



THE LONG WAVELENGTH ARRAY

Jayce Dowell (UNM)
on behalf of the LWA Collaboration
December 2, 2015



LWA STATIONS



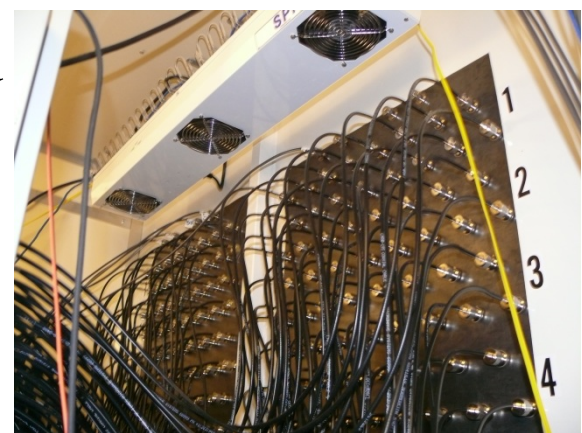
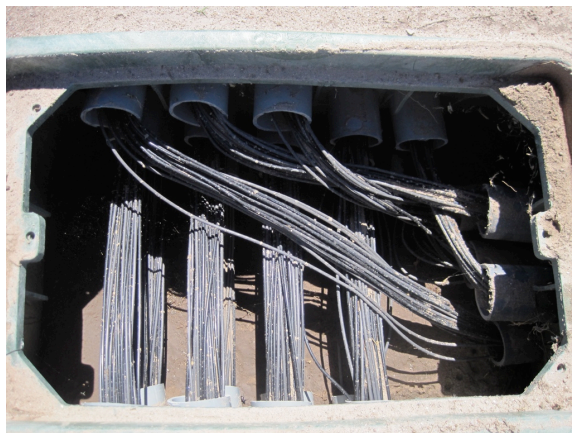
LWA by the Numbers



- 4 years operating the first LWA station
- 10 to 88 MHz frequency range
- 4 independent beams with 2 x 16 MHz tunings
- Triggered observing mode – on sky in < 2 minutes
- 81 active projects on LWA1
 - 117 active users for 42 institutions
 - 23 refereed publications on LWA1 results
- 1 new station being commissioned in Owens Valley
 - See Gregg Hallinan's talk
- 1 new station under construction at Sevilleta, NM



LWA1



LWA-SV

- 3-88 MHz frequency range
- All 256 antennas + outrigger deployed
- Advanced Digital Processor being commissioned



LWA Radio Observatory Staff - UNM



Faculty and Staff

Greg Taylor	Trish Henning
Jayce Dowell	John Dickel
Lanie Dickel	HAL

Students

Karishma Bansal
Joe Malins

Veronica Dike
Jessica Lopez

Postdocs

Frank Schinzel

Kevin Stovall



HAL

HAL has successfully created the following schedule for UTC 2014/09/04:

Schedule:

```
* 2014/09/04 00:04:00 /home/op1/MCS/exec/acquireTBWAndProcess.py
* 2014/09/04 01:30:00 /home/op1/LO001/runLO001_split.py 14280
* 2014/09/04 05:39:00 /home/op1/MCS/sch/INIIdp.sh
* 2014/09/04 05:54:00 /home/op1/MCS/exec/setASP_LS003.sh
* 2014/09/04 05:59:03 LS003001, session 1806 starts on beam 2
* 2014/09/04 05:59:03 LS003001, session 1807 starts on beam 1
* 2014/09/04 05:59:03 LS003001, session 1808 starts on beam 4
* 2014/09/04 05:59:03 LS003001, session 1809 starts on beam 3
* 2014/09/04 06:59:13 LS003001, session 1806 stops on beam 2
* 2014/09/04 06:59:13 LS003001, session 1807 stops on beam 1
* 2014/09/04 06:59:13 LS003001, session 1808 stops on beam 4
* 2014/09/04 06:59:13 LS003001, session 1809 stops on beam 3
* 2014/09/04 07:01:00 /home/op1/MCS/sch/INIIdp.sh
* 2014/09/04 07:02:00 /home/op1/MCS/exec/setLEDA64_split.sh
* 2014/09/04 07:20:00 /home/op1/MCS/sch/startTBN_split38.sh
* 2014/09/04 07:24:00 /home/op1/MCS/exec/acquireTBWAndProcess.py
* 2014/09/04 08:31:00 /home/op1/MCS/sch/INIIdp.sh
* 2014/09/04 08:51:55 LS006, session 112 starts on beam 2
* 2014/09/04 08:51:55 LS006, session 113 starts on beam 4
* 2014/09/04 09:52:05 LS006, session 112 stops on beam 2
* 2014/09/04 09:52:05 LS006, session 113 stops on beam 4
* 2014/09/04 09:54:00 /home/op1/MCS/sch/INIIdp.sh
* 2014/09/04 10:59:55 LH011, session 121 starts on TBN/TBW
* 2014/09/04 12:00:05 LH011, session 121 stops on TBN/TBW
* 2014/09/04 12:04:00 /home/op1/MCS/sch/startTBN_split38.sh
* 2014/09/04 12:08:00 /home/op1/MCS/exec/acquireTBWAndProcess.py
* 2014/09/04 12:52:00 /home/op1/MCS/sch/startTBN_split38.sh
* 2014/09/04 16:11:00 /home/op1/MCS/exec/acquireTBWAndProcess.py
* 2014/09/04 20:26:00 /home/op1/MCS/exec/acquireTBWAndProcess.py
* 2014/09/04 20:52:00 /home/op1/MCS/sch/INIIdp.sh
* 2014/09/04 21:12:30 LS006, session 114 starts on beam 2
* 2014/09/04 21:12:30 LS006, session 115 starts on beam 4
* 2014/09/04 22:12:40 LS006, session 114 stops on beam 2
* 2014/09/04 22:12:40 LS006, session 115 stops on beam 4
* 2014/09/04 22:14:00 /home/op1/MCS/sch/INIIdp.sh
* 2014/09/04 22:34:00 /home/op1/MCS/sch/startTBN_split38.sh
```

Fully autonomous operation of LWA1

- Generation of static schedule
- Dynamic scheduling of filler observations
- Accepts automatic triggering of observations
- Reacts on environmental issues:
 - Electronics overheating
 - Lightning protection mode
- Automatic station health checks

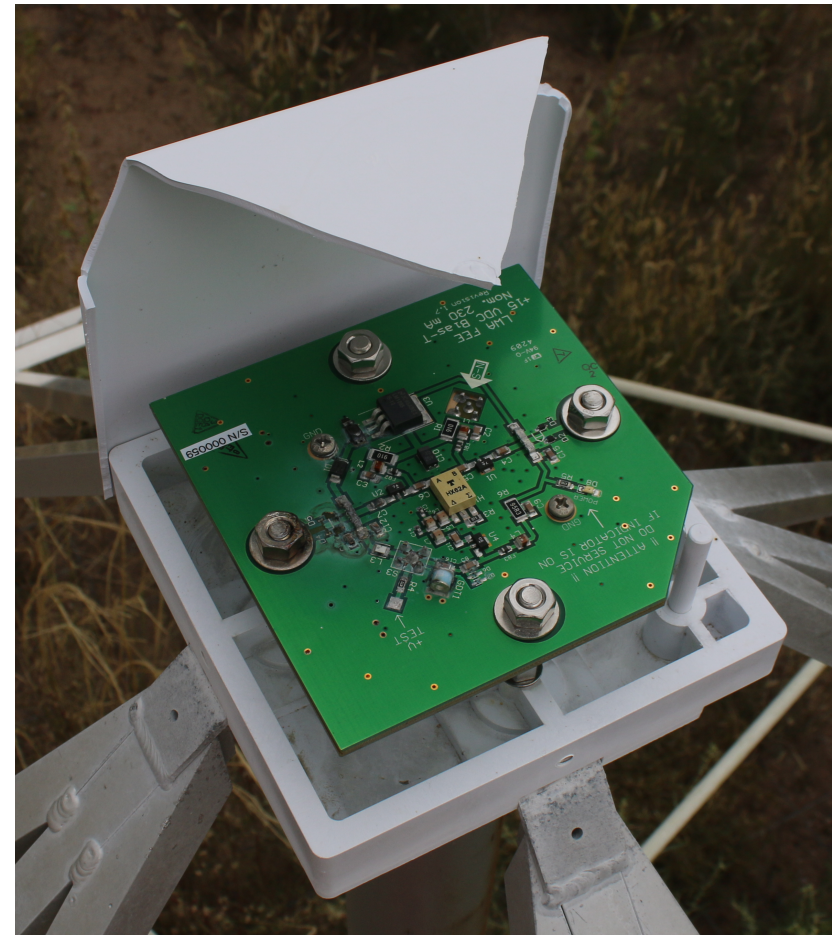
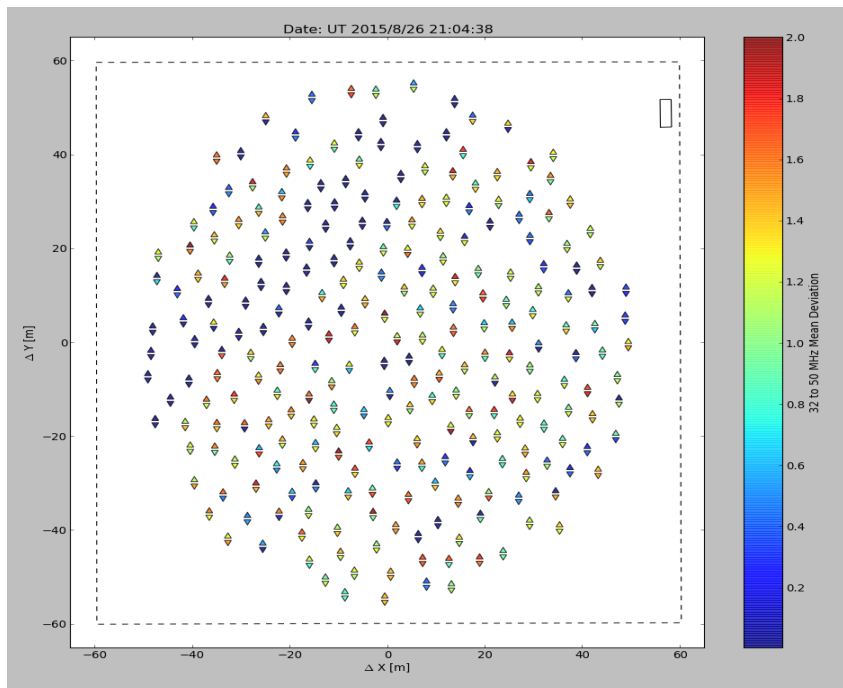
HAL is taking control of LWA1 in order to protect ASP and DP by shutting both systems down. Observations will be canceled until the lightning has cleared the area.

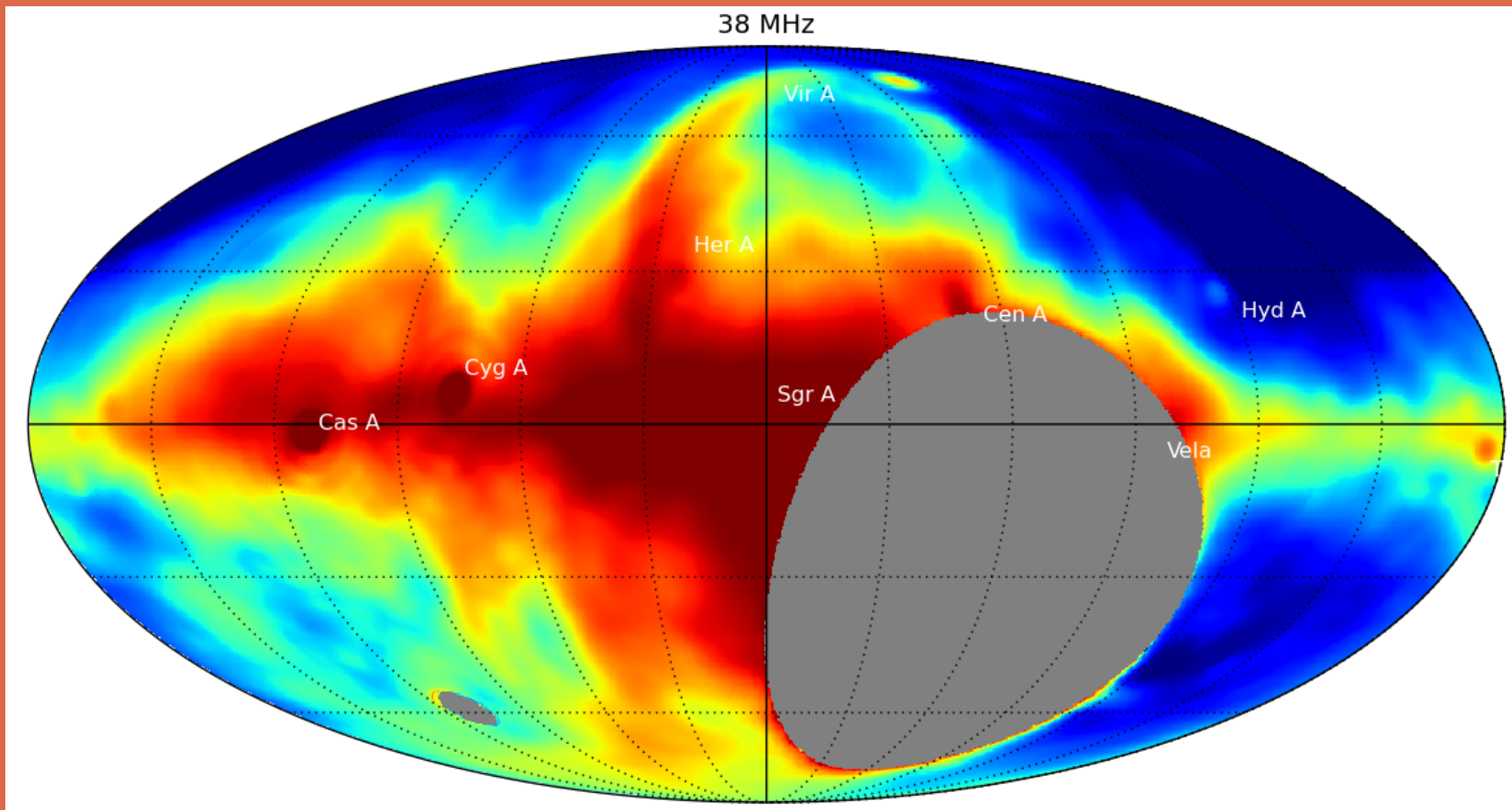
The following 'at' commands have been canceled:

```
* 5920
* 5921
```

LWA1 Lightning Damage

“Lucky” Stand #83 was struck at 4:59pm on August 21





SCIENCE HIGHLIGHTS

LWA1 & LWA-SV Science

Astrophysics

- **Cosmology**
Observing cosmic dawn through redshift 30 absorption of the 21 cm line. High redshift radio galaxies, containing the earliest black holes
- **Acceleration, Propagation & Turbulence in the ISM**
Origin, spectrum & distribution of Galactic cosmic rays, Supernova remnants & Galactic evolution, Pulsars and their environments
- **Solar Science & Space Weather**
Radio heliography of solar bursts & coronal mass ejections, Solar magnetic fields
- **Exploration of the Transient Universe**
New coherent sources, GRB prompt emission, poorly explored parameters space ...

Iono- & Atmospheric Physics

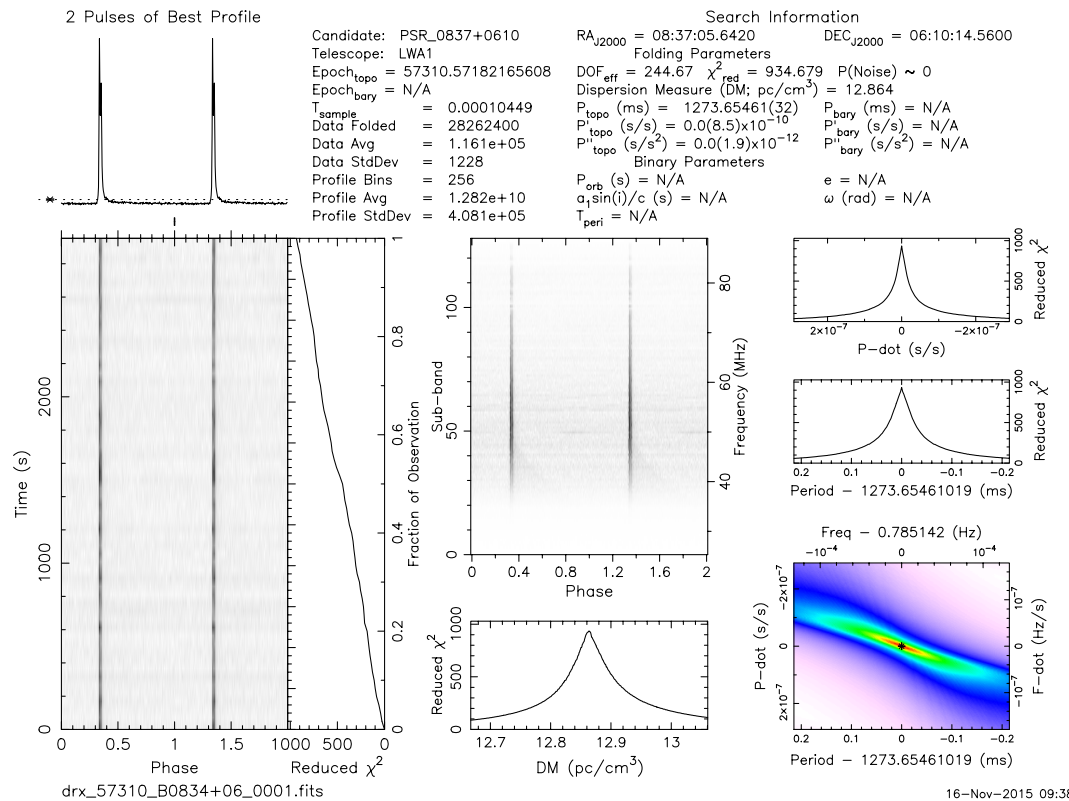
- Unprecedented continuous spatial & temporal imaging of the ionosphere
- Test and improve global ionospheric models
- High-time-resolution Imaging of Lightning

Your ideas?

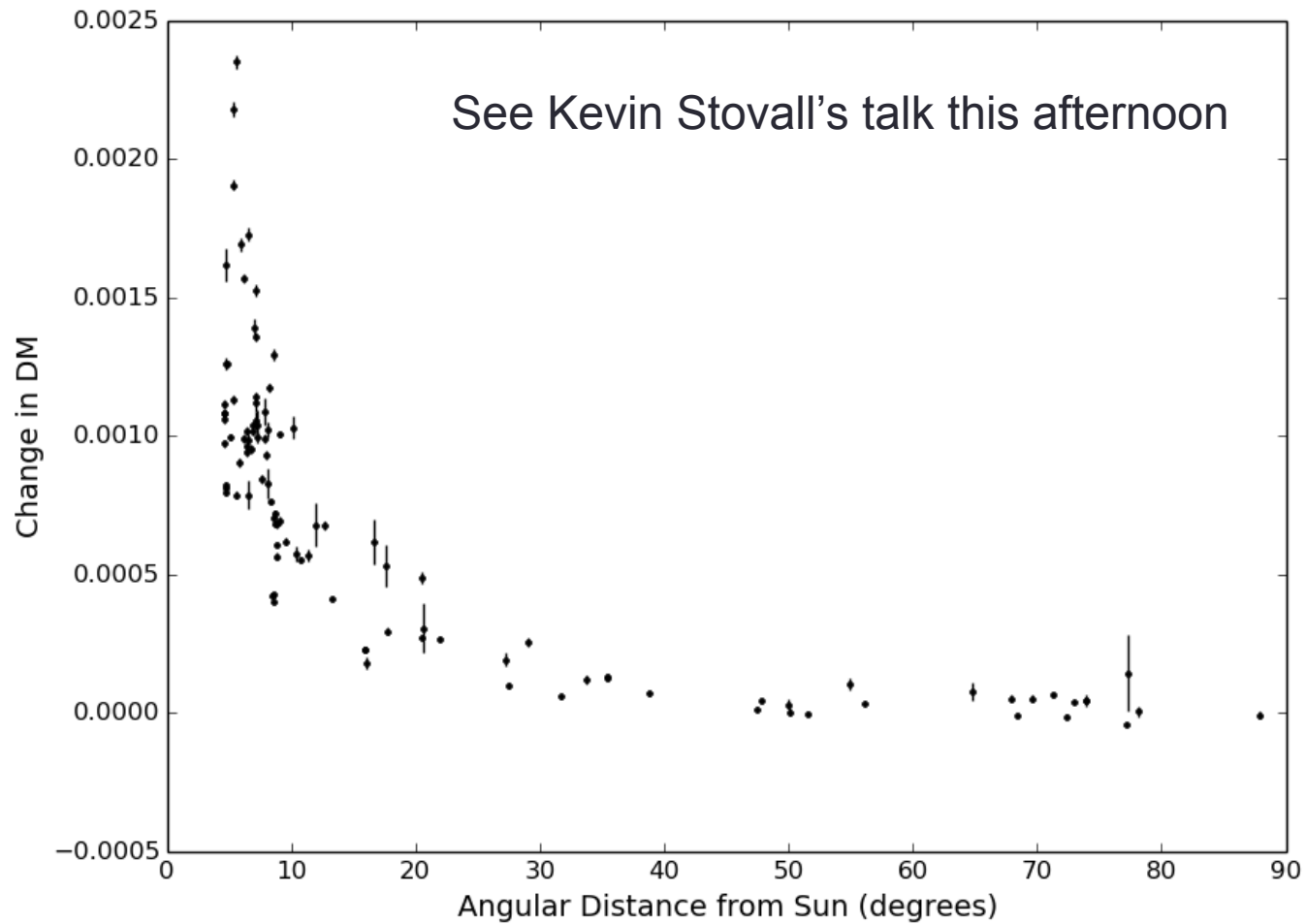
All of LWA1 time is open skies. Your observing proposals are welcome!

LWA1 Pulsar Detections

- 61 pulsars detected
 - 58 pulsations
 - 3 single pulses
 - Many in the LWA Pulsar Archive
- MSPs
 - J0030+0451
 - J0034-0534
 - J2145-0750
- GPs and RRATs
 - J0054+66
 - J0207+7005
 - B0531+21
 - J2325-0530
- Mode Switching
 - B0823+26
 - B0943+10



Solar DM Enhancement



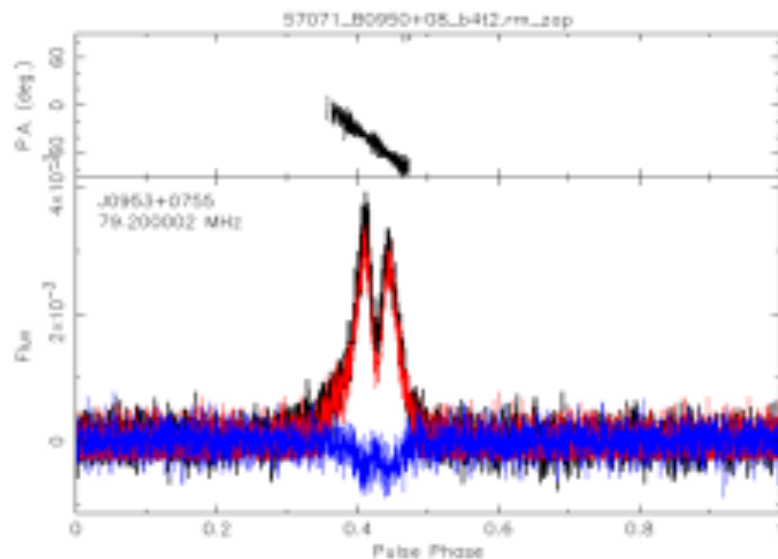
Preliminary Polarization Profiles (sight-lines far from the Sun)

PSR B0950+08

$$RM_{\text{Meas}} = 2.36(4)$$

$$RM_{\text{lon}} = 1.2(1)$$

$$Rm_{\text{cor}} = 1.2(2)$$

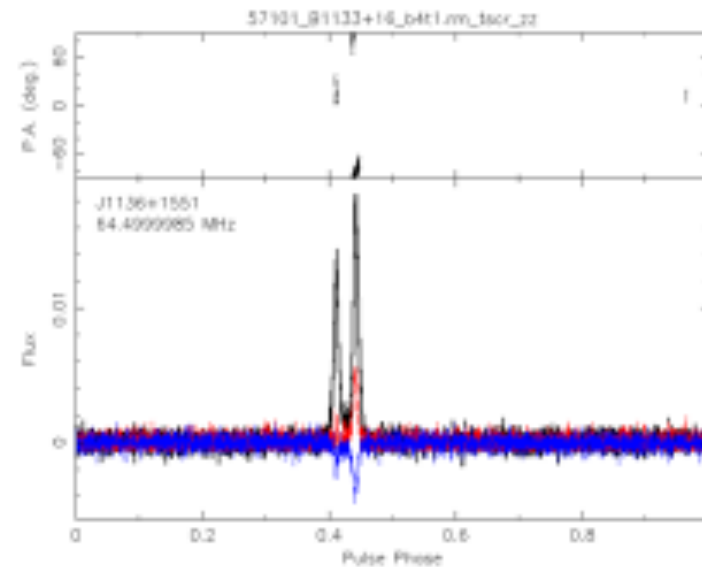


PSR B1133+16

$$RM_{\text{Meas}} = 4.61(1)$$

$$RM_{\text{lon}} = 0.84(4)$$

$$Rm_{\text{cor}} = 3.77(5)$$



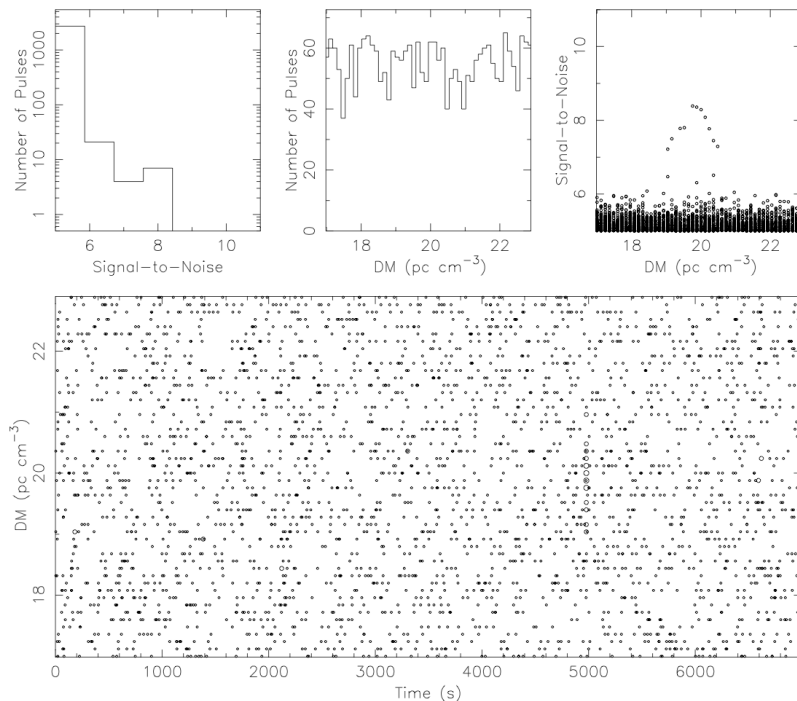
Rotating Radio Transients (RRATs)

J0203+7005

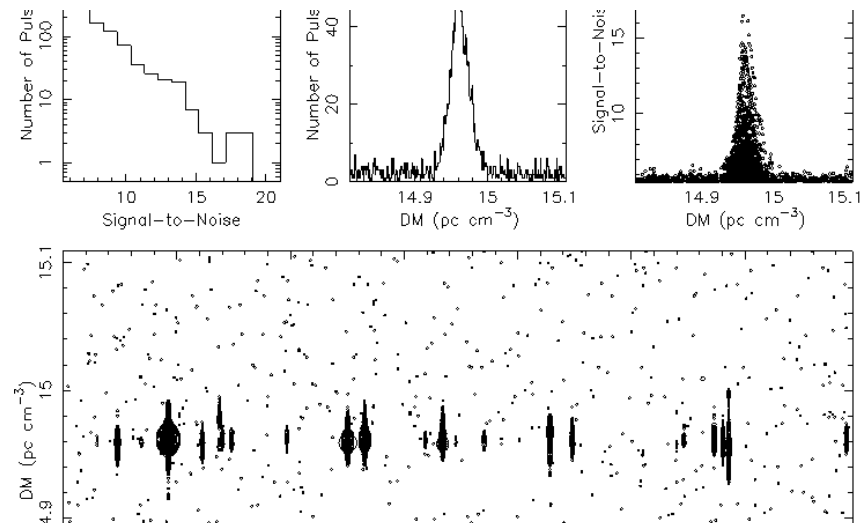
J2324-0530

Single pulse results for 'J0203+70'

Source: J0203+70 RA (J2000): 02:03:00.0000 N samples: 33473720
Telescope: LWA DEC (J2000): 70:23:00.0000 Sampling time: 208.98 μ s
Instrument: DRX MJD_{bary}: 56582.270608784478 Freq_{ctr}: 64.0 MHz



rmiller 13-Dec-2013 00:05

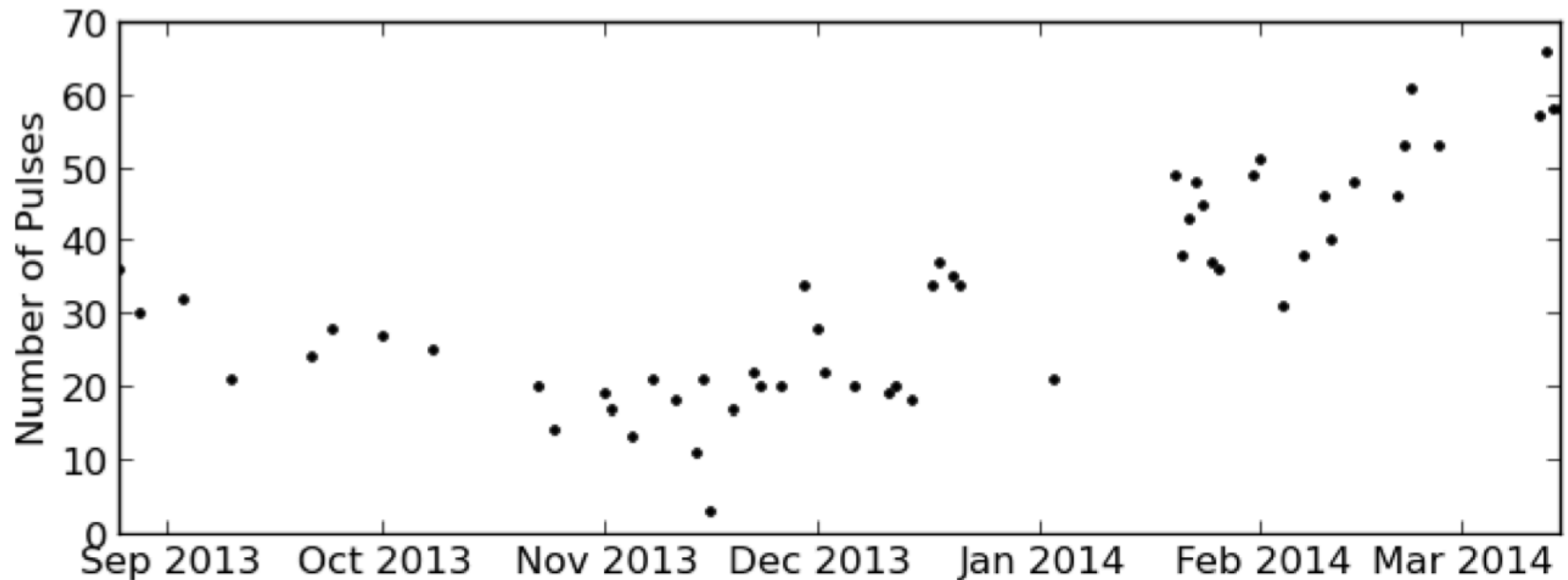


3 detected so far, search is ongoing
McCrackan et al., in prep

Crab Giant Pulses

Flux densities 10-120 kJy

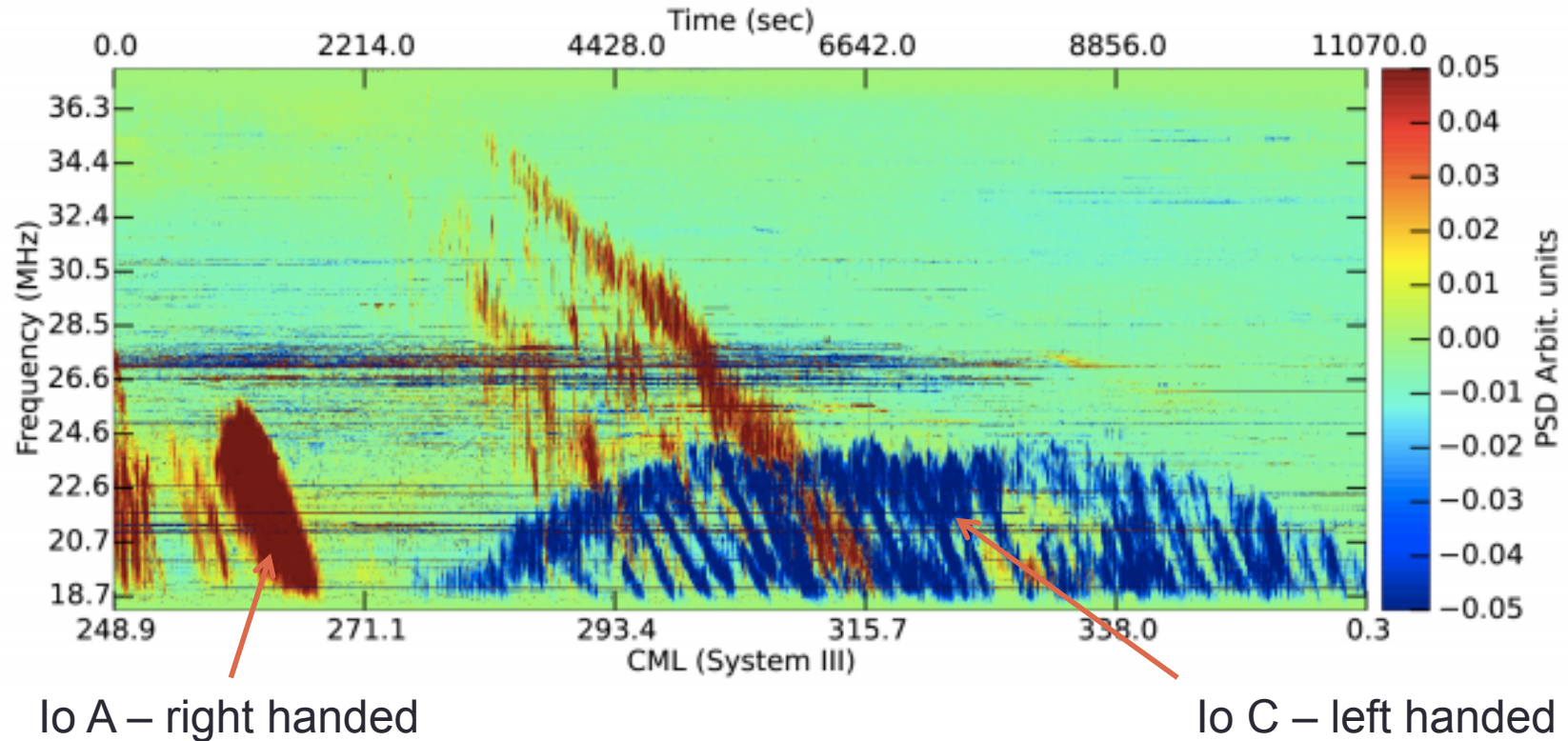
Number of pulses/hour went up by factor 3 over 6 month period



Eftekhari et al. in prep

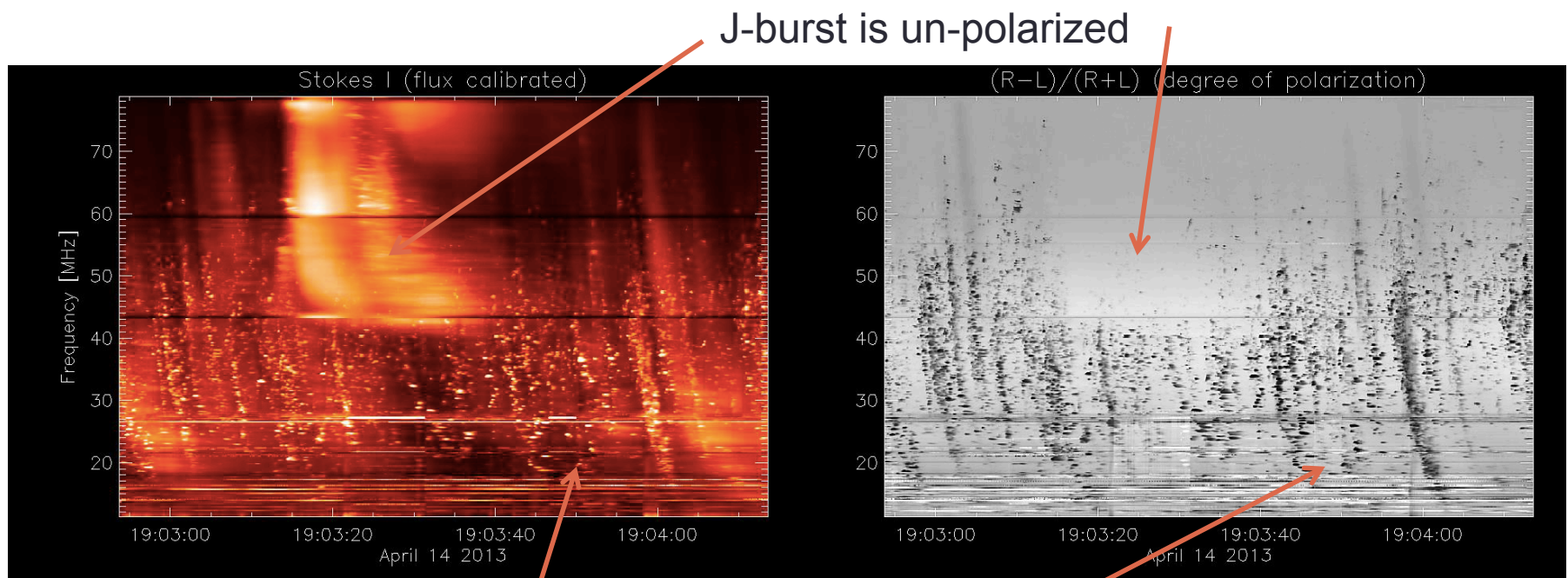
Jovian Decametric Emission

Io A/C bursts from 2013 March 06
Clarke et al. 2014



See Chuck Higgins' talk this afternoon

The Dynamic Sun



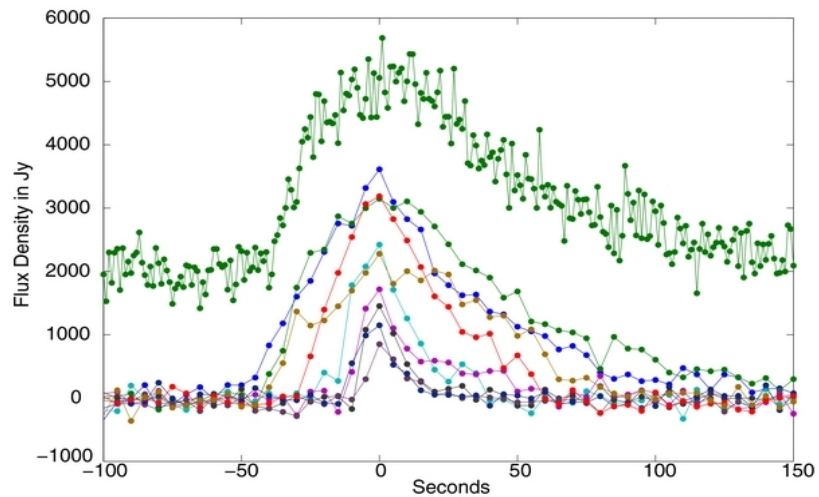
Type IIIb bursts are strongly left-hand polarized

Tun, Cutchin, & White, 2015, submitted

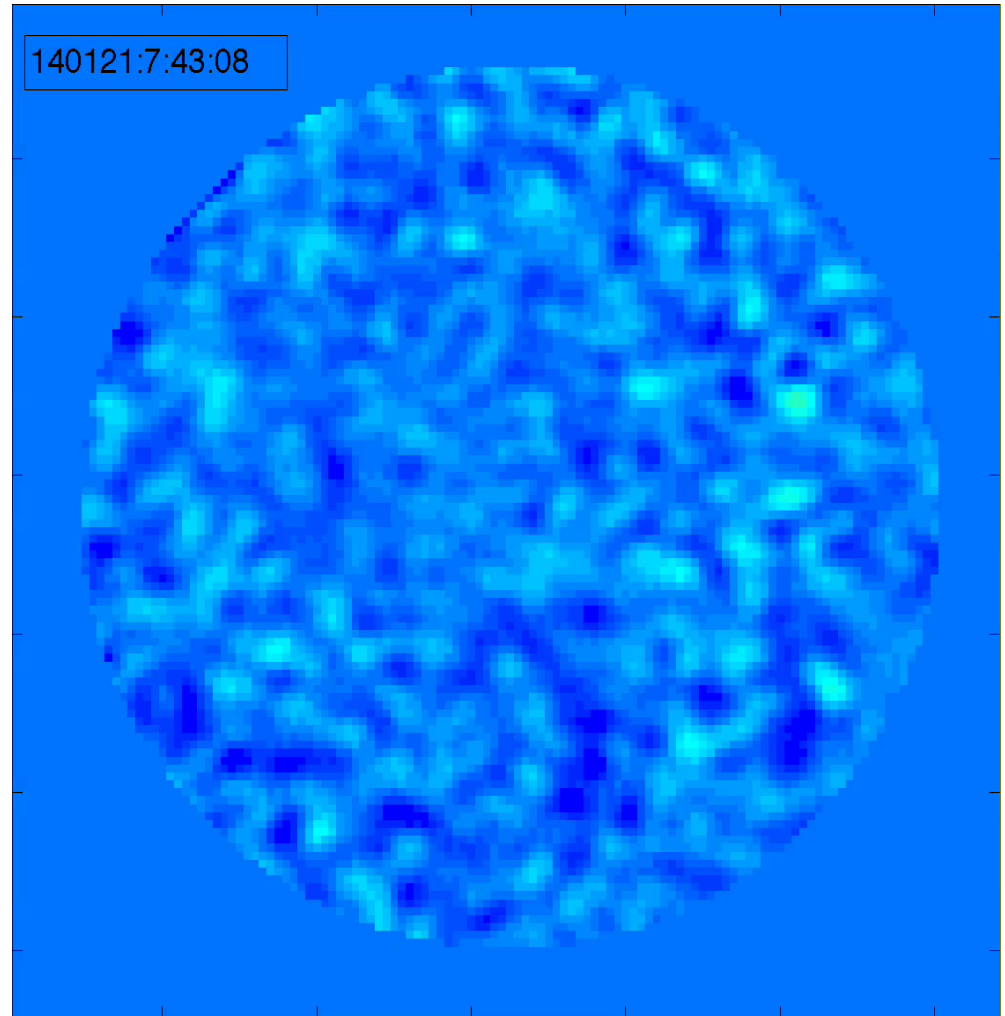
Fireball Emission

[http://
ida10g.alliance.unm.edu/
PASI_Fireballs/](http://ida10g.alliance.unm.edu/PASI_Fireballs/)

Ken Obenbergers's talk
Friday morning



Light curves of the brightest transients

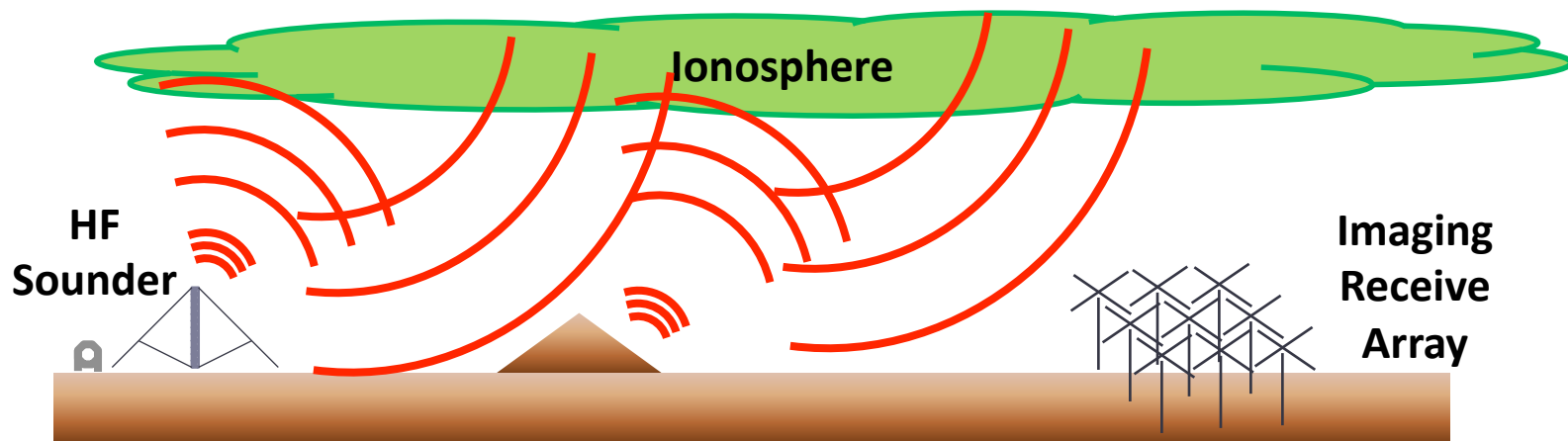




Ionospheric & Terrain Mapping



- In addition to transmitter direct echoes, terrain features illuminated by sounder also potential control points for ionospheric reconstruction
 - Appear as “2nd-hop” echoes
 - Echo strength depends on surface tilt and roughness
 - Readily computed from digital elevation models



Ionospheric Mapping

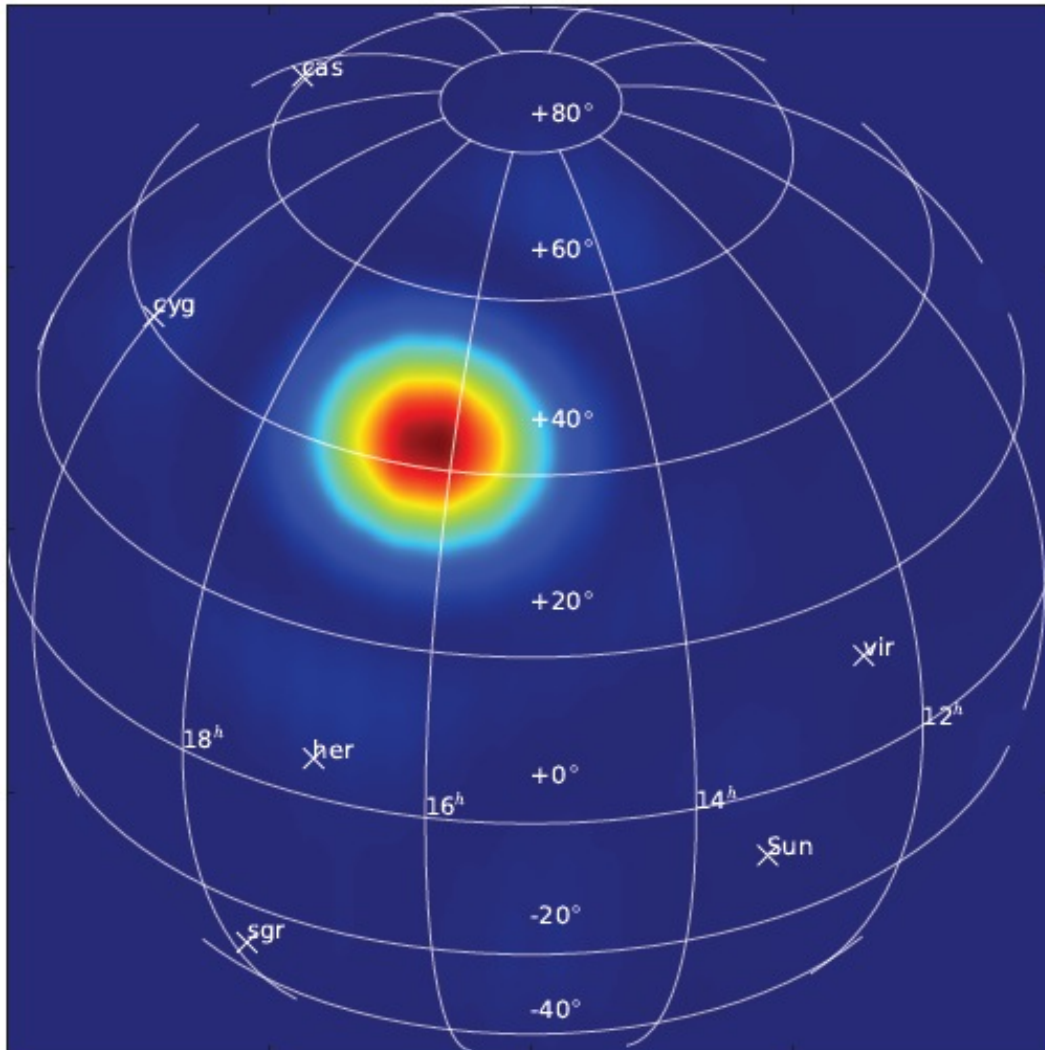
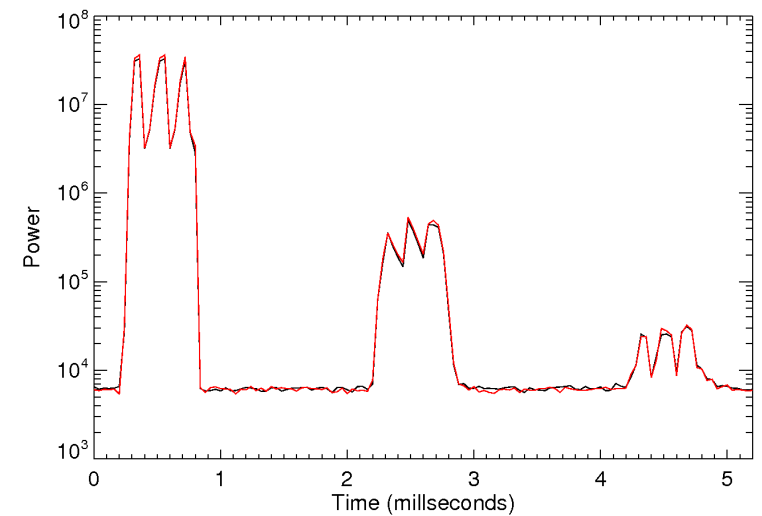


Image of the Kirtland digisonde reflecting off the ionosphere



THE FUTURE

The LWA

- Goal of 53 LWA stations
- Baselines up to 400 km for resolution 2” at 80 MHz with mJy sensitivity
- Cost is ~\$1M/station



The Expanded LWA



- Leverage the new 4-band system on the VLA to build a commensal system
- VLA 4-band + 2 LWA stations doubles the resolution and triples the sensitivity of the VLA

The Expanded LWA

- Early tests show that this is feasible
 - 3C196 with LWA1 and six VLA antennas
- See Frank Schinzel's talk Friday afternoon

