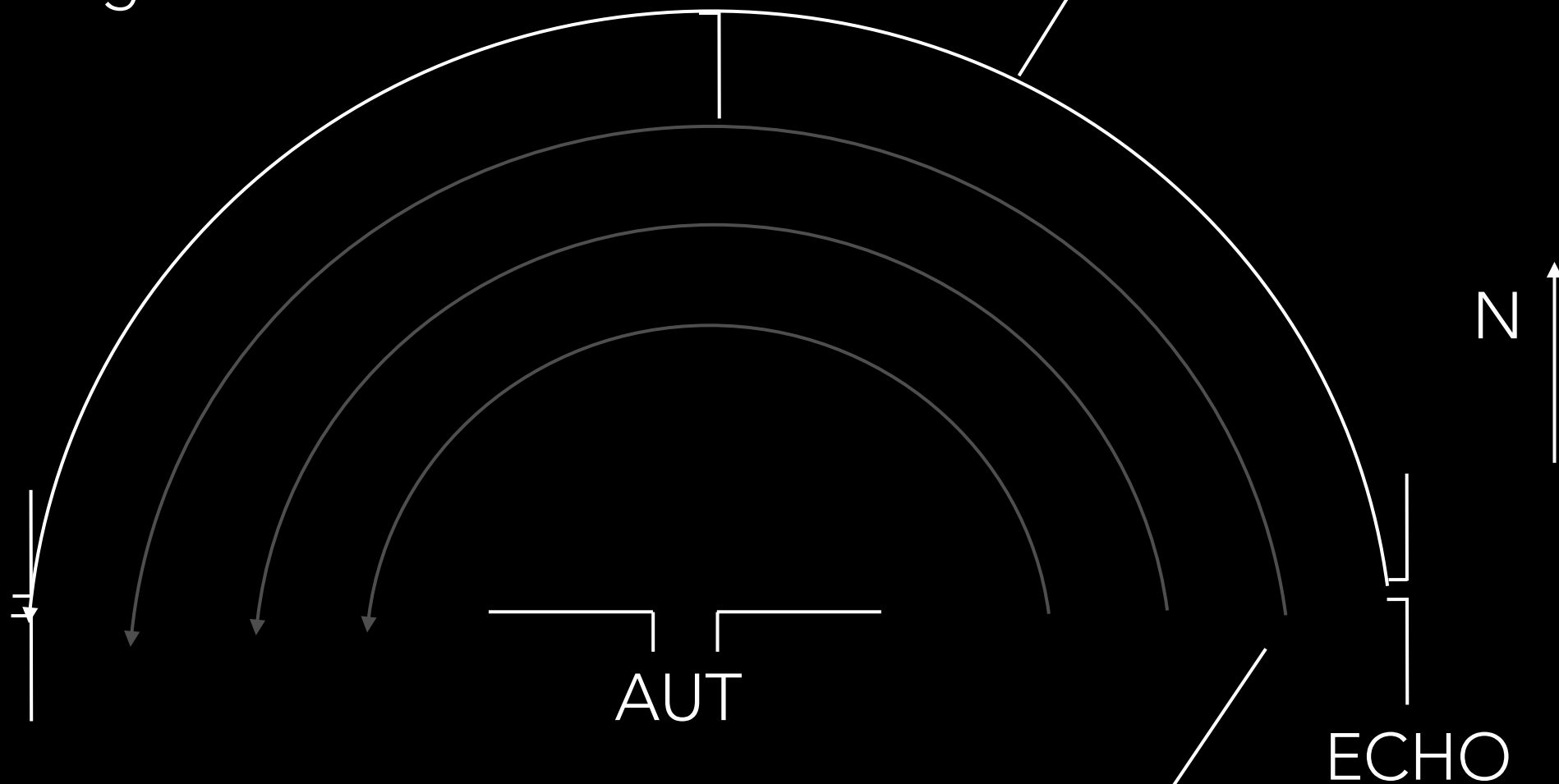
AUT

ECHO

Polarization locked to cardinal directions

Looking from above

Healpix ring flight paths

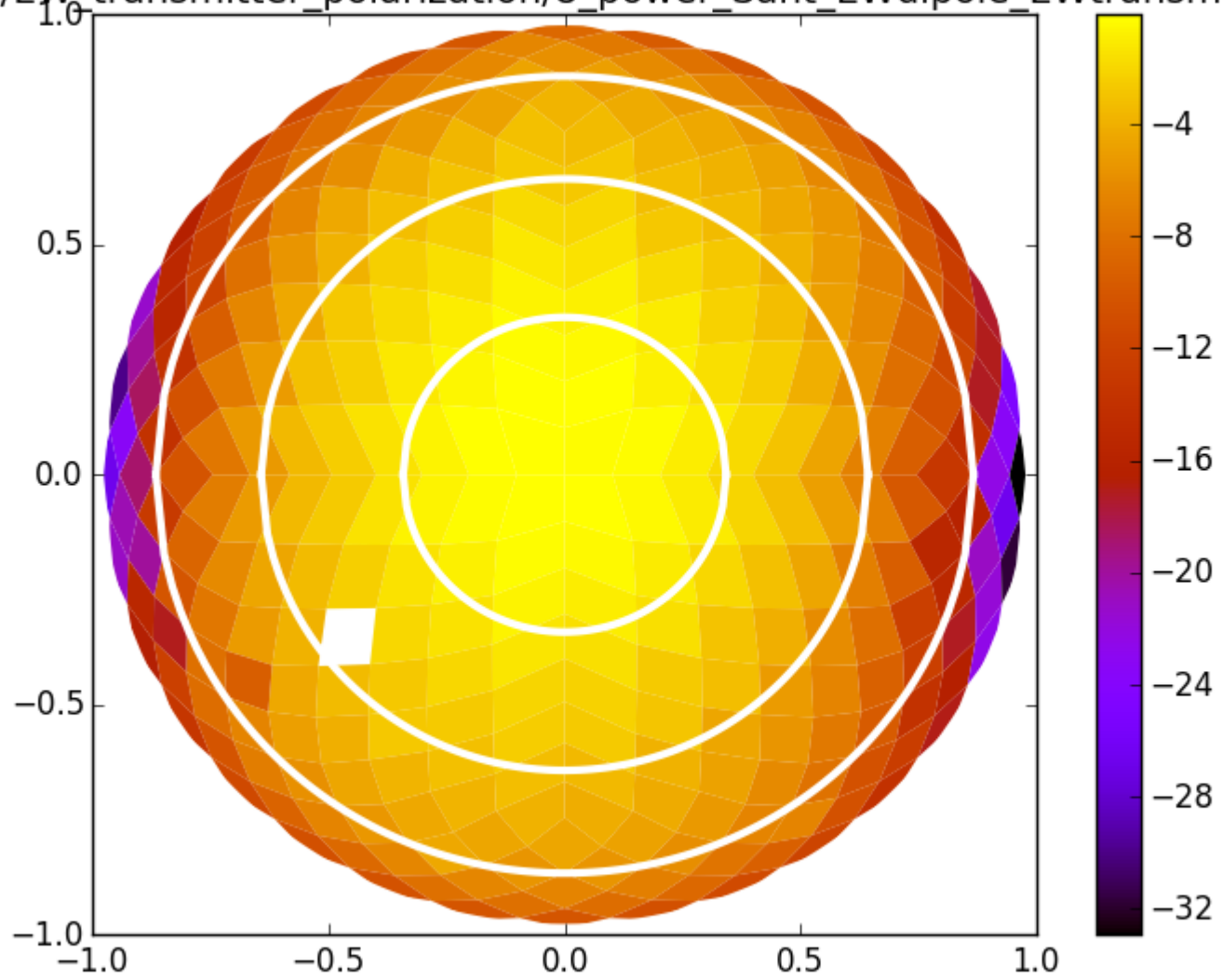


Polarization locked to cardinal directions*

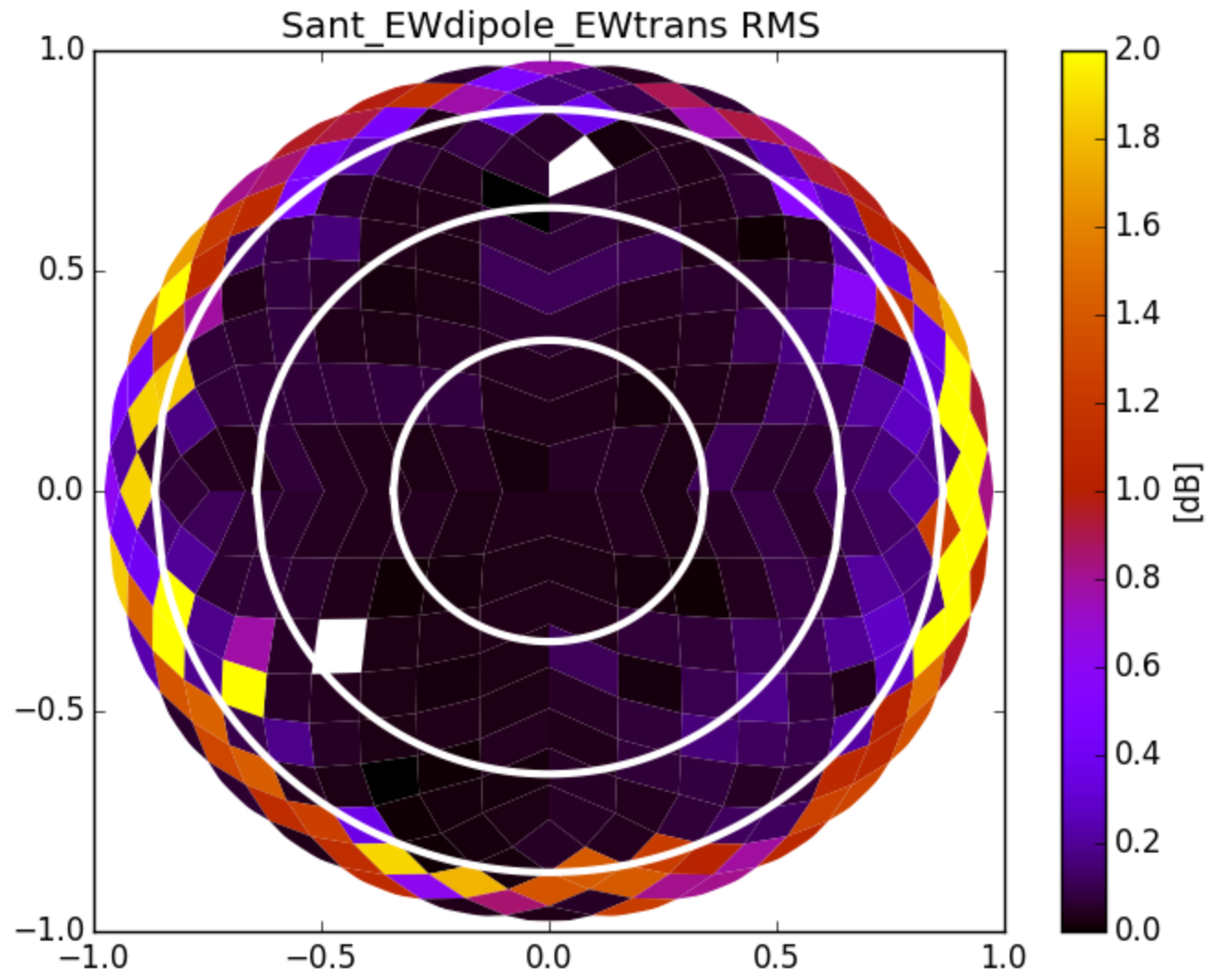
*does not give equal weight to all pols at all sky locations

EXAMPLE MAP

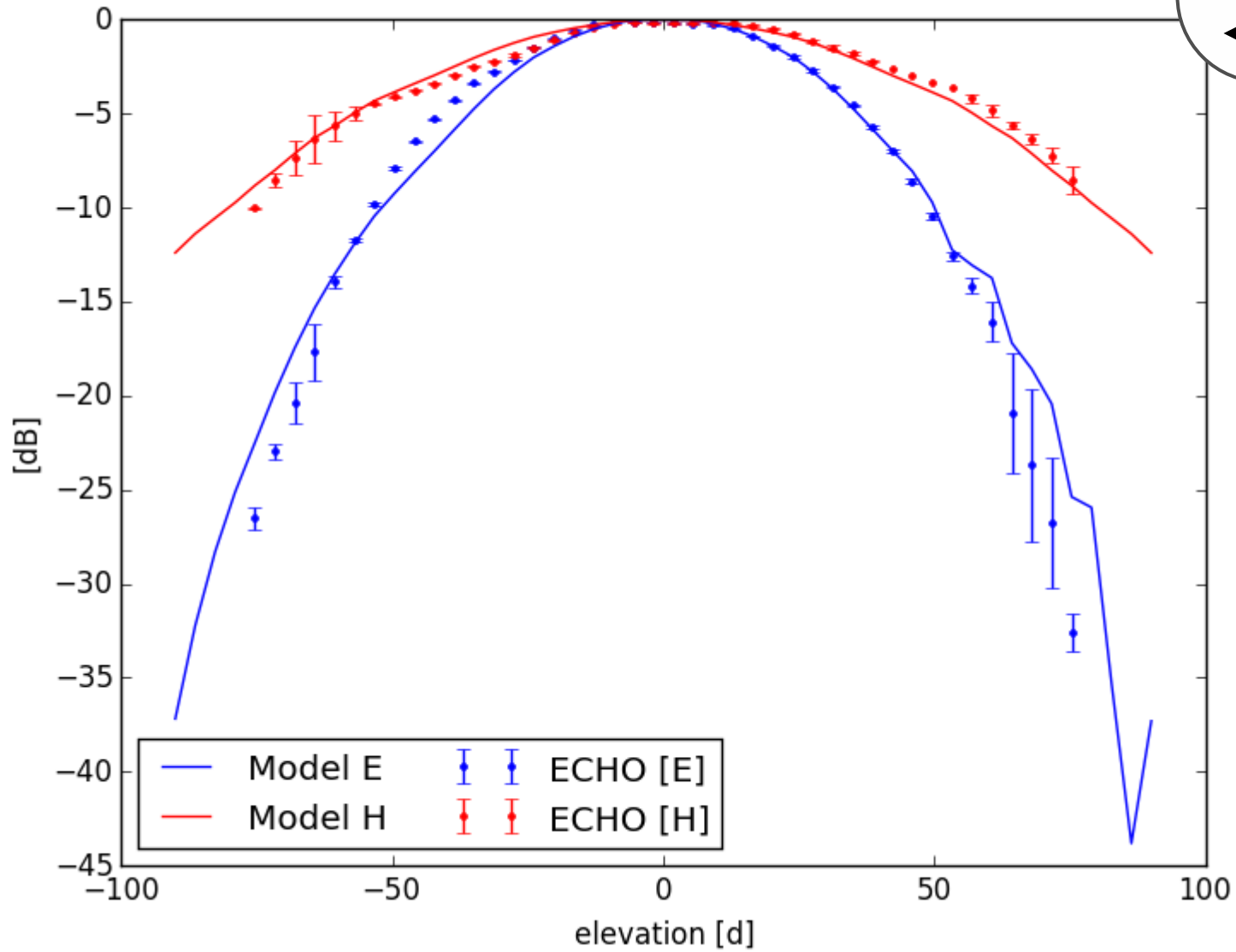
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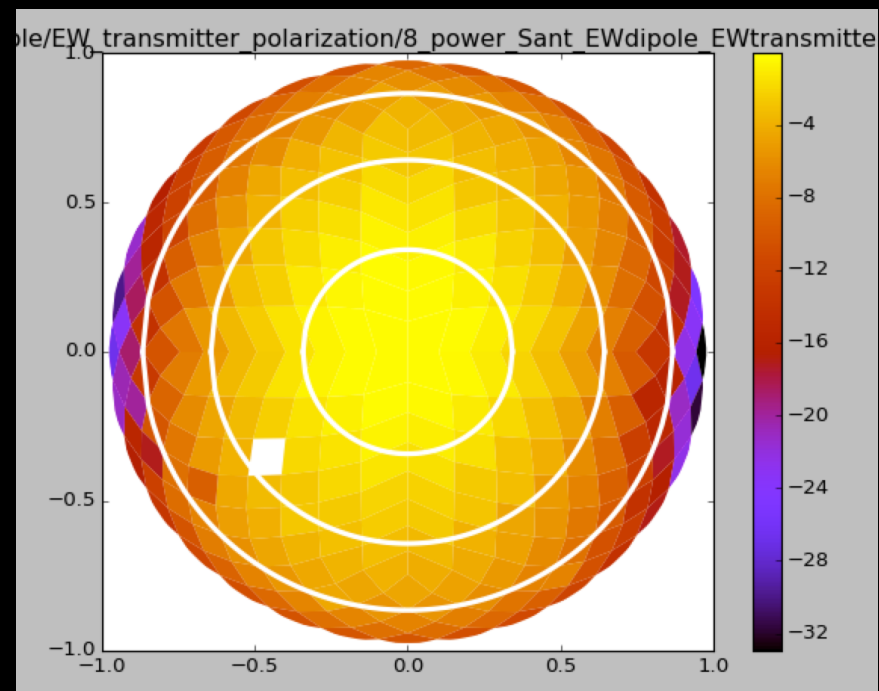
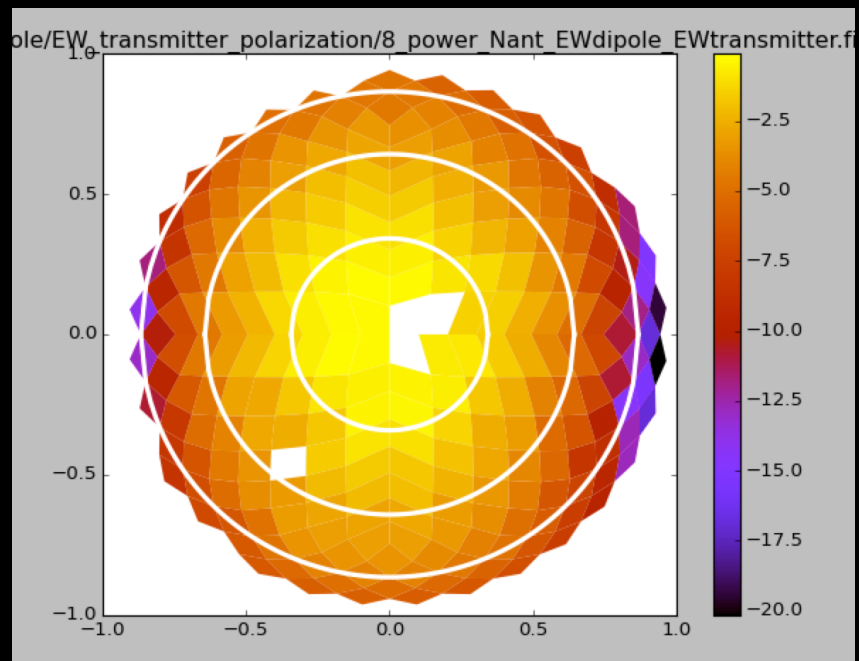


MEASUREMENT PRECISION

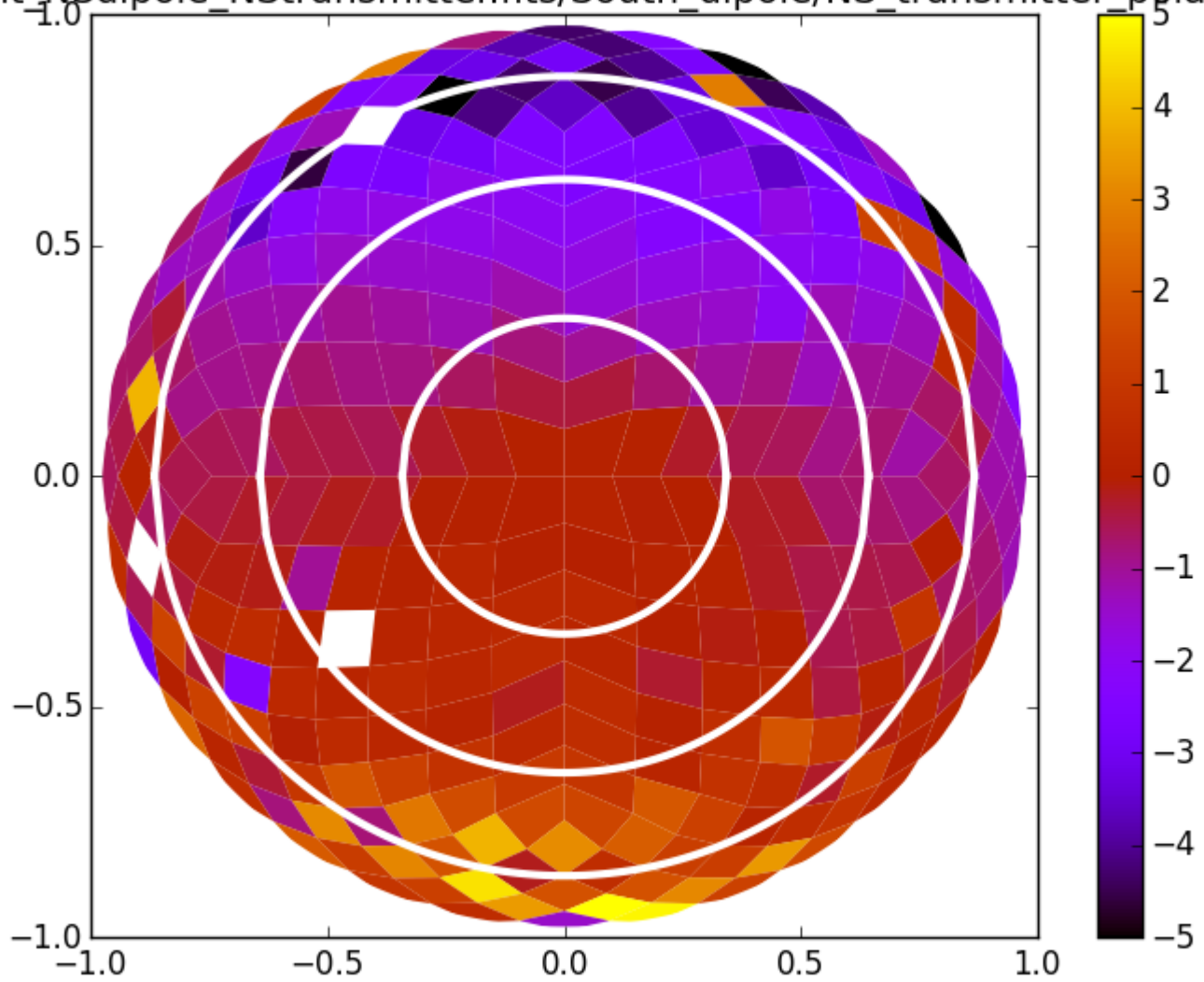


BEAM CUT

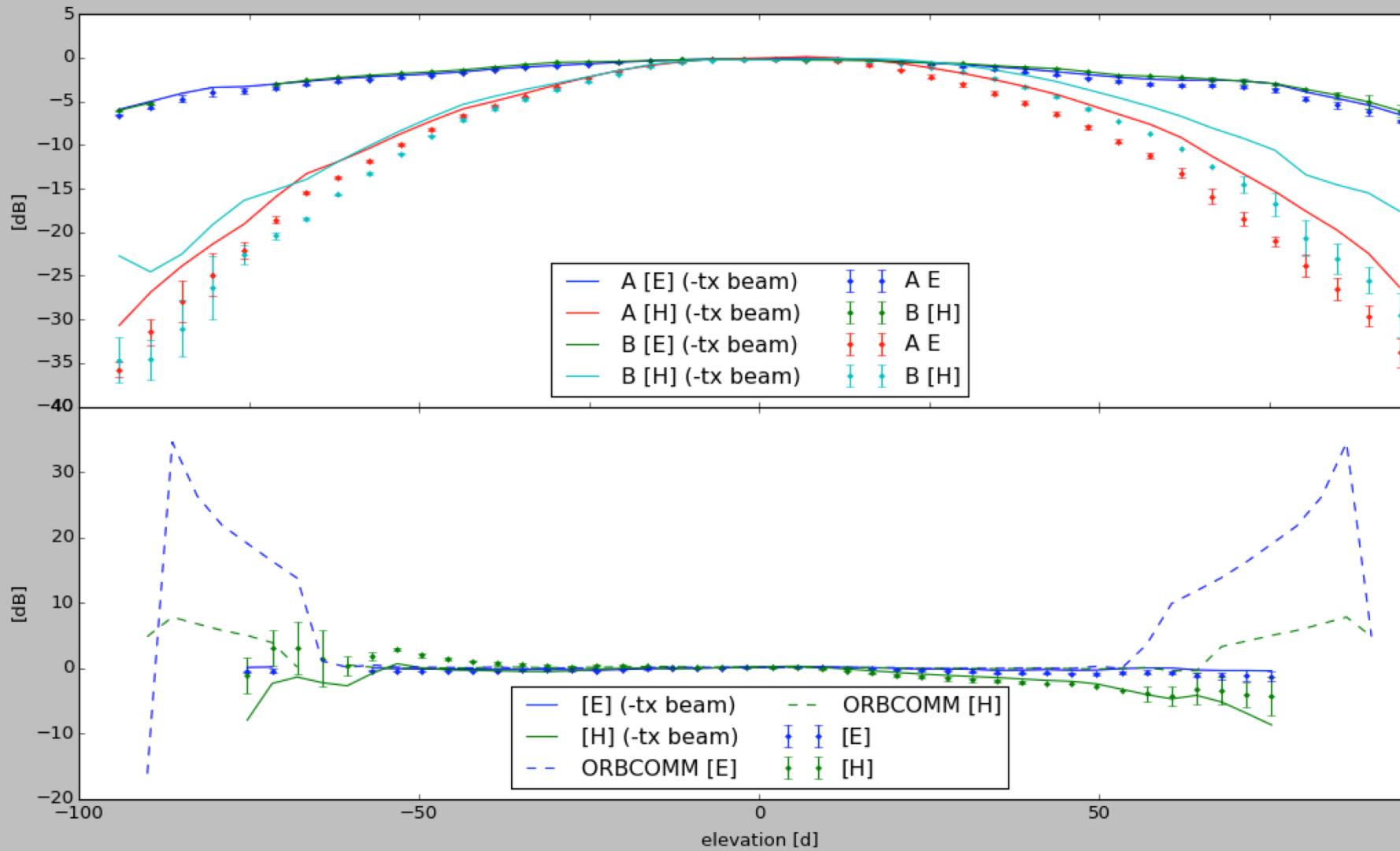




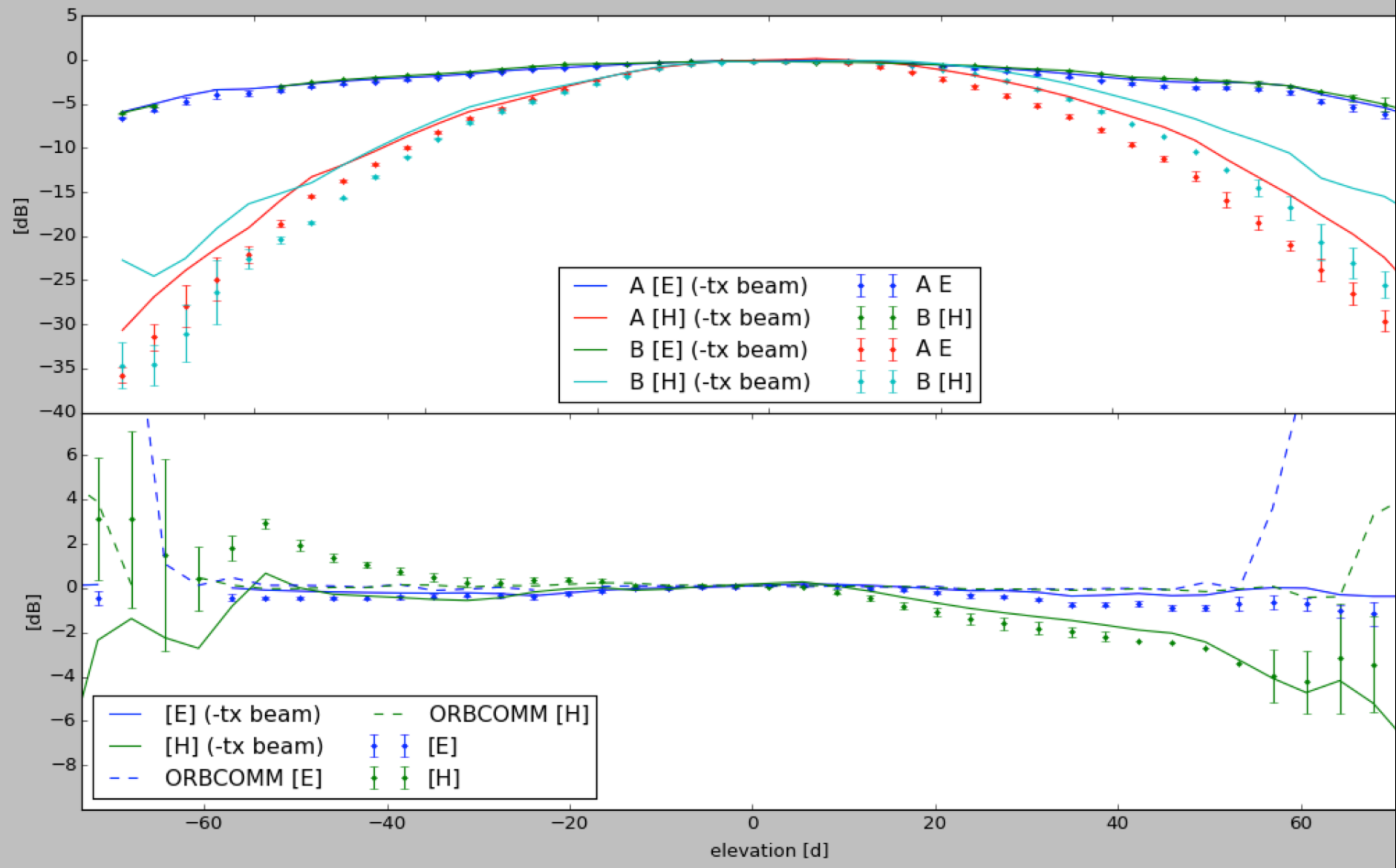
r_Nant_NSdipole_NStransmitter.fits/South_dipole/NS_transmitter_polarization/8_p



North_dipole/NS_transmitter_polarization/8_power_Nant_NSdipole_NStransmitter.fits
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North_dipole/NS_transmitter_polarization/8_power_Nant_NSdipole_NStransmitter.fits
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 --a_tx_correct --b_tx_correct --trans=NS --ratio=null4_north_over_south_ratio_ns.fits --fit_tx



FIRST TEST CONCLUSIONS

An aerial photograph of a radio astronomy site. In the foreground, a large green field contains several small structures and a dirt path. In the middle ground, two large satellite dishes are visible, one on the left and one on the right. The background shows a dense forest of green trees and rolling hills under a cloudy sky.

- Most places much better than 0.1 dB (2%) precision on each map
- Accuracy can be much better than 1 dB (25%) but is subject to systematics
-

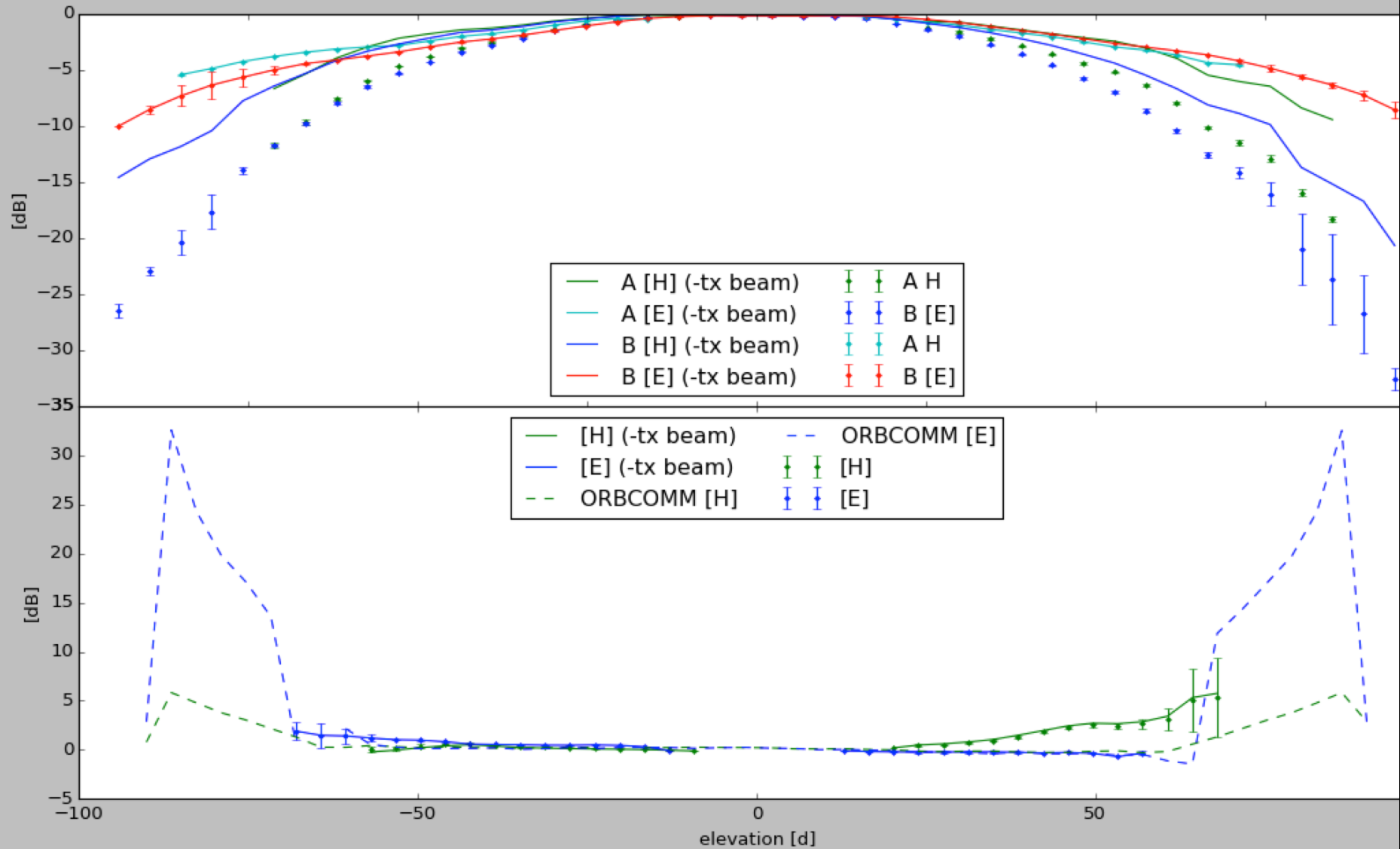
CURRENT\FUTURE WORK

- **Mapping HERA dishes/MWA tiles/[insert your favorite telescope here]**
- Systematic modeling
- real time mapping
- platform stability
- guest instruments
- full polarization coverage
- Site coordination
- EM simulation and anechoic chamber mapping of calibration source
- Identify systematics "as they happen"
- RTK GPS positioning and attitude, longer duration flight time, better EM isolation
- Measure reflections in time domain.
- Add vertical transmitter dipole?
- Goal: routine operation on radio quiet sites

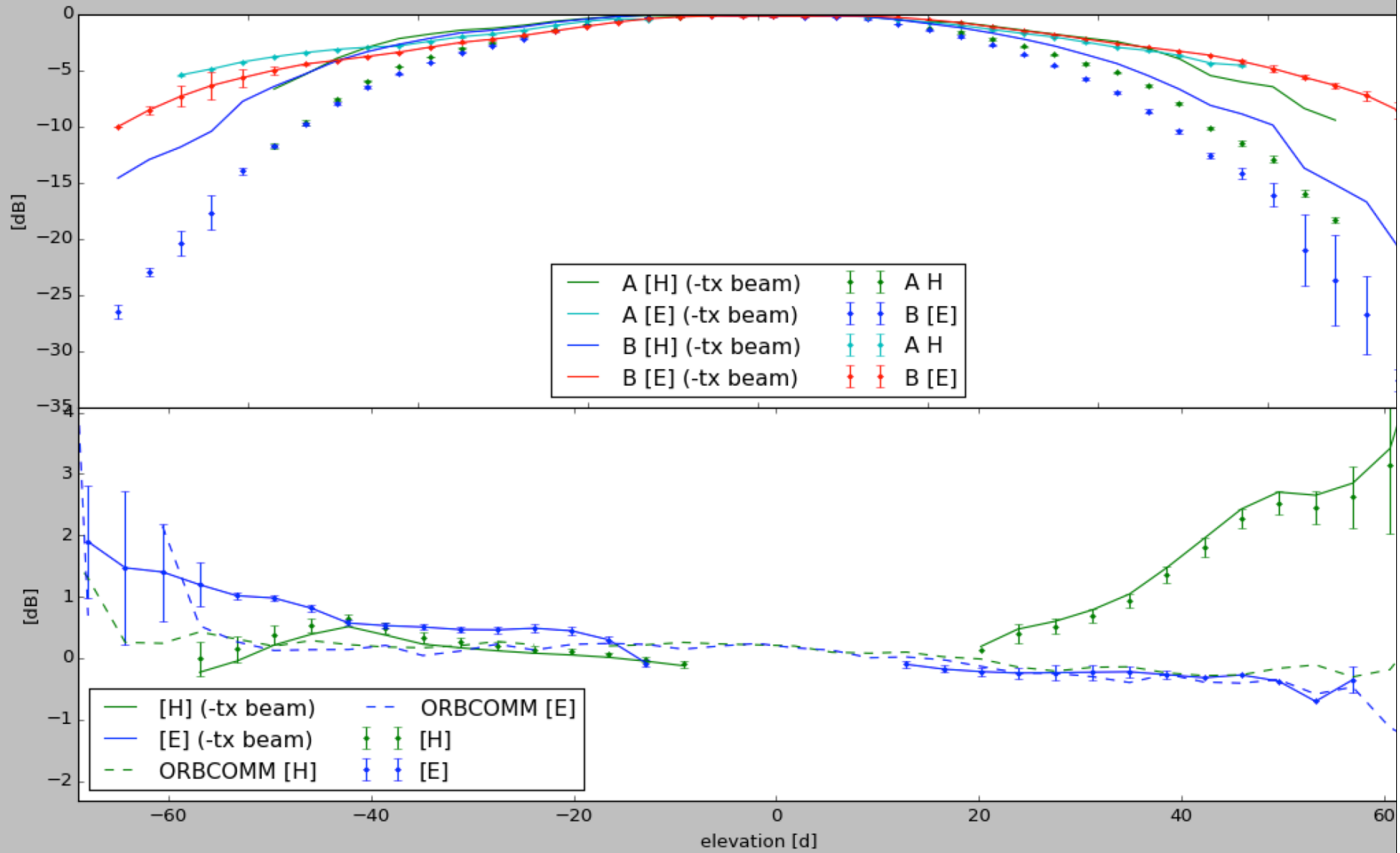


The ECHO Team

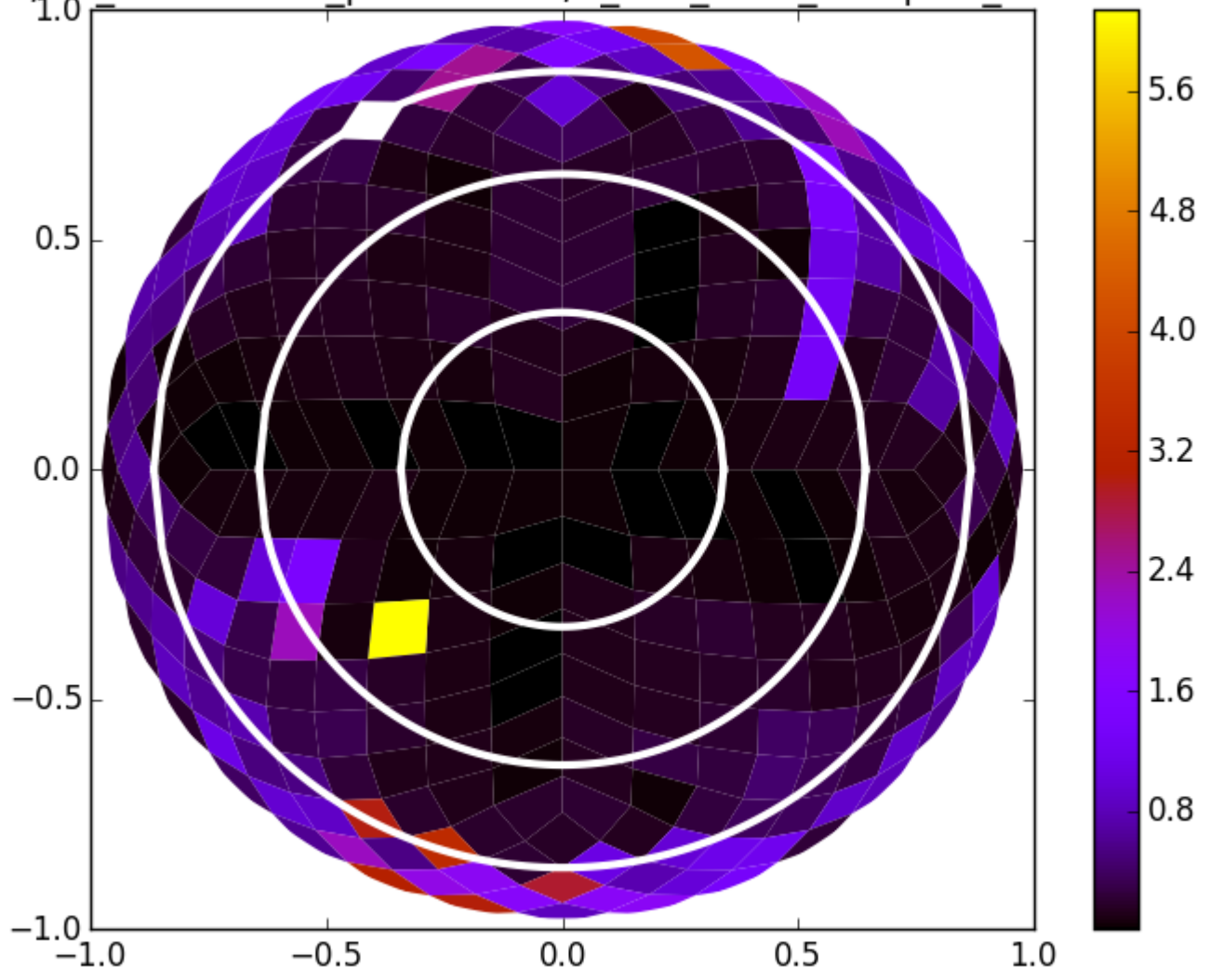
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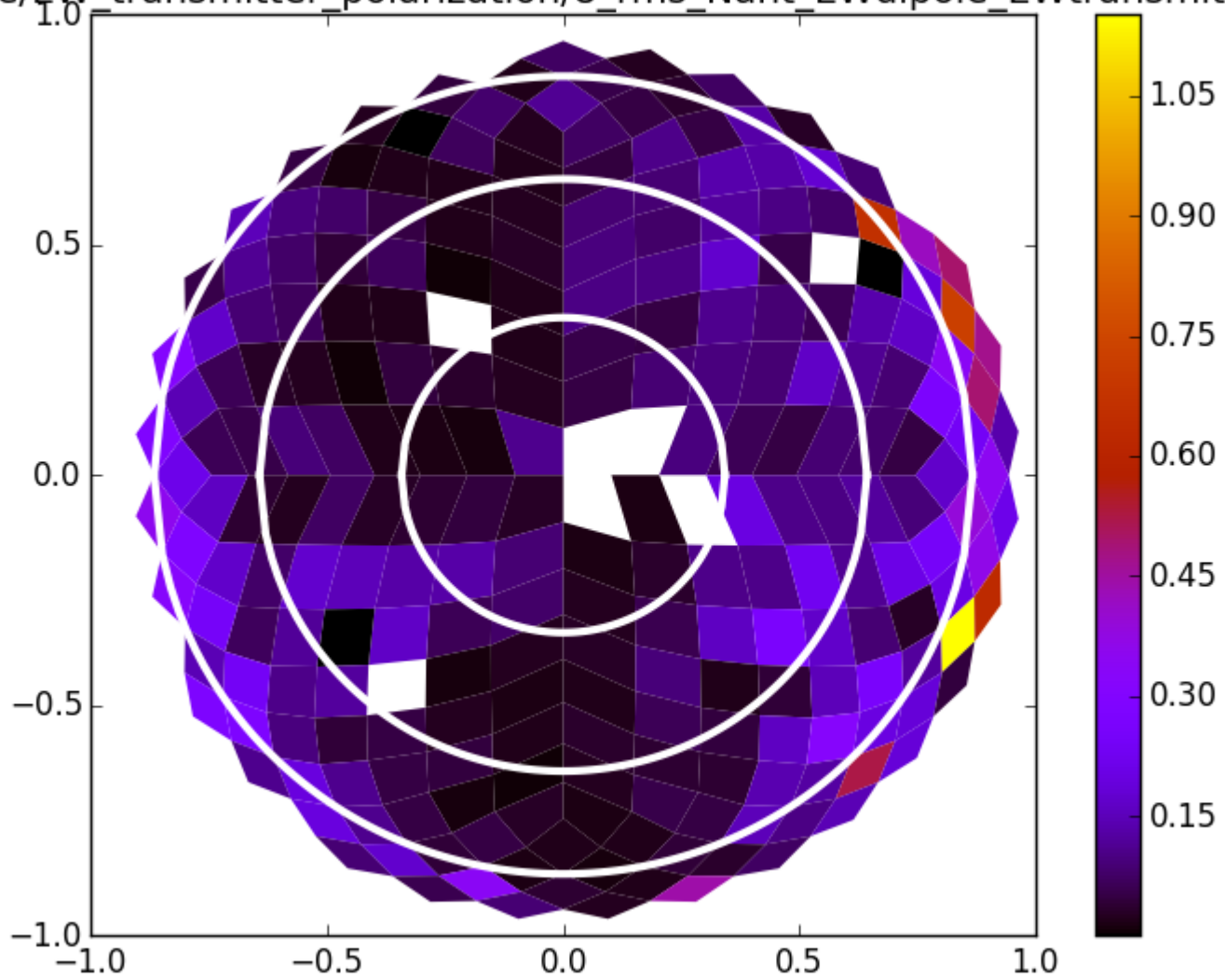
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dipole/NS_transmitter_polarization/8_rms_Nant_NSdipole_NStransmitter.fits



dipole/EW_transmitter_polarization/8_rms_Nant_EWdipole_EWtransmitter.fits



Nant_EWdipole_EWtransmitter.fits/South_dipole/EW_transmitter_polarization/8_p

